

# Double Chooz

*(first multi-detector appetiser)*

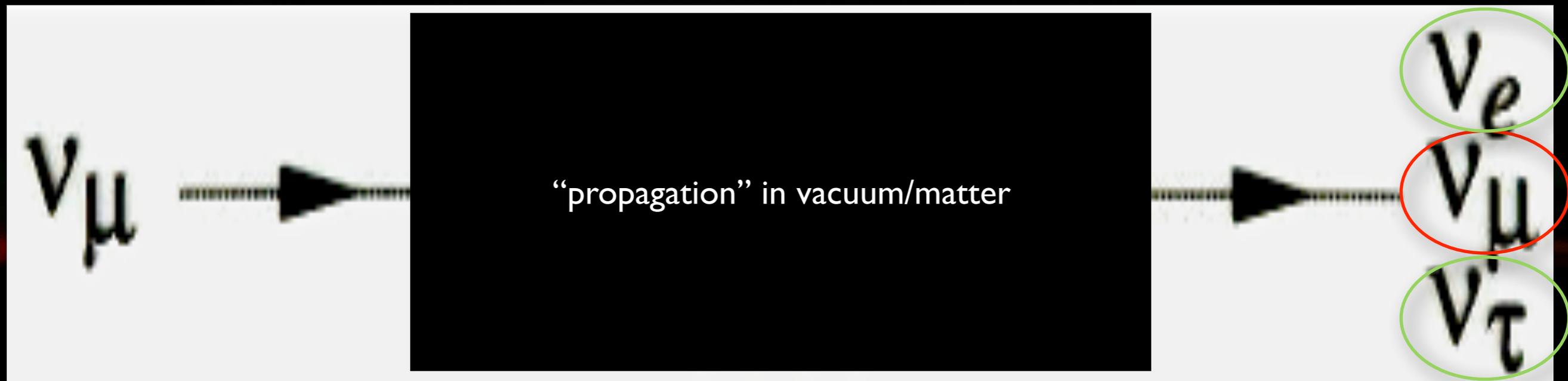
*FNAL laboratory (Chicago, USA)  
March 2016*

**Anatael Cabrera**

CNRS / IN2P3 @ APC (Paris)

Let's take  $\nu_\mu$  (a popular example) to start with...

*disappearance*  
*appearance*



observations: **disappearance** ( $\rightarrow$  Nobel Prize 2015) & **appearance** ( $>2010$ )

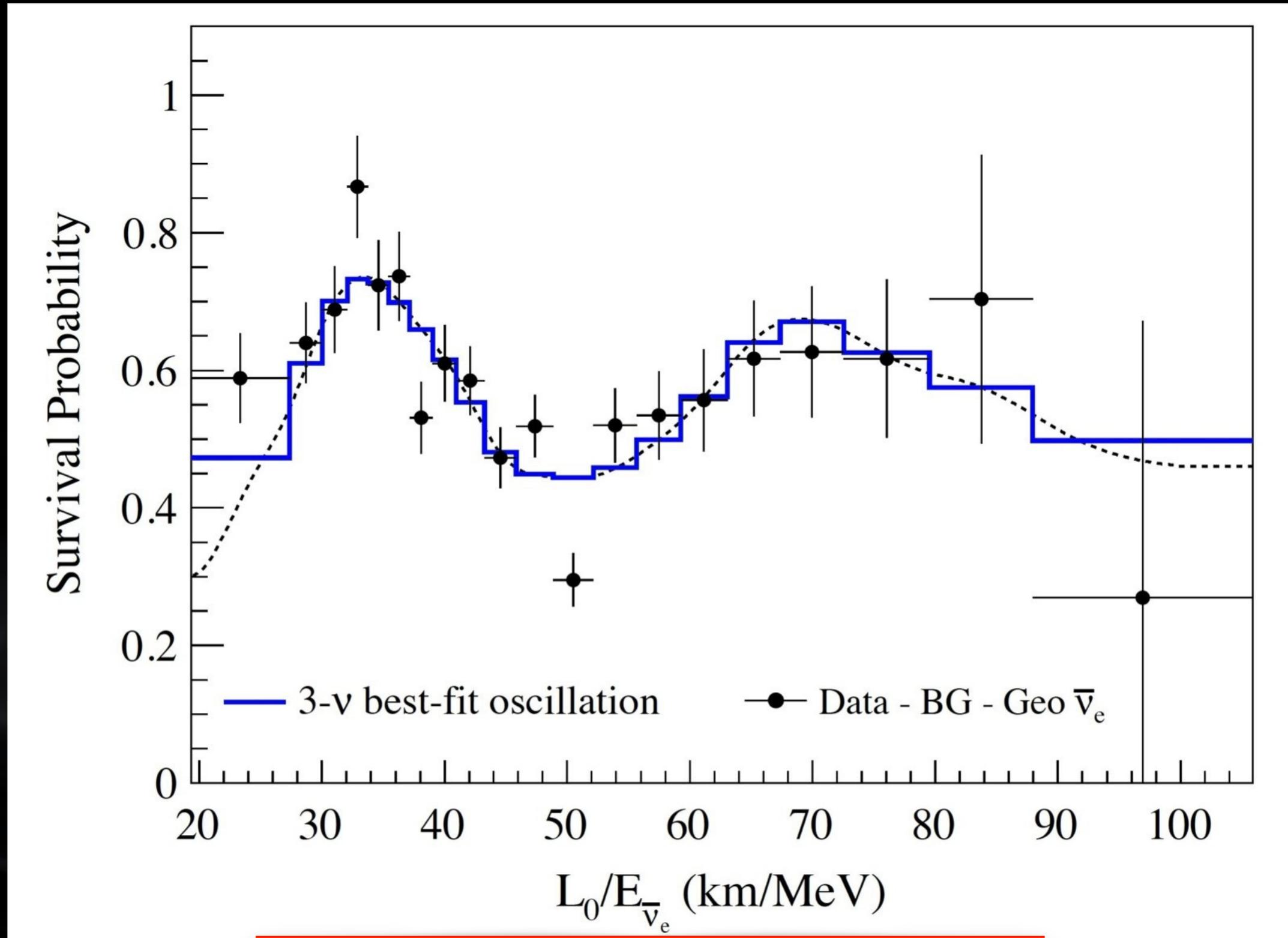
**all observations (many!) follow well one model: 3 $\nu$  oscillation**

# “mixing”: a common phenomenon...



# the latest KamLAND's $P(\nu_e \rightarrow \nu_e)$ ...

the most beautiful **E/L disappearance** so far... (to me)



arguably a Noble prize worth E/L...

“atmospheric”  $\Rightarrow \theta_{23} \sim 45^\circ$

$\theta_{13}$  & “dirac”  $\delta_{CP}$

“solar”  $\Rightarrow \theta_{12} \sim 33^\circ$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{matrix} \text{sub-leading} \\ \leftarrow \end{matrix} \begin{pmatrix} c_{13} & 0 & e^{-i\delta} s_{13} \\ 0 & 1 & 0 \\ -e^{i\delta} s_{13} & 0 & c_{13} \end{pmatrix} \begin{matrix} \text{sub-leading} \\ \leftarrow \end{matrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$\Delta m_{31}^2$                        $\Delta m_{31}^2$                        $\Delta m_{21}^2$

atmos+LBL(dis)                      Chooz+LBL(app)                      solar+KamLAND

$P(\nu_\mu \rightarrow \nu_\mu)$                        $P(\nu_e \rightarrow \nu_e)$  &  $P(\nu_\mu \rightarrow \nu_e)$                        $P(\nu_e \rightarrow \nu_x)$

knowledge on  
 $\theta_{13}$  &  $\delta_{CP}$   
[**this talk**]

$(\nu_e, \nu_\mu, \nu_\tau)^T = U(\nu_1, \nu_2, \nu_3)^T$ , where  $U^{PMNS}$  looks like

$$\begin{pmatrix} \blacksquare & \blacksquare & \blacksquare \\ \blacksquare & \blacksquare & \blacksquare \\ \blacksquare & \blacksquare & \blacksquare \end{pmatrix}$$

$U^{PMNS}$

$\theta_{13}$  drives this!!!

**note:**  $\theta_{13}$  dominant term @ position “1-3” of matrix (not mixing of  $m_1$  and  $m_3$  states)

$\sin^2(\theta_{13})$ 

At present time direct measurements of  $\sin^2(\theta_{13})$  are derived from the reactor  $\bar{\nu}_e$  disappearance at distances corresponding to the  $\Delta m_{32}^2$  value, i.e.  $L \sim 1\text{km}$ . Alternatively, limits can also be obtained from the analysis of the solar neutrino data and accelerator-based  $\nu_\mu \rightarrow \nu_e$  experiments.

VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.19 ± 0.12 OUR AVERAGE</b>				
2.15 ± 0.13		<sup>1</sup> AN	15 DAYA	DayaBay, Ling Ao/Ao II reactors
2.3 <sup>+</sup> 0.9 - 0.8		<sup>2</sup> ABE	14H DCHZ	Chooz reactors
2.12 ± 0.47		<sup>3</sup> AN	14B DAYA	DayaBay, Ling Ao/Ao II reactors
2.5 ± 0.9 ± 0.9		<sup>4</sup> ABE	13C DCHZ	Chooz reactors
2.9 ± 0.3 ± 0.5		<sup>5</sup> AHN	12 RENO	Yonggwang reactors

• • • We do not use the following data for averages, fits, limits, etc. • • •

Citation: K.A. Olive *et al.* (Particle Data Group), Chin. Phys. C, **38**, 090001 (2014) and 2015 update

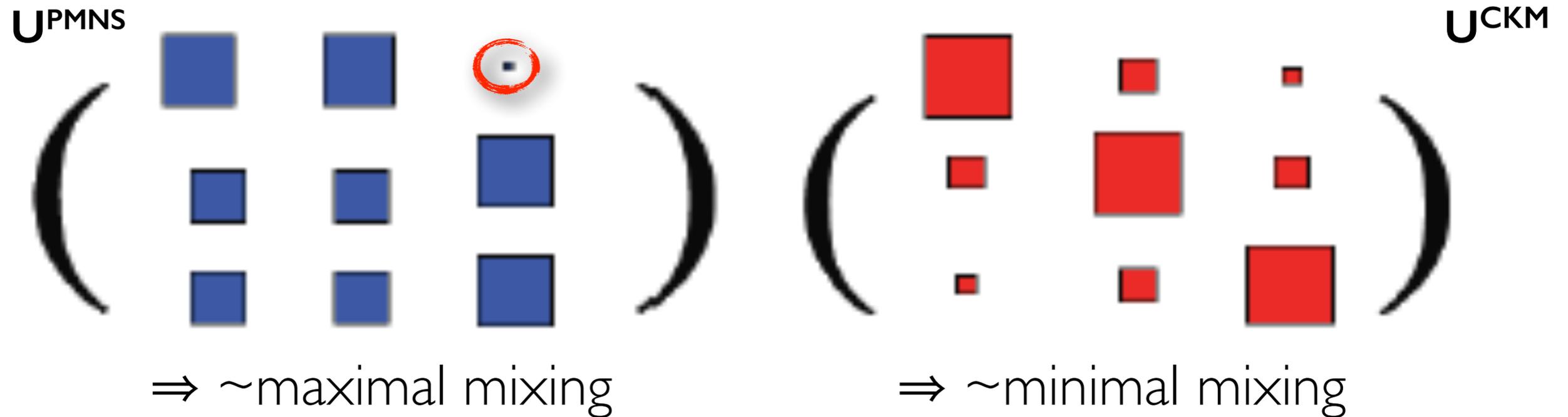
(the boring but still real argument)

**as all parameters in SM are to be measured and plug-in (i.e. no prediction from SM)**

(but models beyond-SM might have predictions our SM values)

**our knowledge on  $\theta_{13}$  dominated by DYB**

also linked to bigger questions...



why are CKM and PMNS so?

if 3 families  $\Rightarrow$  **unitary** (really?)

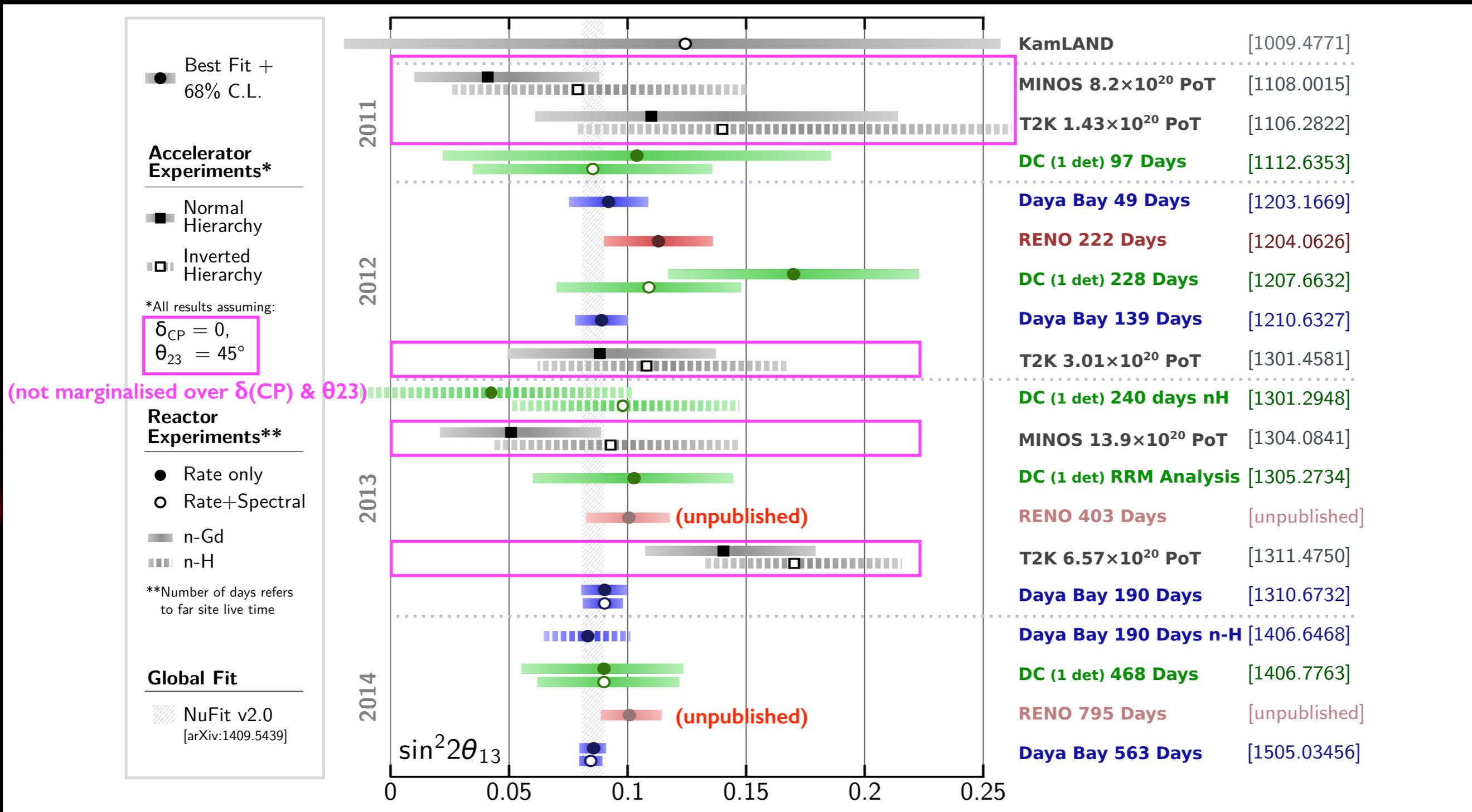
CKM unitarity being tested to high precision  
PMNS very early stages

if not unitary  $\Rightarrow$  indirect evidence of  $>3$  families?

the world's  $\theta_{13}$  knowledge...



# $\theta_{13}$ -reactor measurements (summer 2015)...



unsurpassable reactor precision → “reactor- $\theta_{13}$ ” for several decades to go!  
 (measurement by T2K+MINOS+NOvA, KamLAND⊕Solar, etc)

## Comparing T2K results with reactors

T2K  $\sin^2 2\theta_{13}$  result computed assuming  $\sin^2 \theta_{23} = 0.5$ ,  $\delta_{CP} = 0$ , and normal hierarchy (top), and inverted hierarchy (bottom)

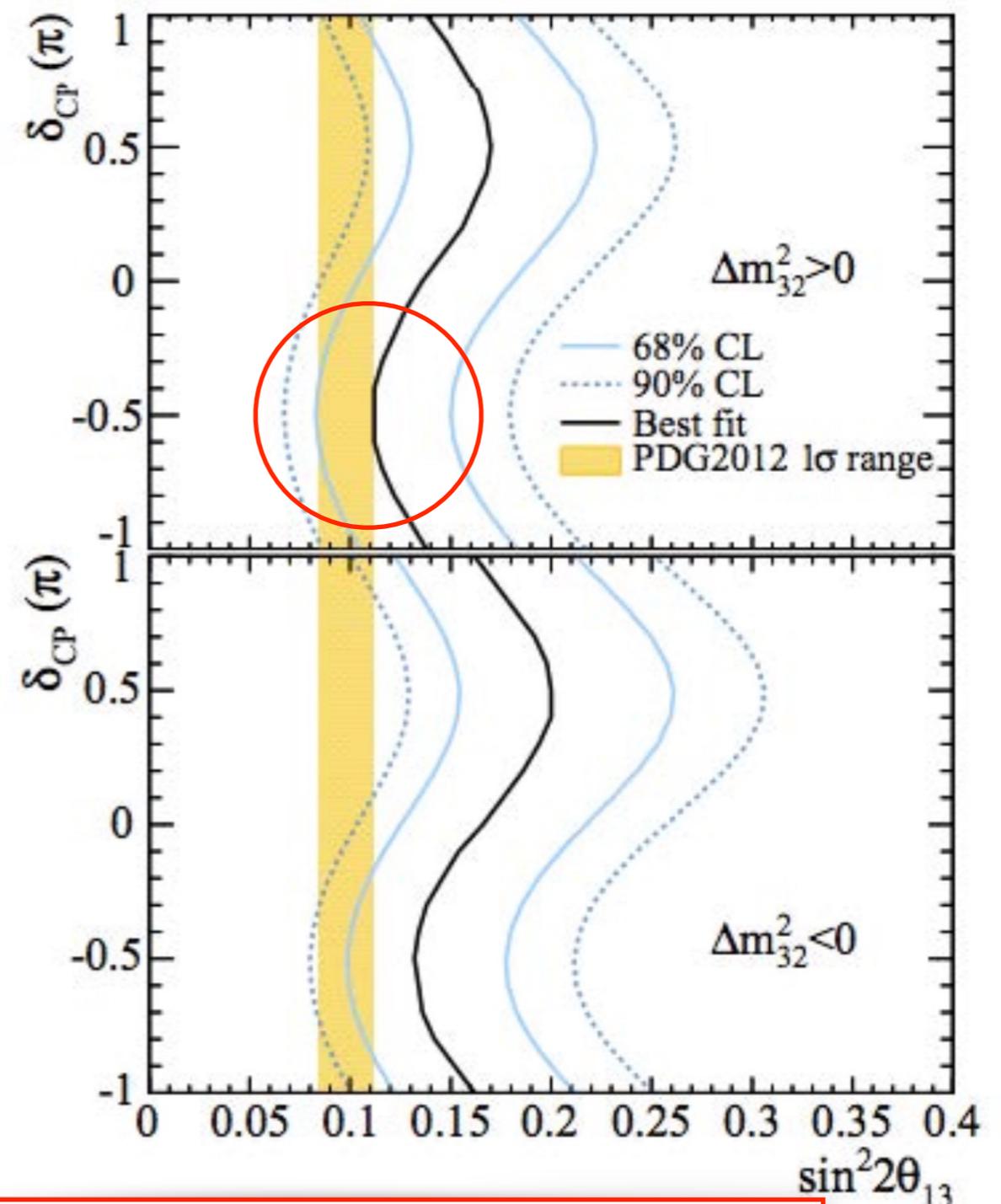
Consistent at 90% CL ( $1.6\sigma$ )

...but excess by T2K nudges all remaining unknowns in direction to increase rates

- normal hierarchy

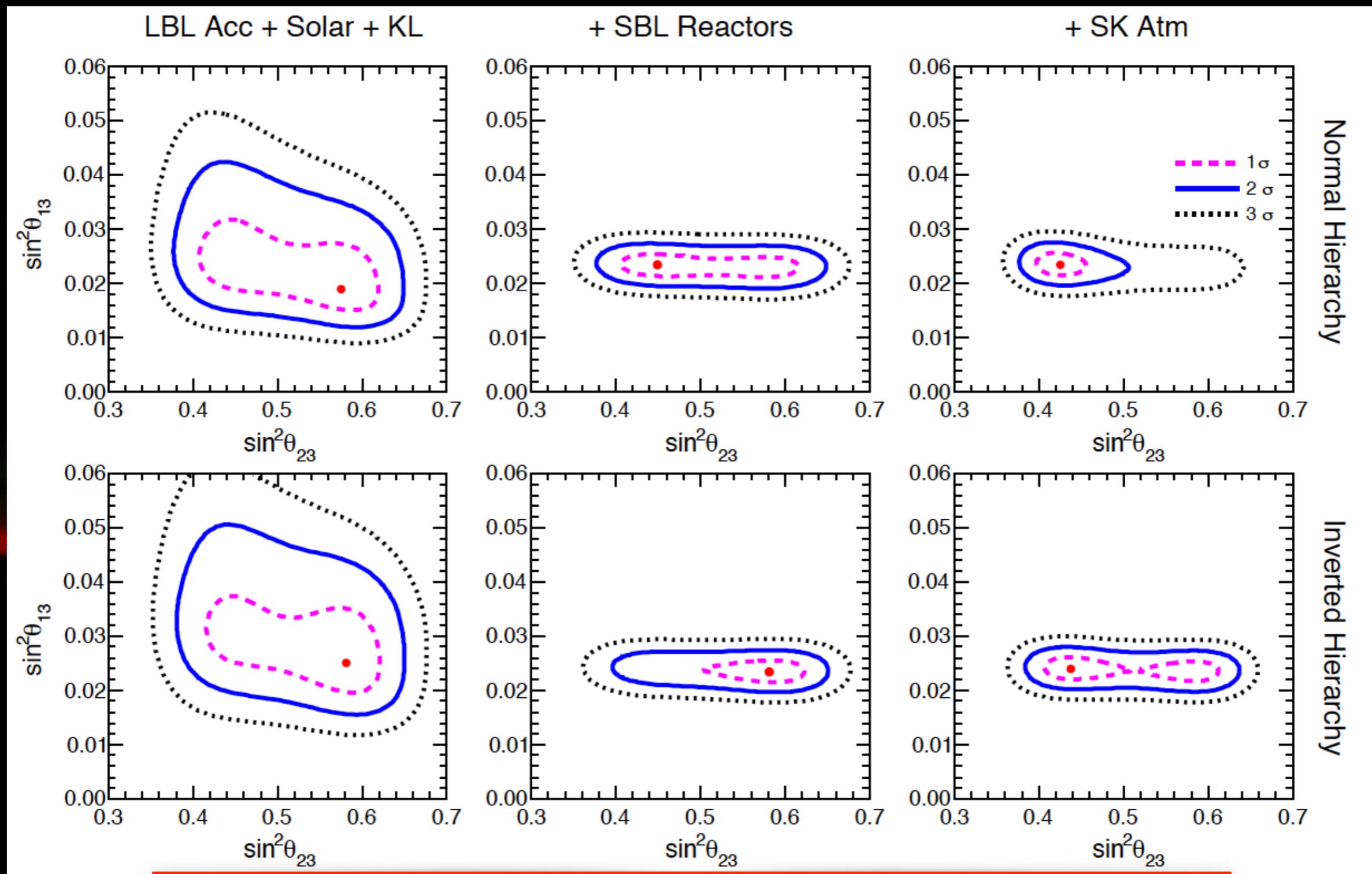
-  $\theta_{23} > 45^\circ$

-  $\delta_{CP} = -\pi/2$  (aka  $3\pi/2$ )



interplay between reactors & beams  $\rightarrow$  critical phenomenology

# || (Lisi et al Jan. 2014) impact to the $\theta_{23}$ octant...



reactors (again): may help to resolve  $\theta_{23}$ -octant ambiguity?

reactor data (only  $\theta_{13}$ ) orthogonality aids beam data  $\Rightarrow$  **a deeper sight together**

# Double Chooz collaboration



## Brazil

CBPF  
UNICAMP  
UFABC



## France

APC  
CEA/DSM/IRFU:  
SPP  
SPhN  
SEDI  
SIS  
SENAC  
CNRS/IN2P3:  
Subatech  
IPHC



## Germany

EKU Tübingen  
MPIK Heidelberg  
RWTH Aachen  
TU München  
U. Hamburg



## Japan

Tohoku U.  
Tokyo Inst. Tech.  
Tokyo Metro. U.  
Niigata U.  
Kobe U.  
Tohoku Gakuin U.  
Hiroshima Inst.  
Tech.



## Russia

INR RAS  
IPC RAS  
RRC Kurchatov



## Spain

CIEMAT-  
Madrid



## USA

U. Alabama  
ANL  
U. Chicago  
Columbia U.  
UCDavis  
Drexel U.  
IIT  
KSU  
LLNL  
MIT  
U. Notre Dame  
U. Tennessee

Spokesperson:  
H. de Kerret (IN2P3)

Project Manager:  
Ch. Veyssière (CEA-Saclay)

Web Site:  
[www.doublechooz.org/](http://www.doublechooz.org/)



# DC @ LNCA laboratory...



## Chooz Reactors

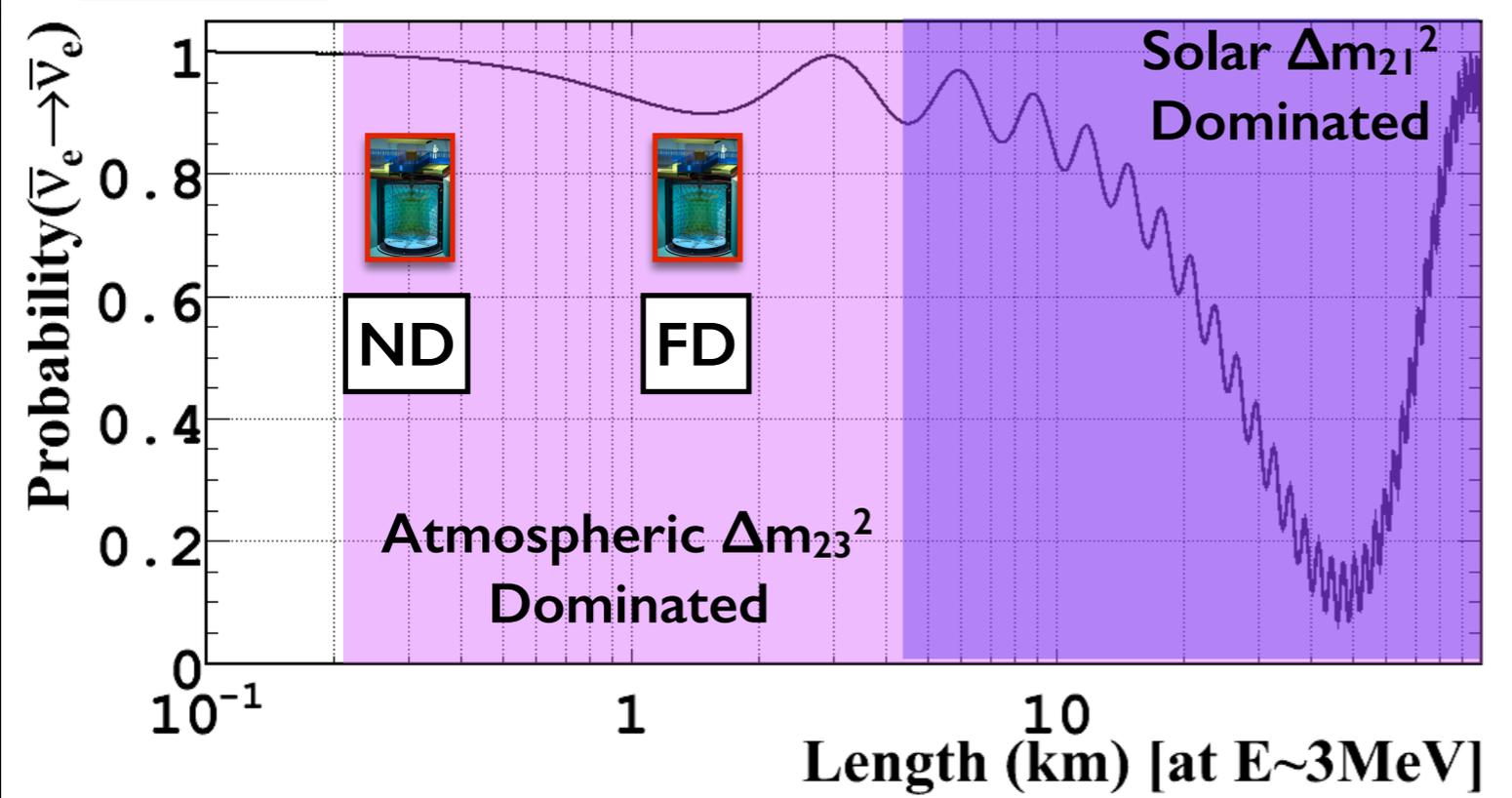
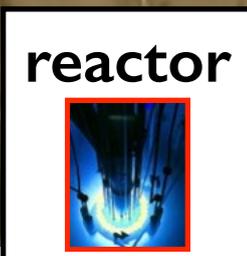
Power: 8.5GWth

$$\Rightarrow \sim 10^{21} \text{ v/s}$$

(N4s: very powerful)

## Near

$\langle L \rangle \approx 410\text{m}$   
~250 IBD/day  
~120 mwe  
Target: 8.2t  
Dec.2014



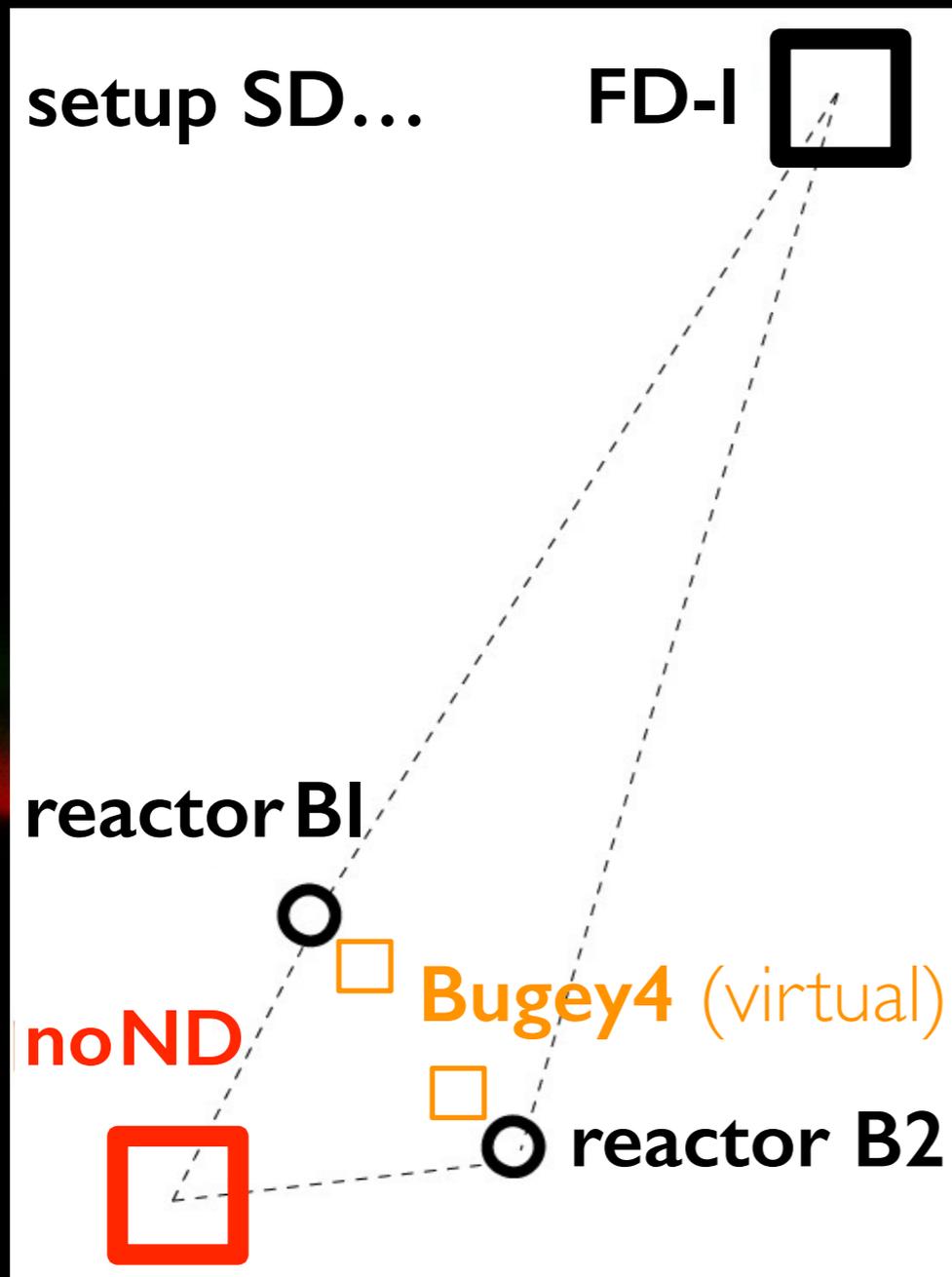
## Far

$\langle L \rangle \approx 1050\text{m}$   
~40 IBD/day  
~300 mwe  
Target: 8.2t  
April 2011

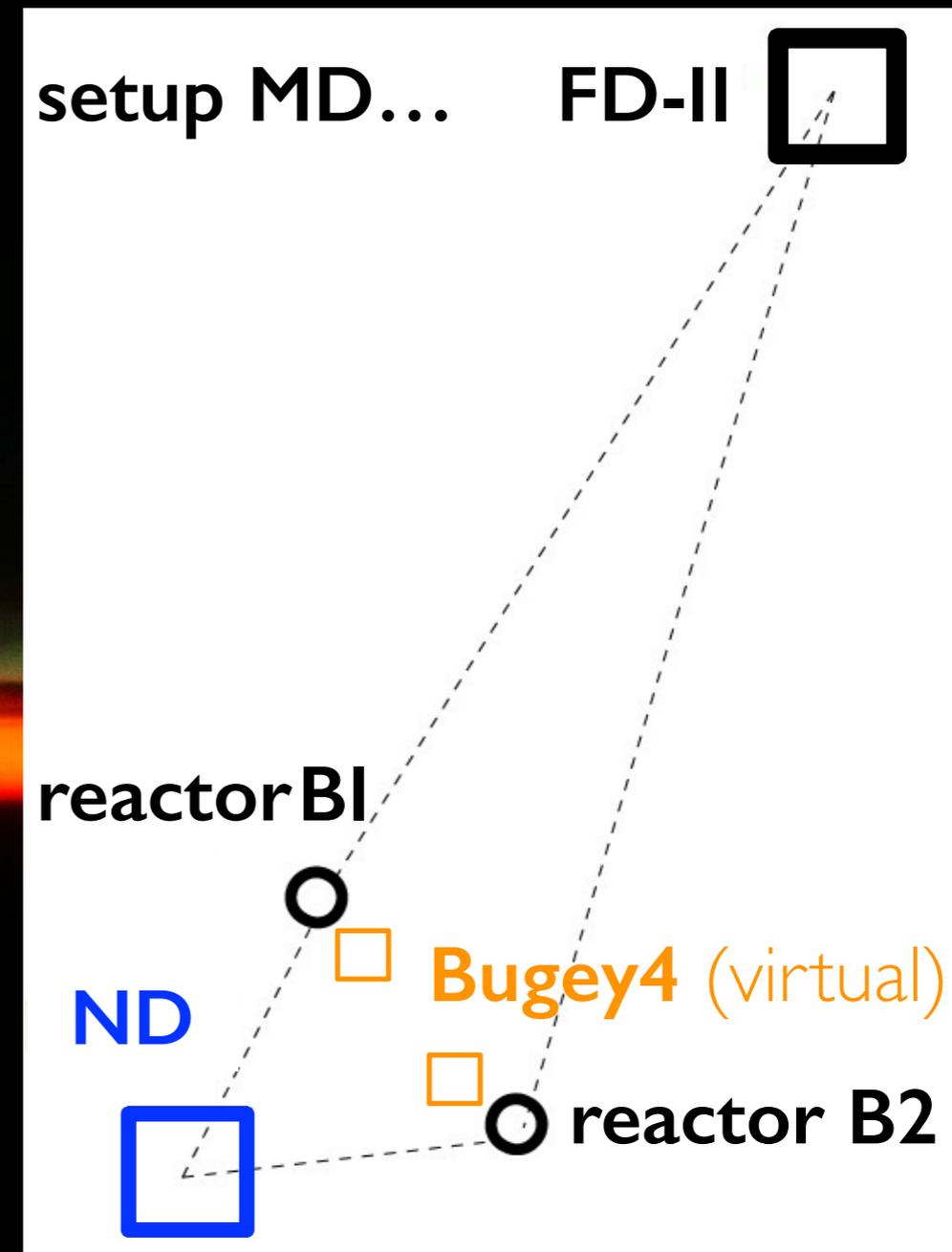


## our experiment @ DC-IV: 2 setup...

(exposure used: ~48months)



(exposure used: ~9months @ Moriond)



- IBD's @ FD →  $\theta_{13}$  info [no ND monitor but B4]
  - Li+He BG + Reactor-OFF →  $\delta(\text{BG})$  systematics
  - cross-check on BG, detection, etc (common FD)
- @Gd-III:  $1\sigma[\sin^2(2\theta_{13})] \approx 0.03$**

- IBD's @ FD →  $\theta_{13}$  info [ND monitor]
- Li+He BG 2 detectors →  $\delta(\text{BG})$  systematics
- lowest  $\delta(\text{flux})$  using ND monitor
- cross-check on BG, detection, etc (FD vs ND)

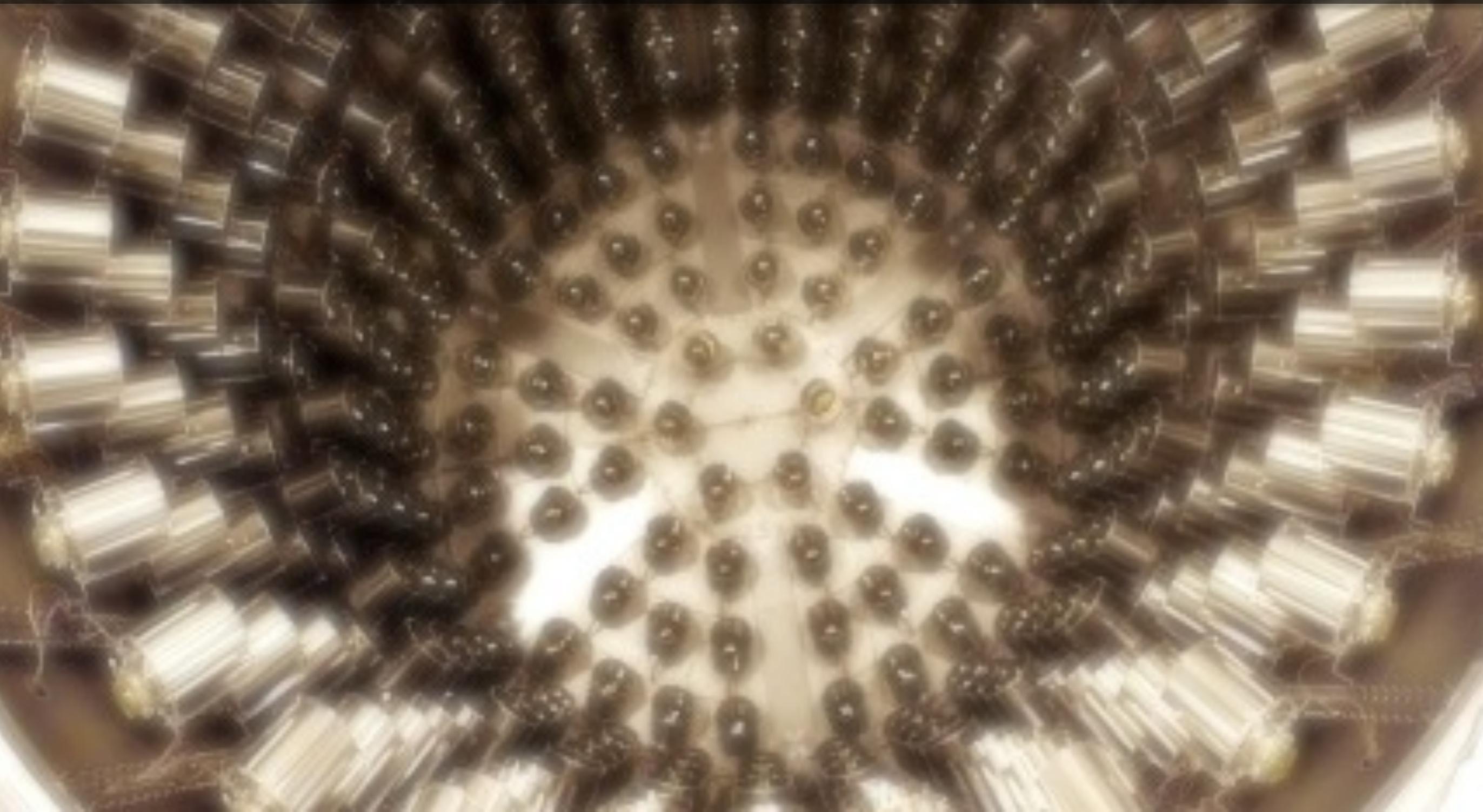
systematics	single detector (SD) (%)	multi-detector (MD) (%)
$\delta(\text{detection})$	$\sim 2.0$ (no fiducial volume)	$\sim 0.2$ ( <b>identical detectors</b> )
$\delta(\text{flux})$	$\sim 3.0$ (prediction) [ $\sim 1.7$ via <b>Bugey4</b> ]	$\leq 0.5$ ( <b>ND reactor monitor</b> )
$\delta(\text{background})$	$\leq 1.0$ (radio-purity+overburden)	$\leq 1.0$ ( <b>no suppression</b> )

— systematics uncertainties  $\sim 1\%$  each —

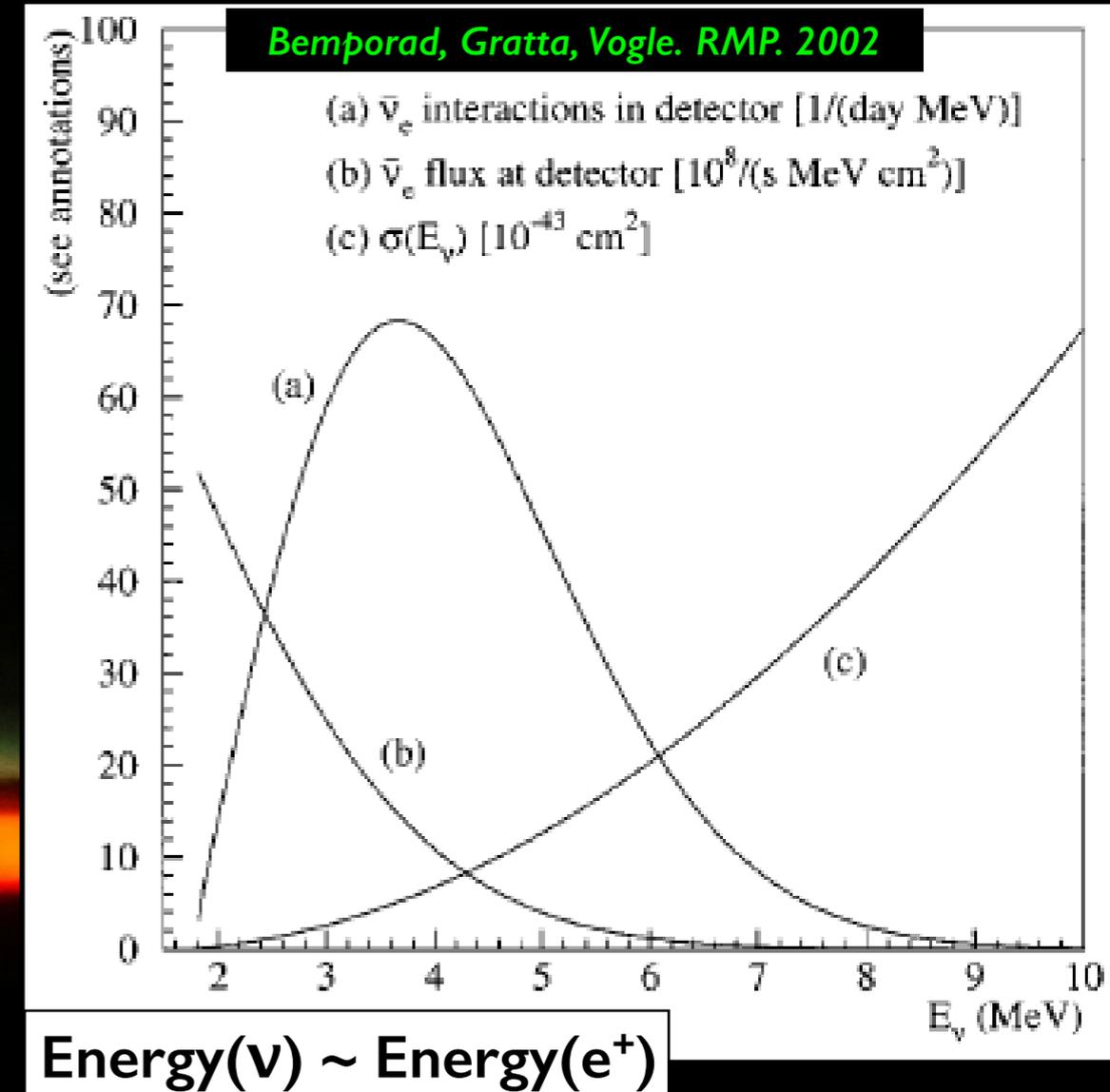
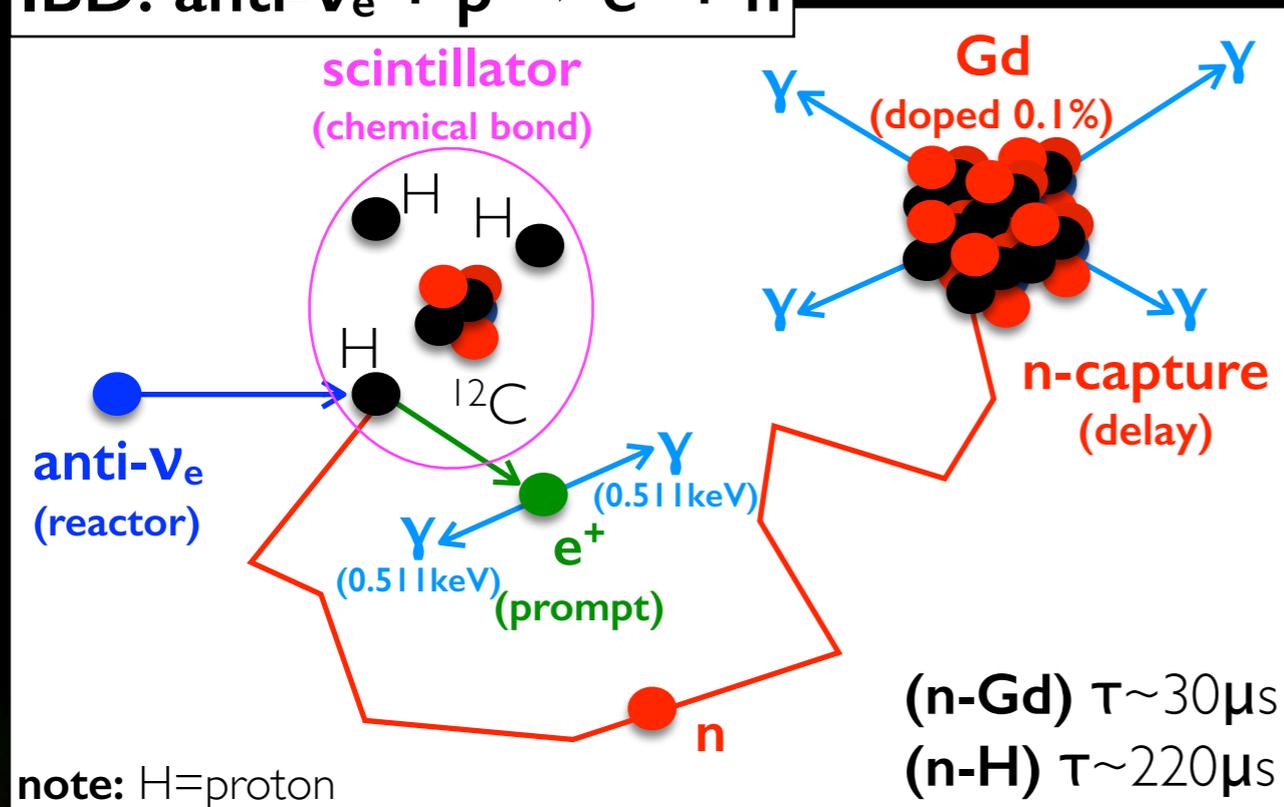
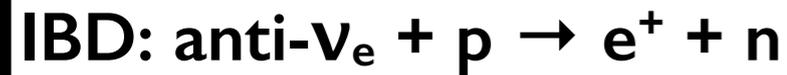
only one experiment providing  
 $\theta_{13}$  with a few ‰ errors?

(remember: LEP, etc)

our detectors...



# inverse- $\beta$ decay (IBD) interaction...



- high & well known  $\sigma^{\text{IBD}}$  [ $\tau_{\text{neutron}} = (881.5 \pm 1.5)\text{s}$ ]
- IBD manifests via **trigger-coincidence**
  - 1st trigger  $\rightarrow e^+$  (prompt) [ionisation  $\oplus$  annihilation]
  - 2nd trigger  $\rightarrow n$ -Gd capture (delay @  $\sim 8\text{MeV}$ )
- Energy( $\nu$ )  $\sim$  Energy( $e^+$ ) + 0.8MeV
- major rejection of radioactivity background...
  - time/space coincidence
  - delay @ 8MeV (radioactivity dominates  $\leq 3\text{MeV}$ )

## why IBD $\oplus$ Gd?

- small & shallow (high S/BG)
- no need for ultra-purity

$\Rightarrow$  **inexpensive % precision!!**

a generic  $\theta_{13}$ -LAND...

Outer  $\mu$ -Veto (OV)

Plastic-Scintillator: strips ( $\rightarrow$ tracking)

$\nu$ -Target (NT)

Liquid-Scintillator + Gd (0.1%)

$\gamma$ -Catcher (GC)

Liquid-Scintillator

Light Buffer

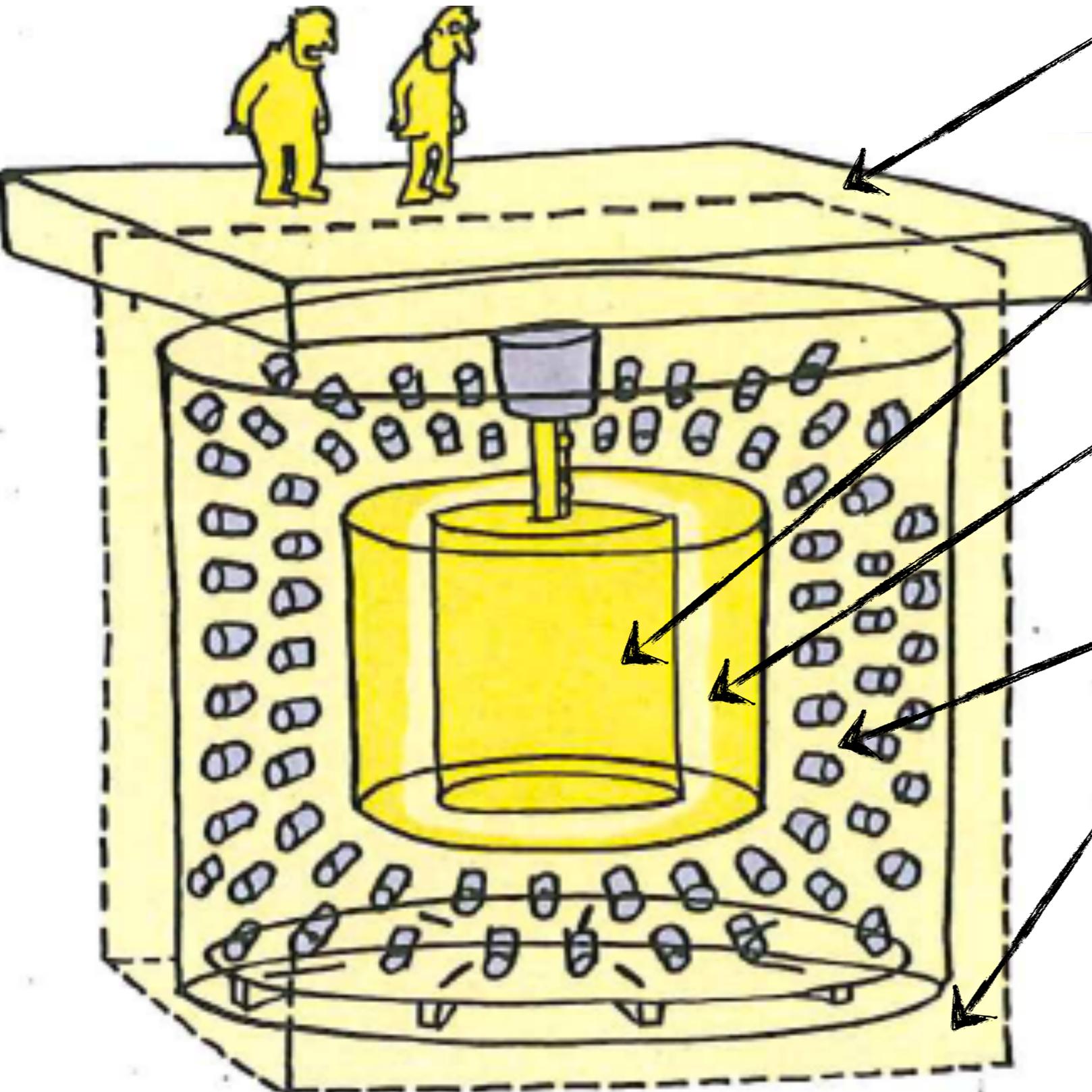
Oil (negligible scintillation)

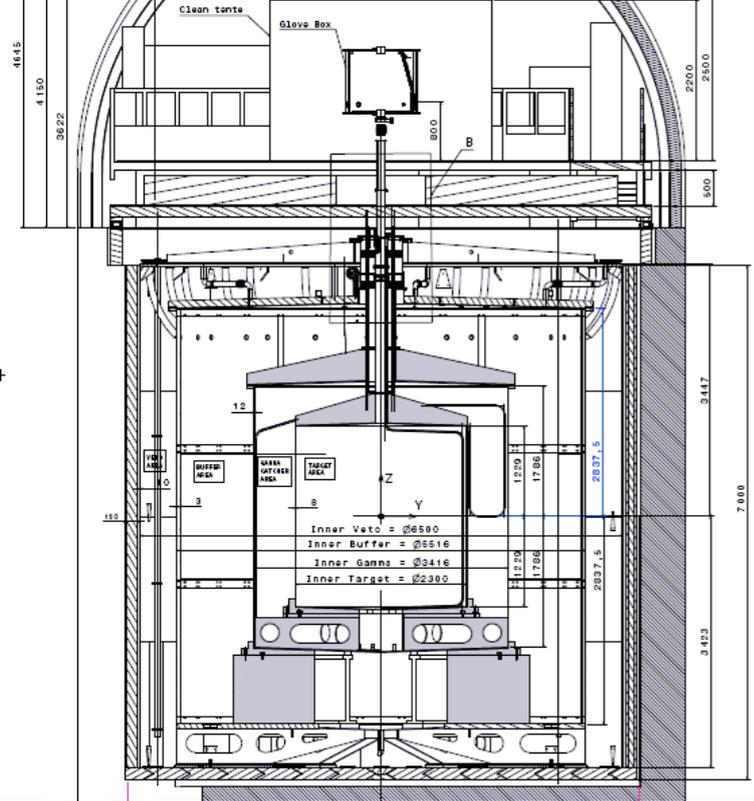
Inner  $\mu$ -Veto (IV)

Liquid-Scintillator

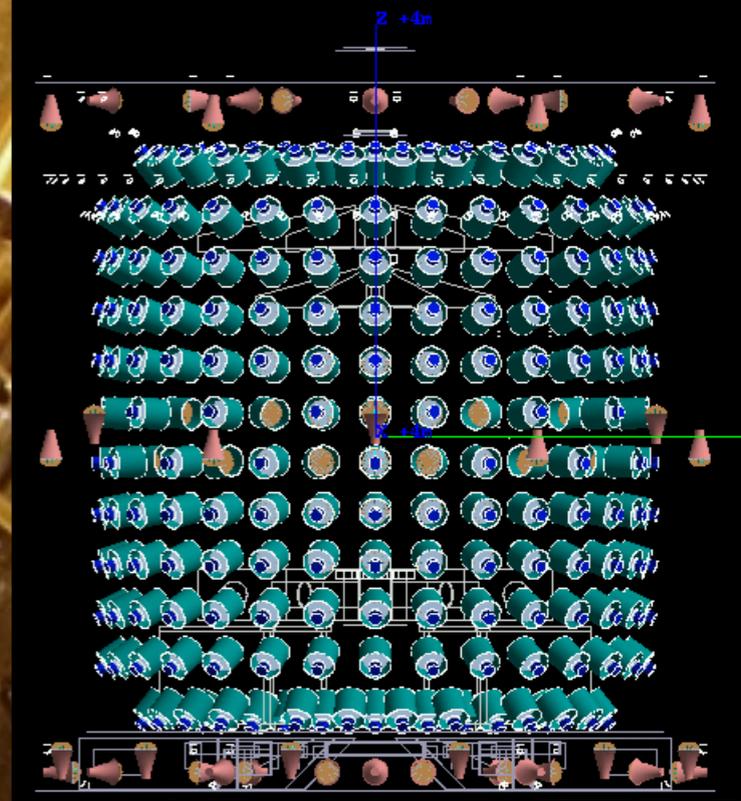
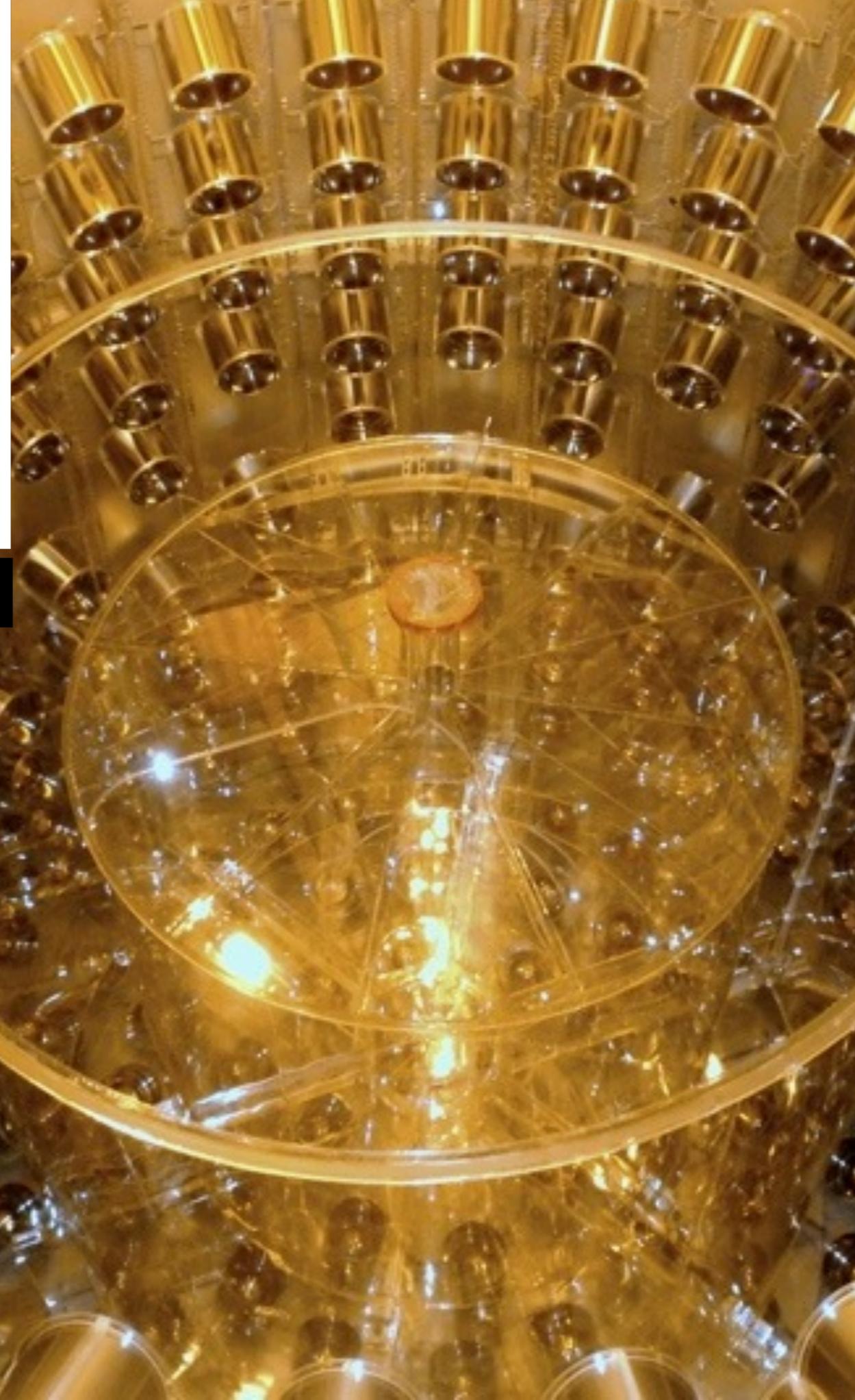
Inert  $\gamma$ -Shield

15cm of steel (around all detector)





engineer's view



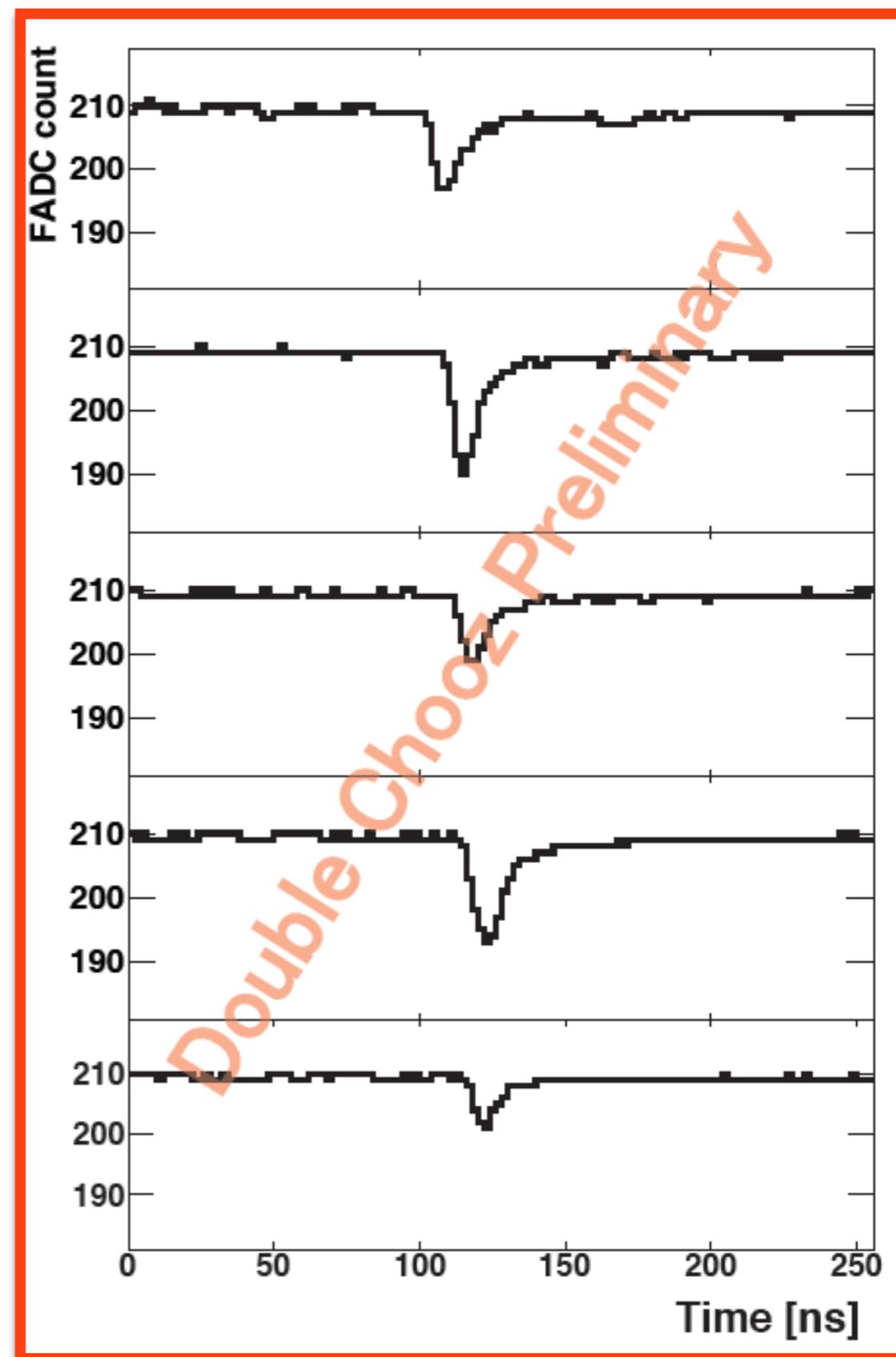
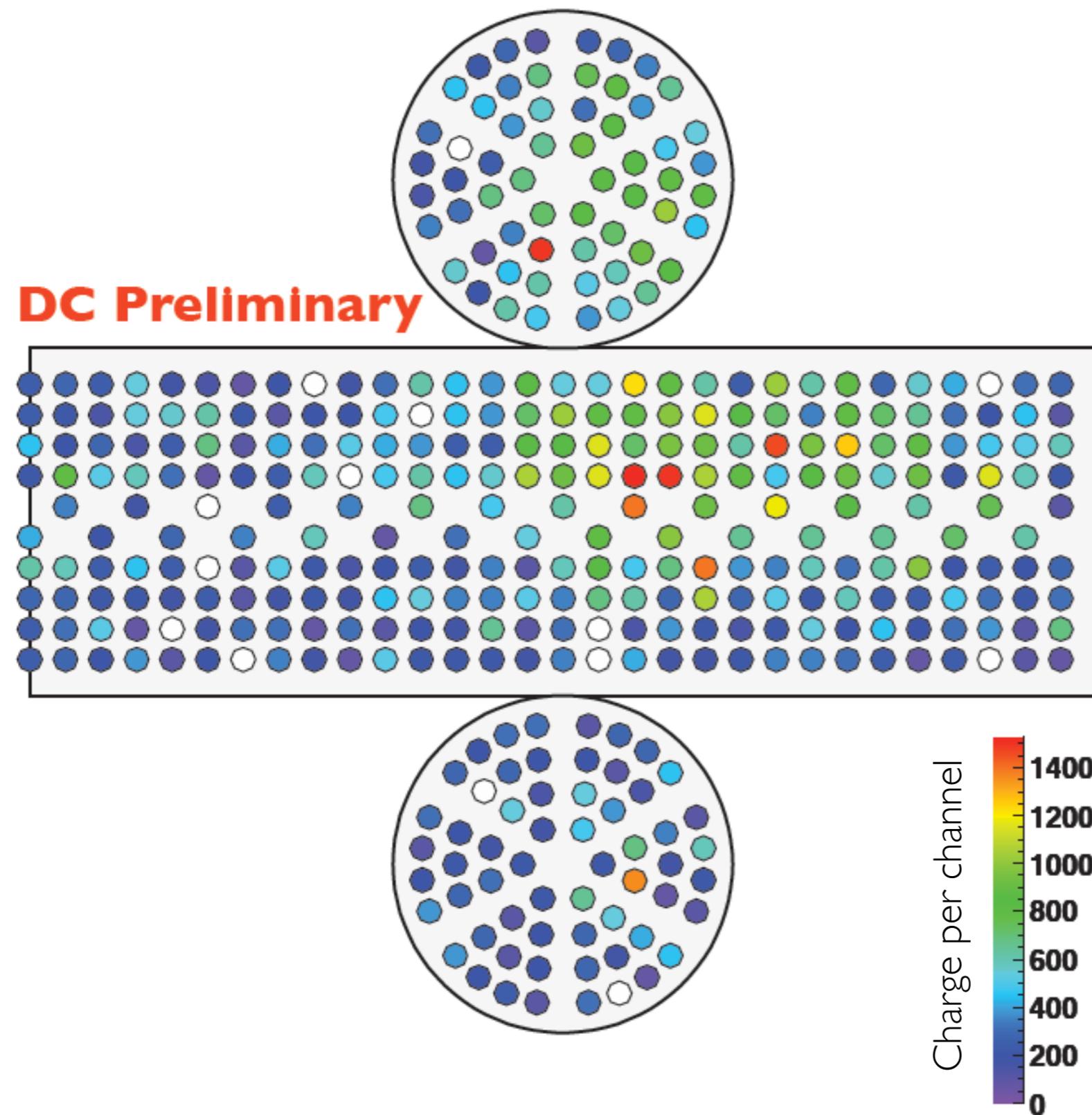
MC's view

our favourite view...

readout...

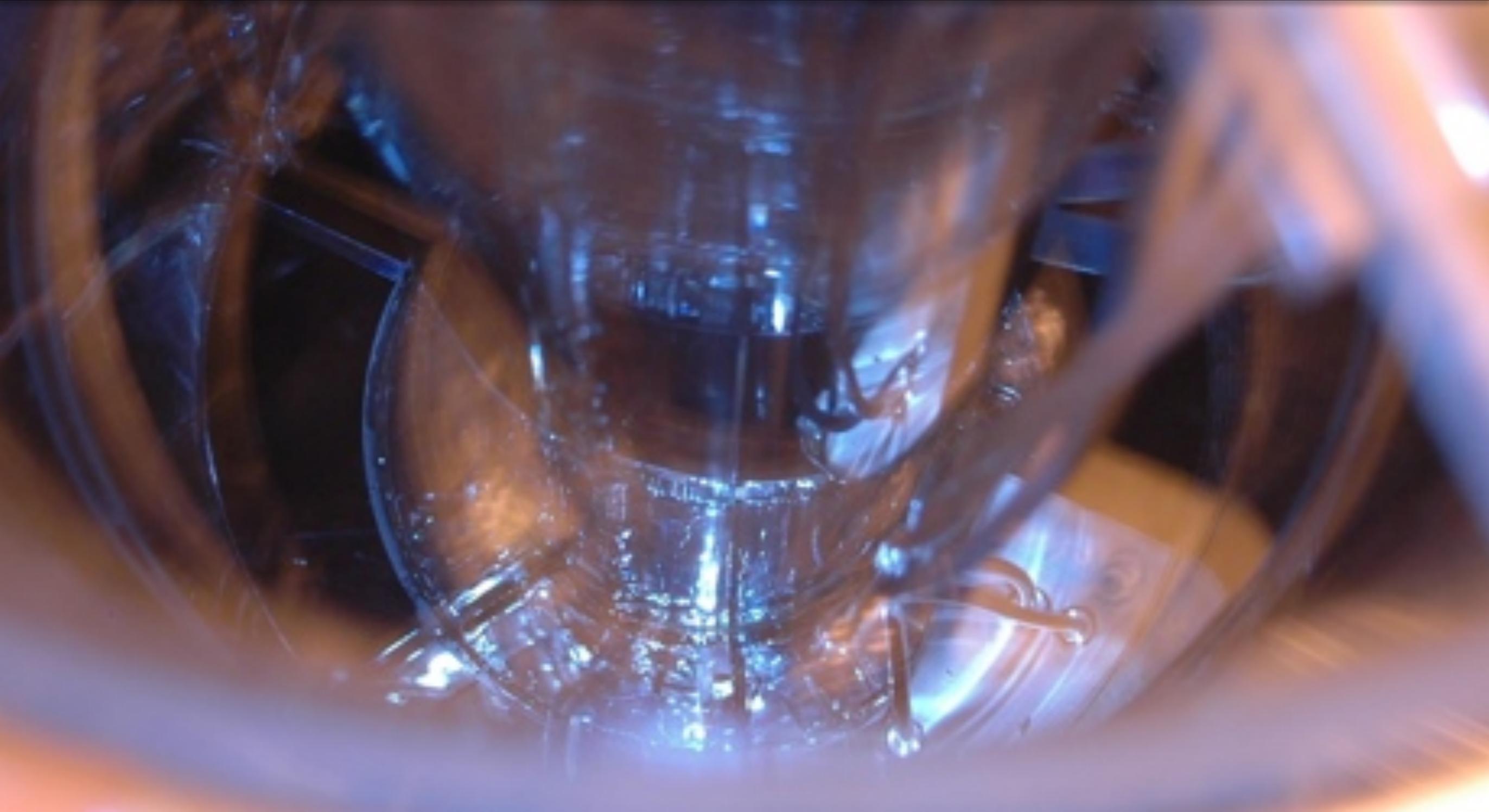


## a point like event (no IV activity)...

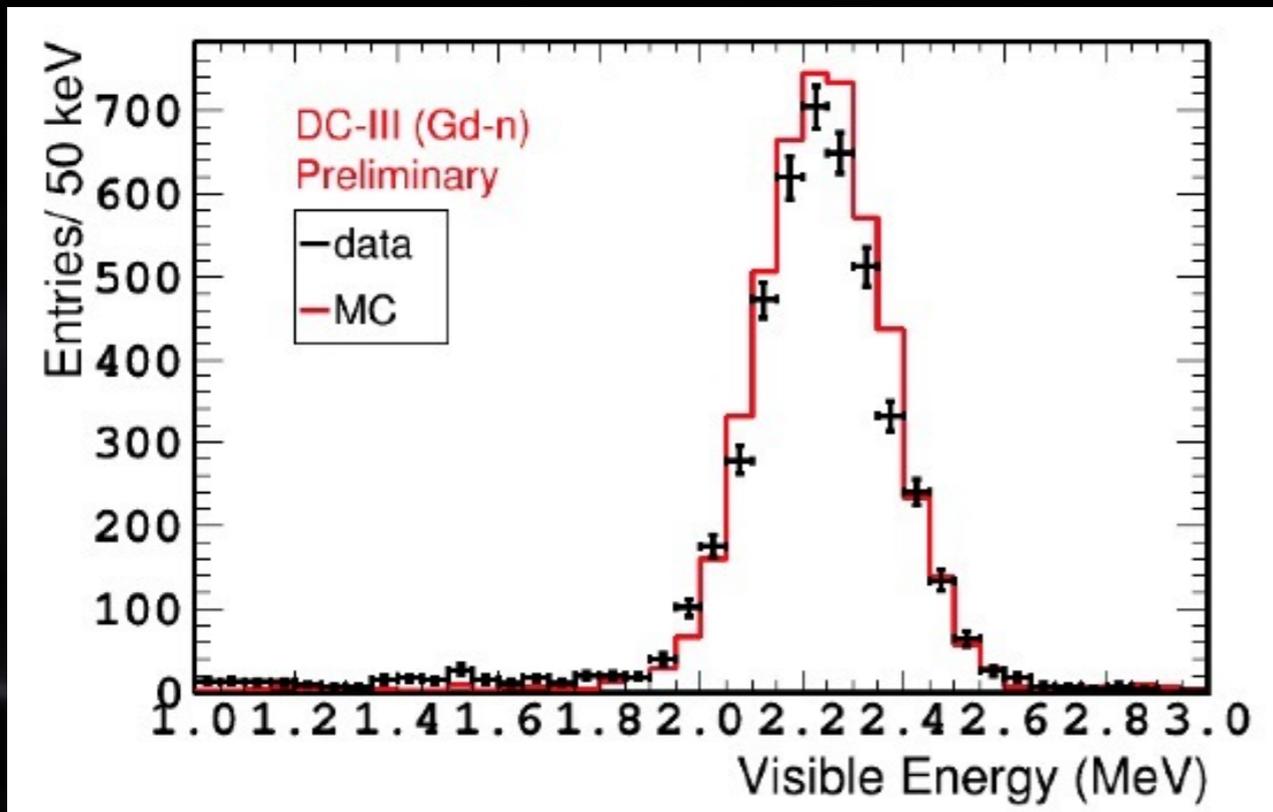
Energy  $\sim 8\text{MeV}$ 

NOTE: all PMTs working (white means: no charge)

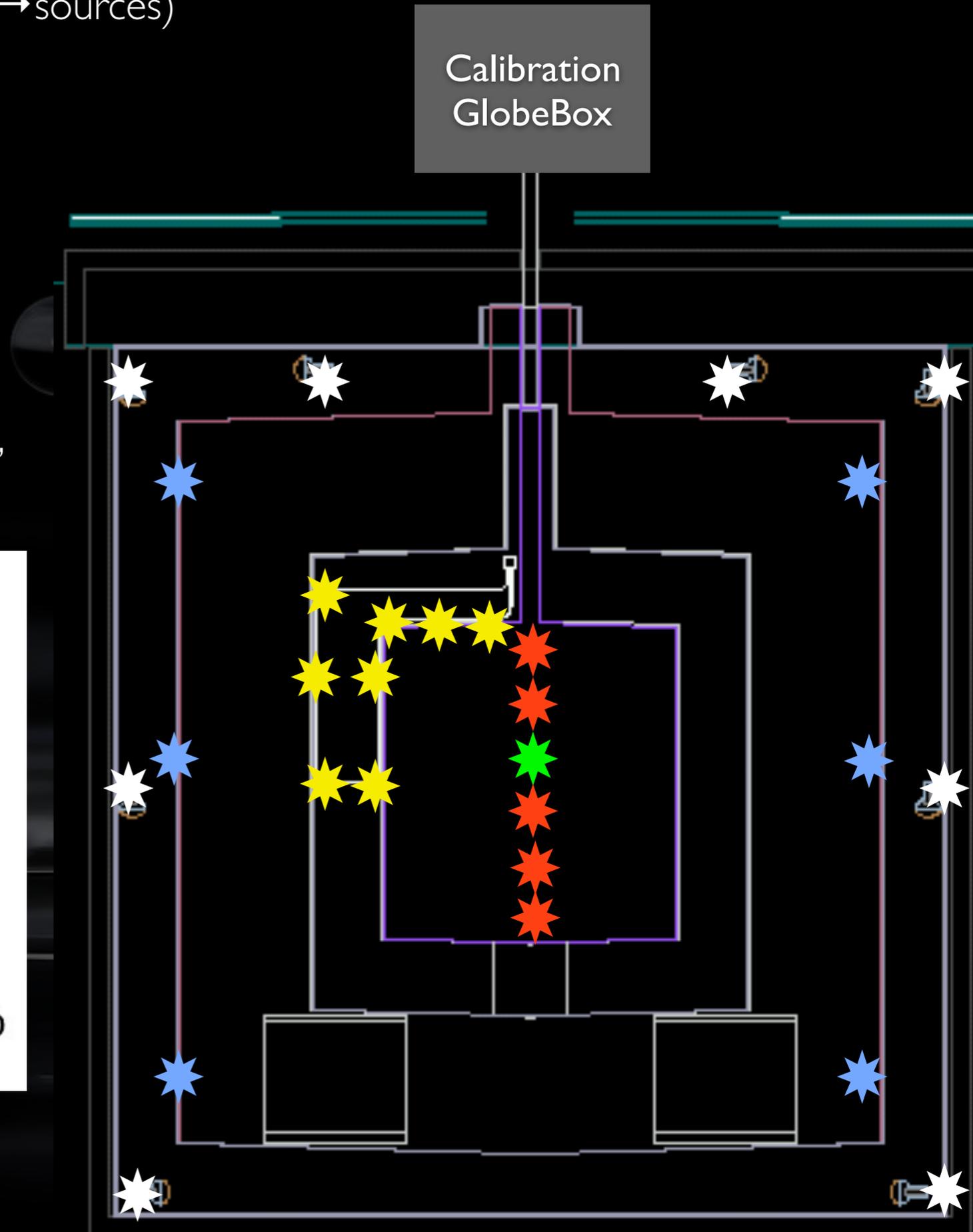
calibration...

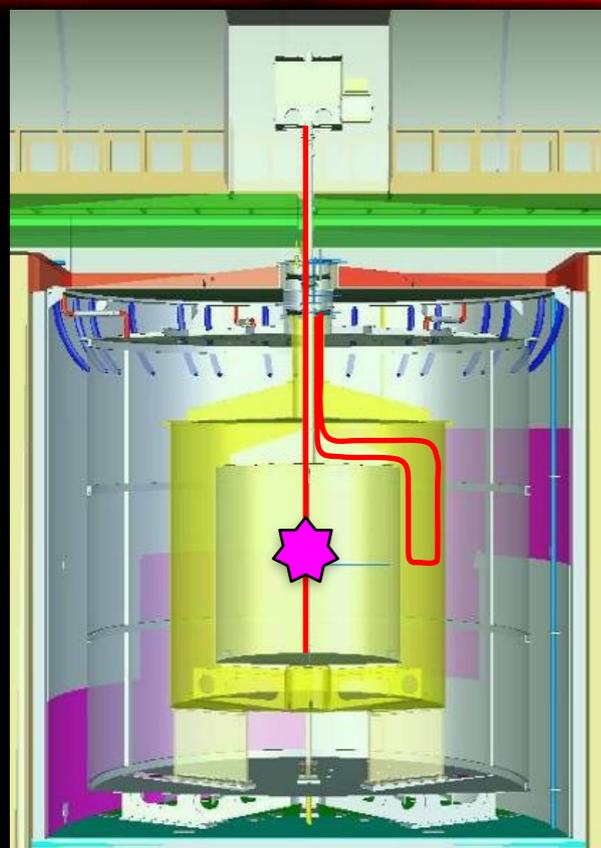
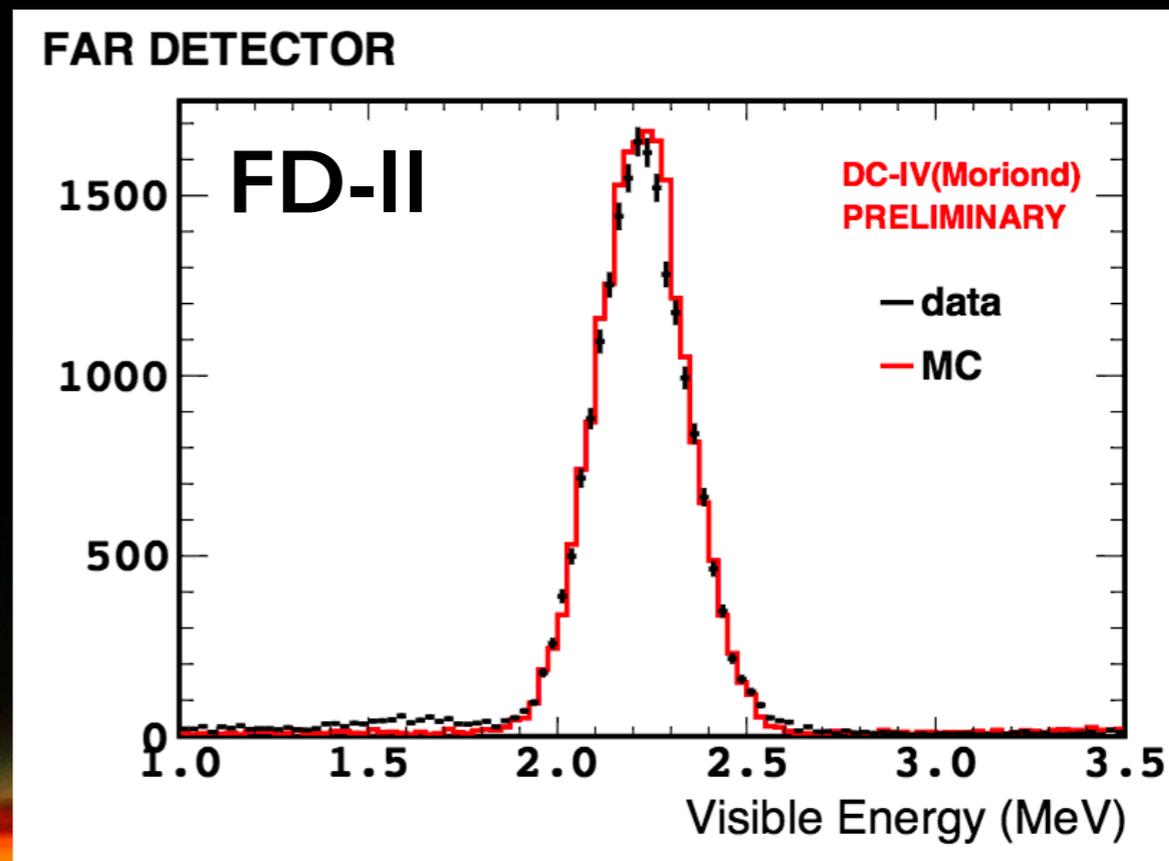
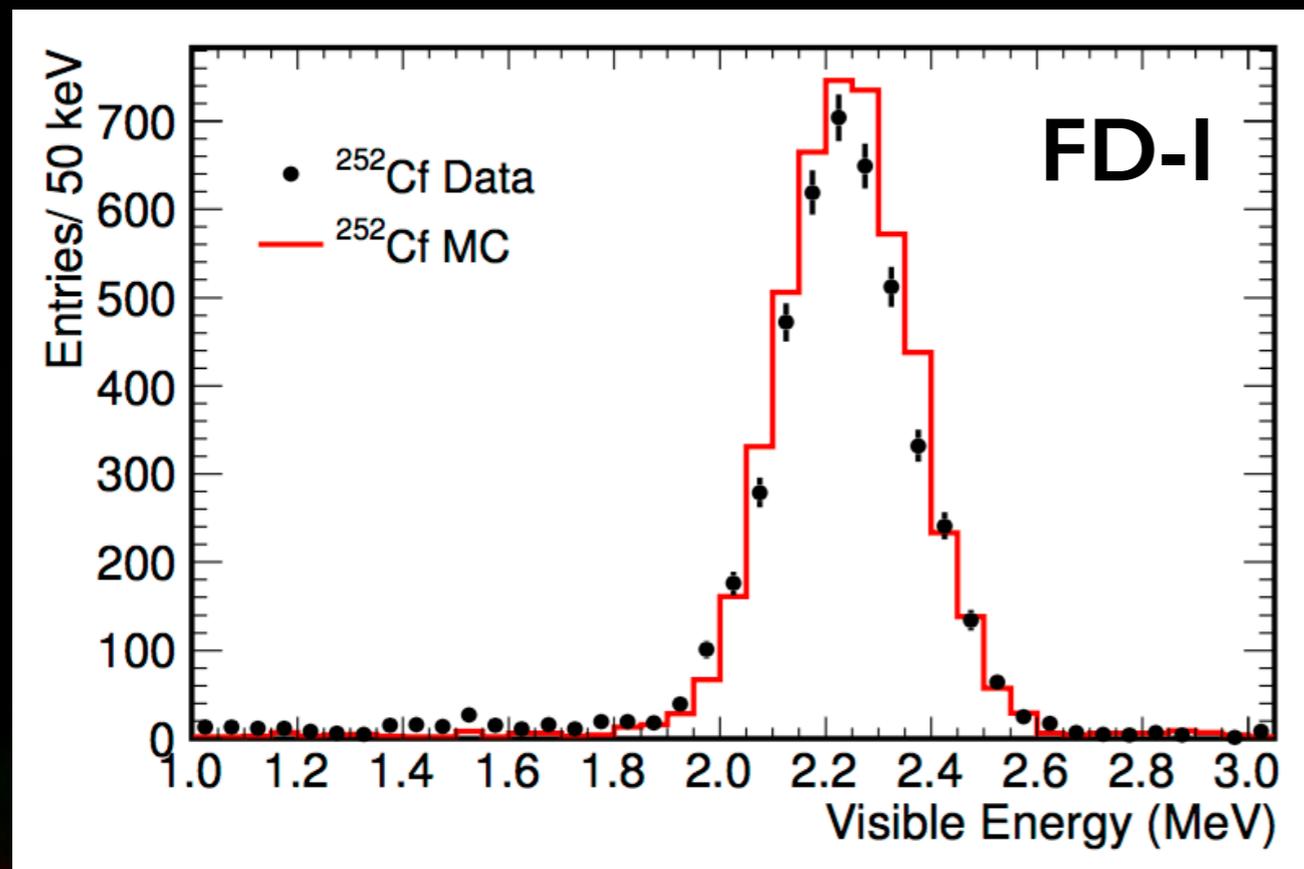


- **principle:** redundancy critical for systematics ( $\rightarrow$  sources)
- **in-built:** light LED (**ID** + **IV**)
- **deployable** ( $^{137}\text{Cs}$ ,  $^{68}\text{Ge}$ ,  $^{60}\text{Co}$ ,  $^{252}\text{Cf}$ , lasers)
  - **z-axis** ( $\rightarrow$   $\nu$ -target sampling)
  - **GC guide-tube** ( $\rightarrow$  GC sampling)
  - (not yet used) **Articulated Arm**
- **natural:** H-n, C-n, Gd-n peaks ( $\mu$ 's fast-n), BiPo, IBD (delay spectrum  $\rightarrow$  validation)

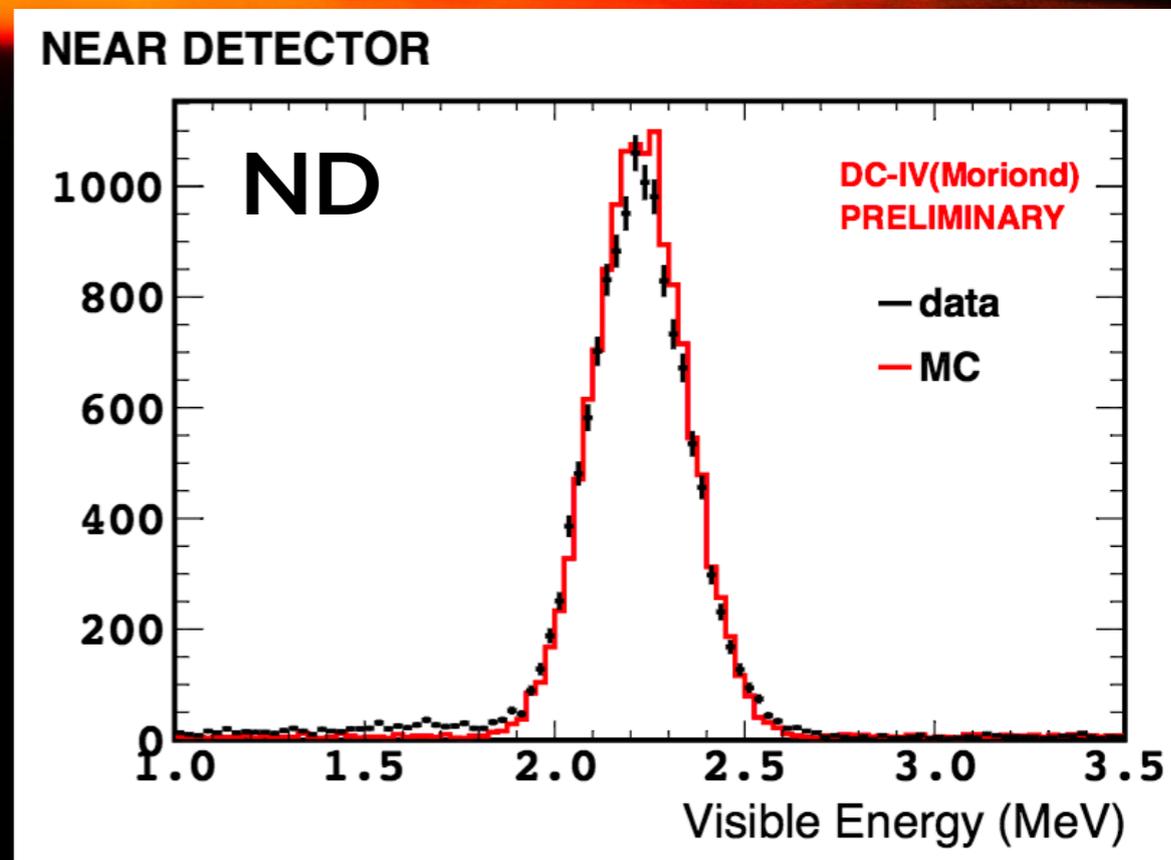


**MeV definition** (H-n peak @ center)  
(our *standard candle*)

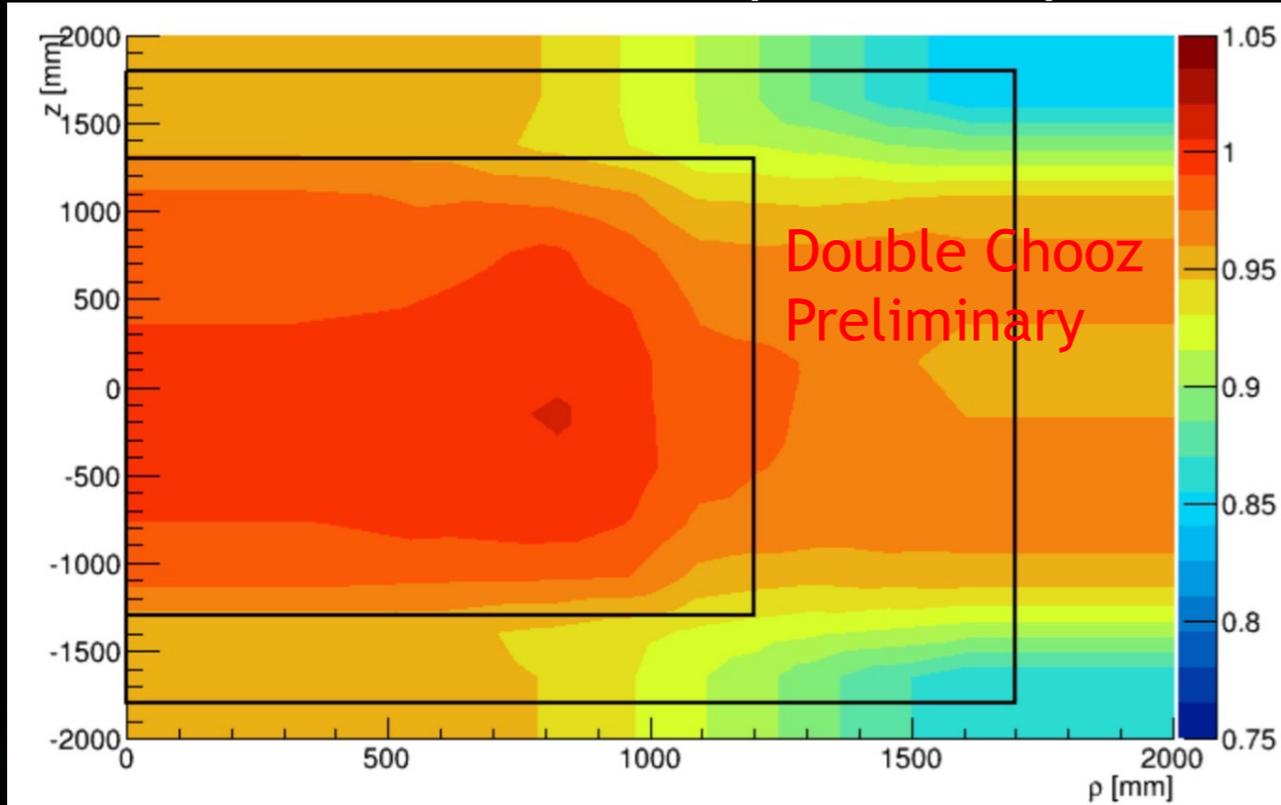




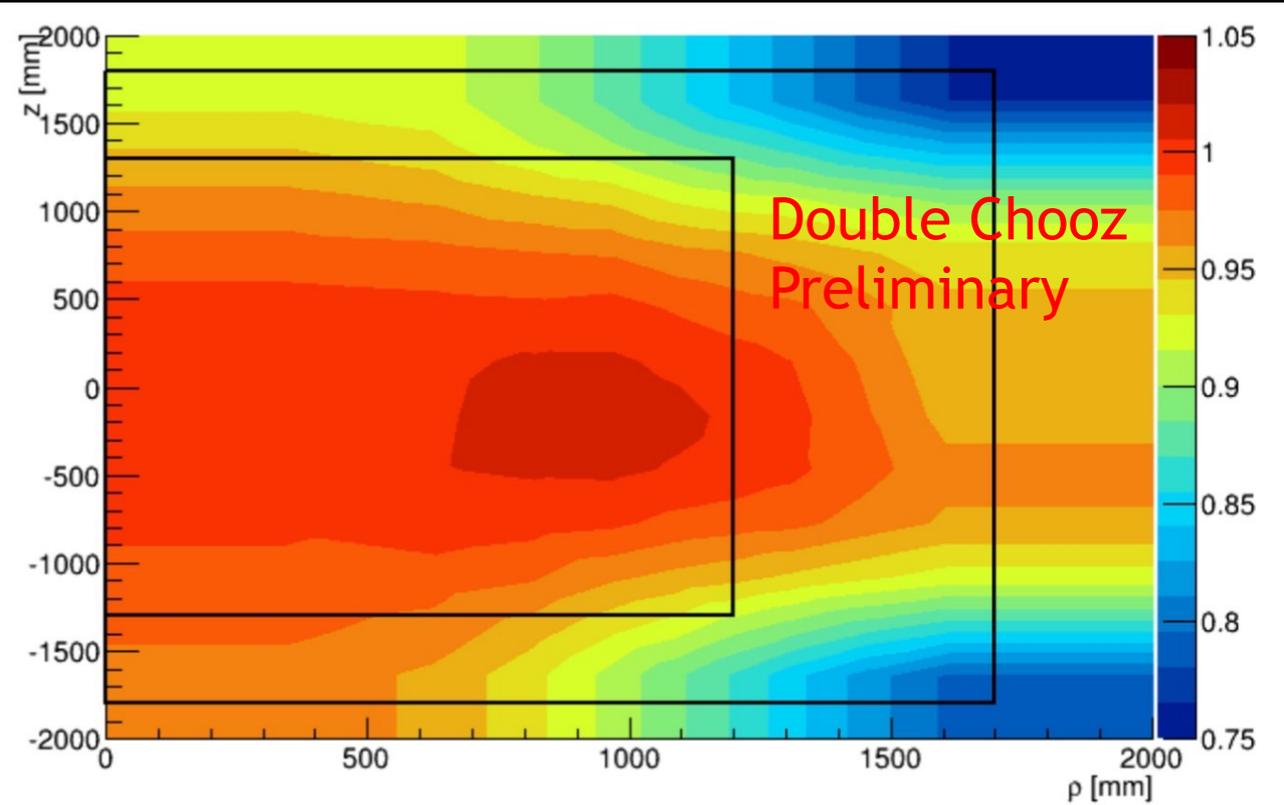
Cf neutron source  
deployed at center



FD-II DATA (9months)

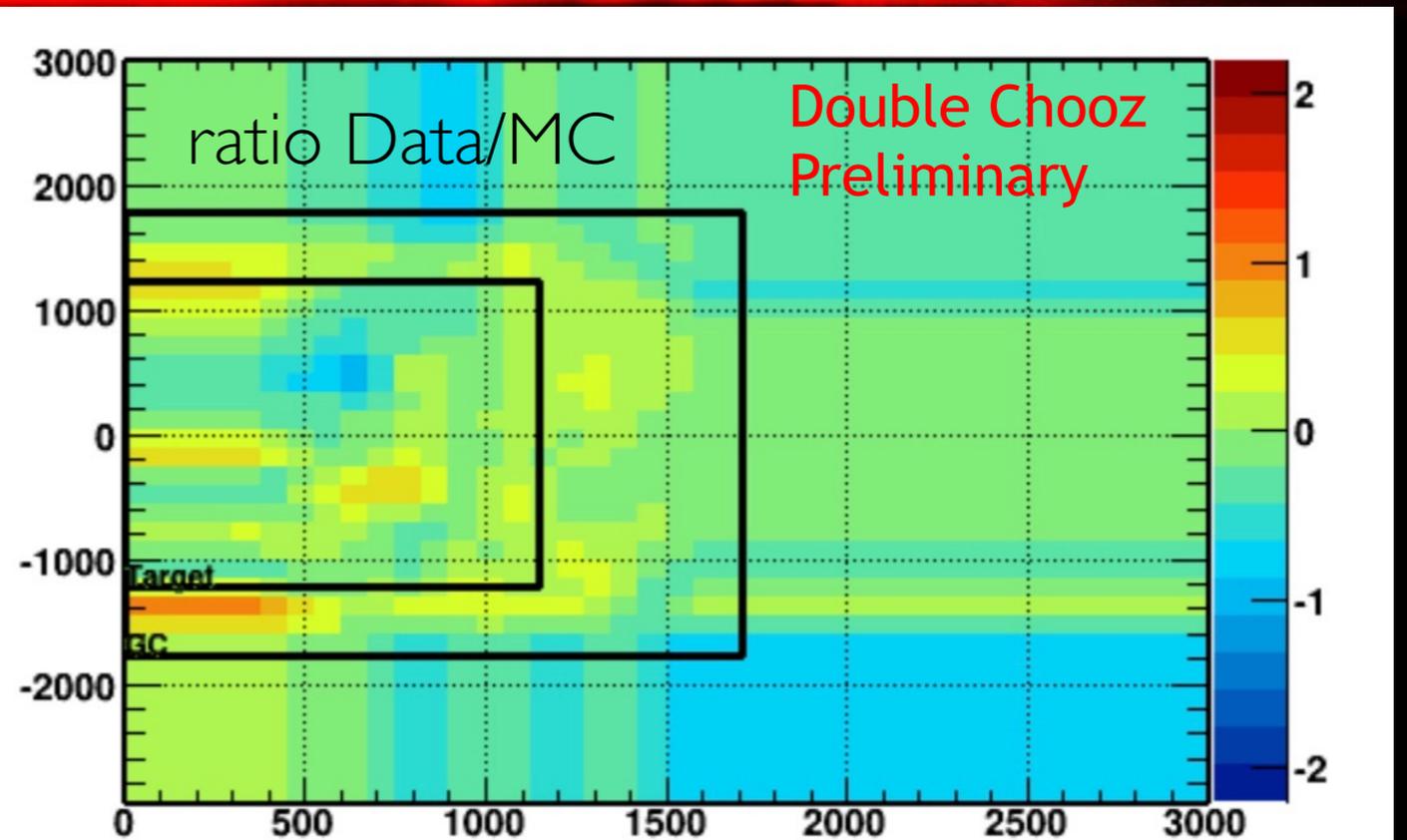


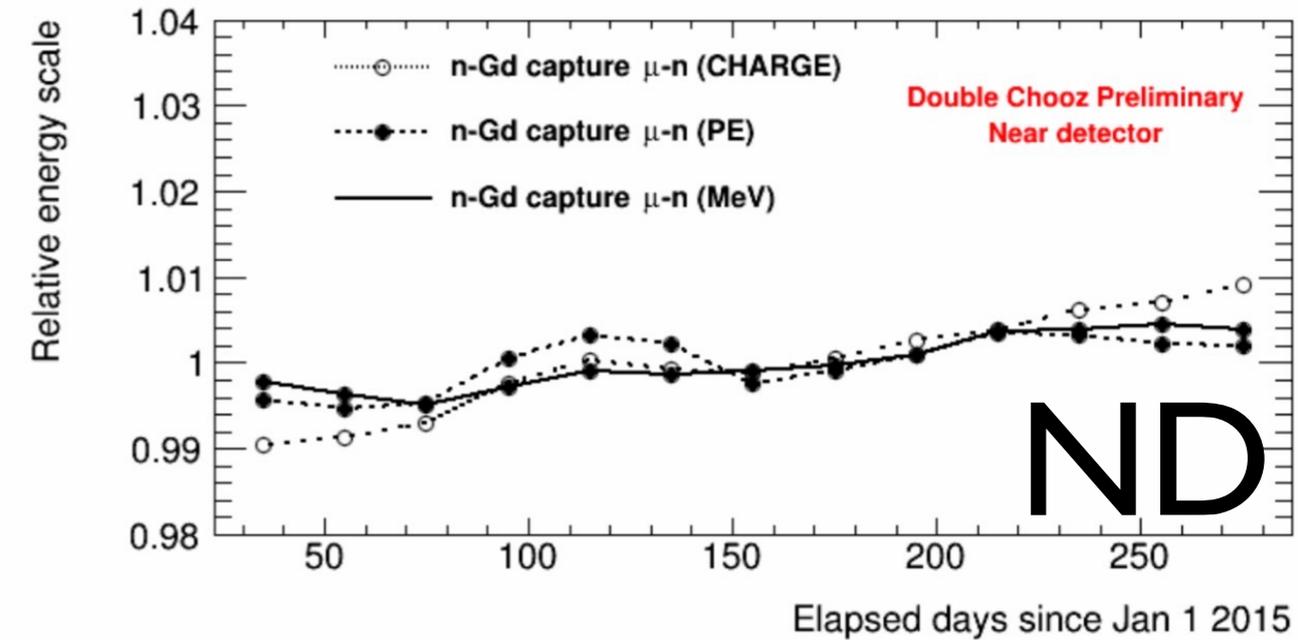
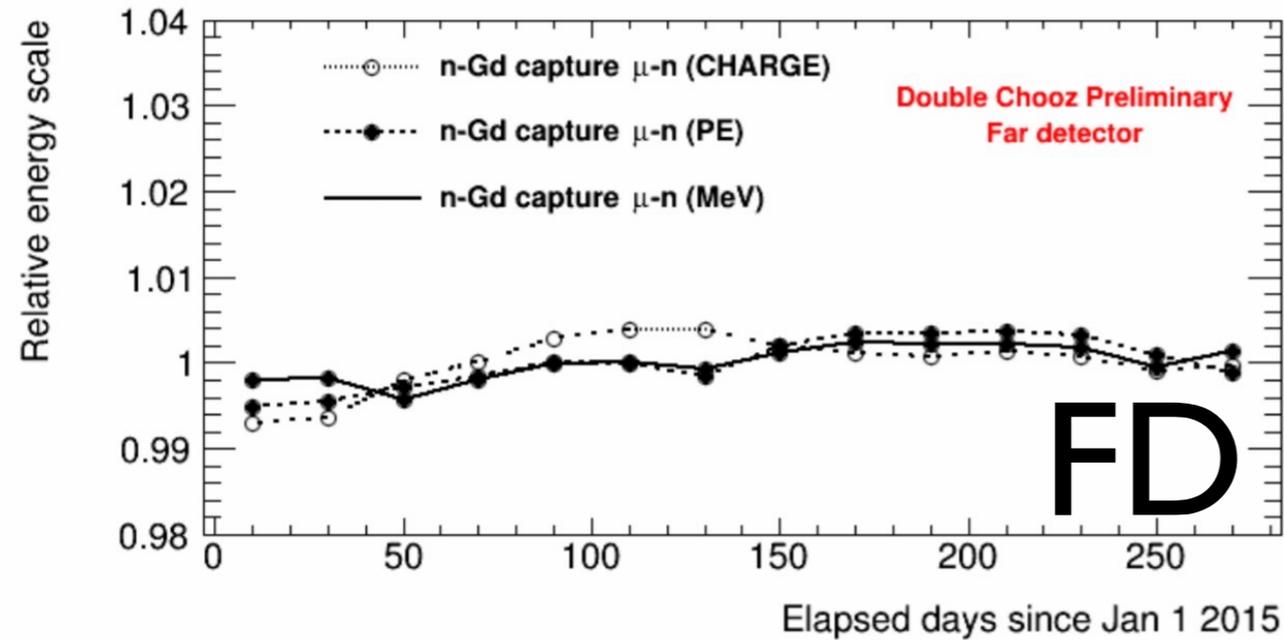
FD-II MC



(swamped by  $\mu$ 's and after- $\mu$  stuff)

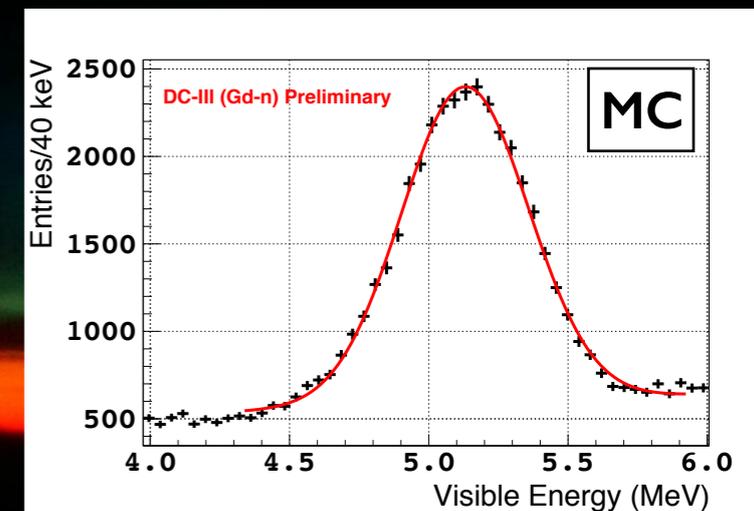
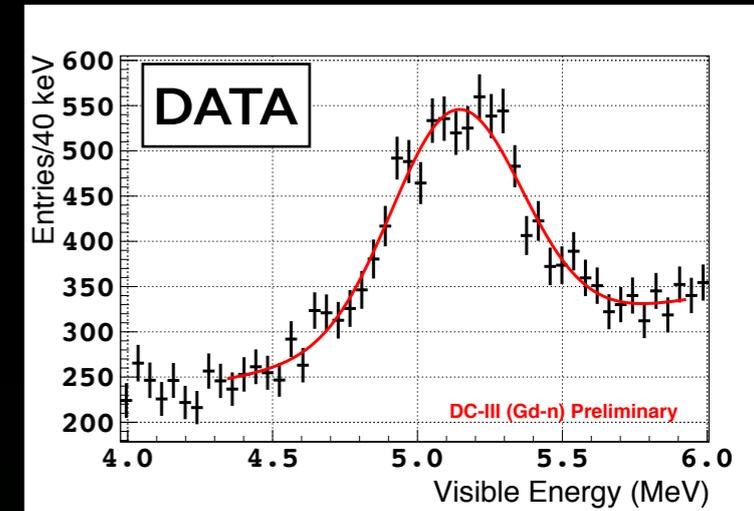
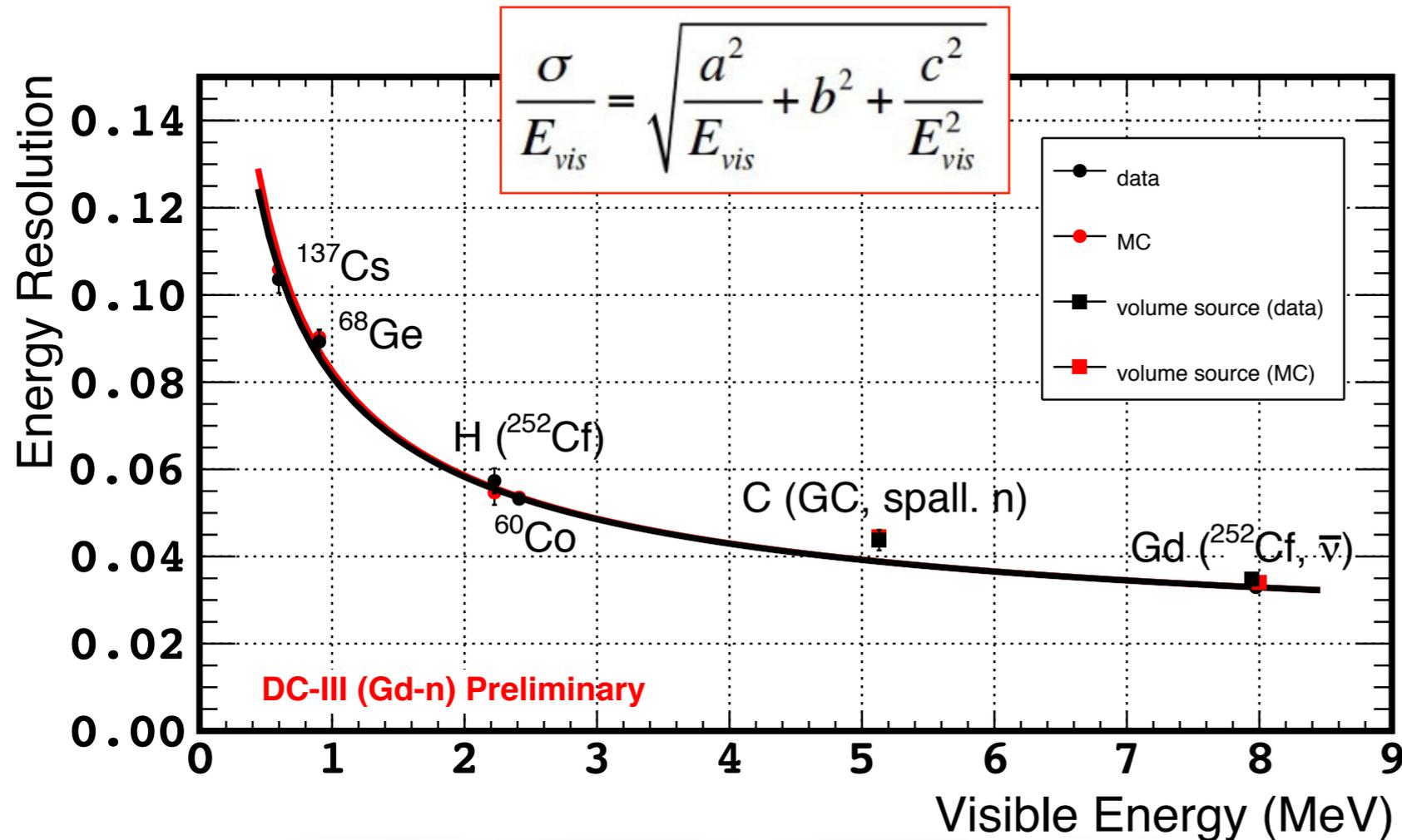
- most DC calibration uses cosmogenic signals (i.e. self-calibrated)
- fast-n H-n captures map:
- ratio DATA /MC: **0.25%<sup>FD</sup>** & **0.39%<sup>ND</sup>**





response increases with time ( $\approx 1\%/year$ )

DC: the only loaded-LS detector not deteriorating?  
(FRoS T-16 workshop @ FNAL)



excellent MC precision & accuracy

a: statistical term  
b: constant term  
c: e.g. electric noise

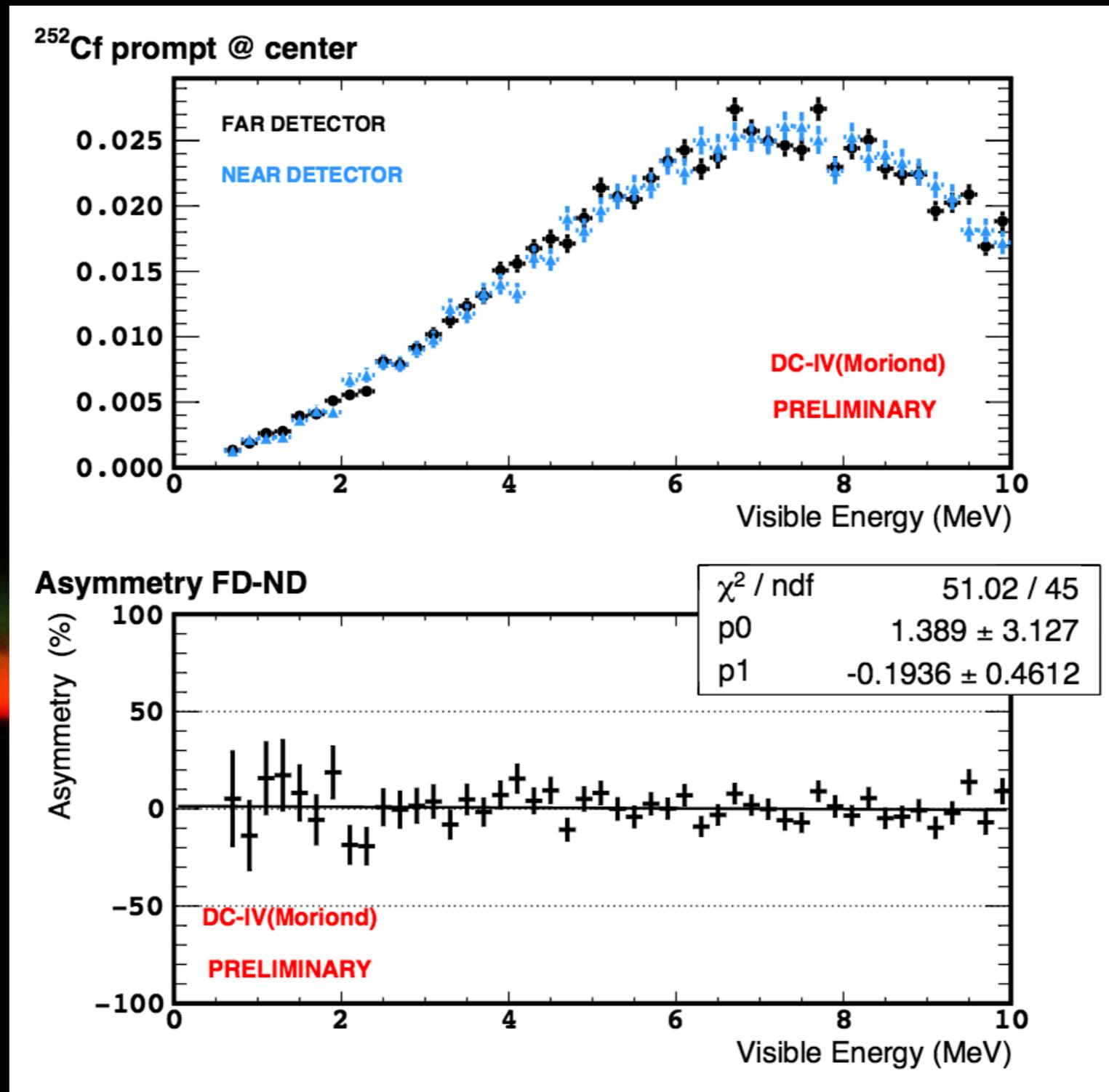
#### Data

a=0.0773±0.0025  
b=0.0182±0.0014  
c=0.0174±0.0107

#### MC

a=0.0770±0.0018  
b=0.0183±0.0011  
c=0.0235±0.0061

- **remarkable agreement data to MC** throughout full energy range
  - identical curves (→ no free knobs in MC)
  - most relevant region for  $\theta_{13}$  is  $\leq 4\text{MeV}$
- **excellent precision:** peak position and widths (highly non-trivial)
  - true for peaks in center or anywhere in NT and GT
  - C-n peak (mainly from GC) → slight different response in GC (worse)
- **constant term of resolution ~0.018** (powerful calorimetry)
  - largely dominated by stochastic term

prompt energy of  $^{252}\text{Cf}$  data (same source)...

- $^{252}\text{Cf}$  emits  $\sim 10$   $\gamma$  with 1 MeV in average.
- **comparison FD to ND data (no MC)  $^{252}\text{Cf}$  @ the center**  $\rightarrow$  identical

more calibration needed (poor-ish statistics)

our **new** analyses (I,II,III  $\rightarrow$  **IV**)...

@51st Moriond (March 2016)

**PRELIMINARY**

(further details in our 2014 paper  $\rightarrow$  similar)

Improved measurements of the neutrino mixing angle  $\theta_{13}$  with the Double Chooz detector

---

Double Chooz Collaboration

arXiv:1406.7763v1 [hep-ex] 30 Jun 2014

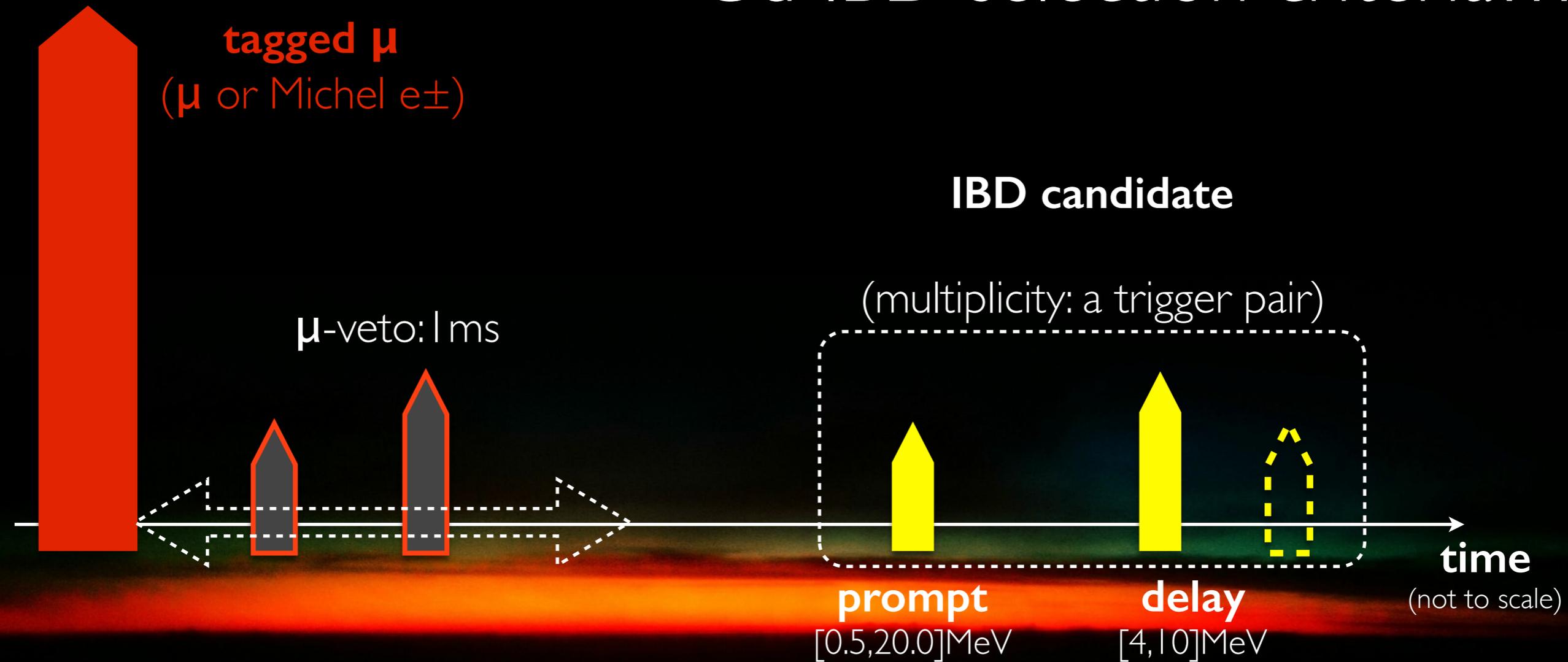
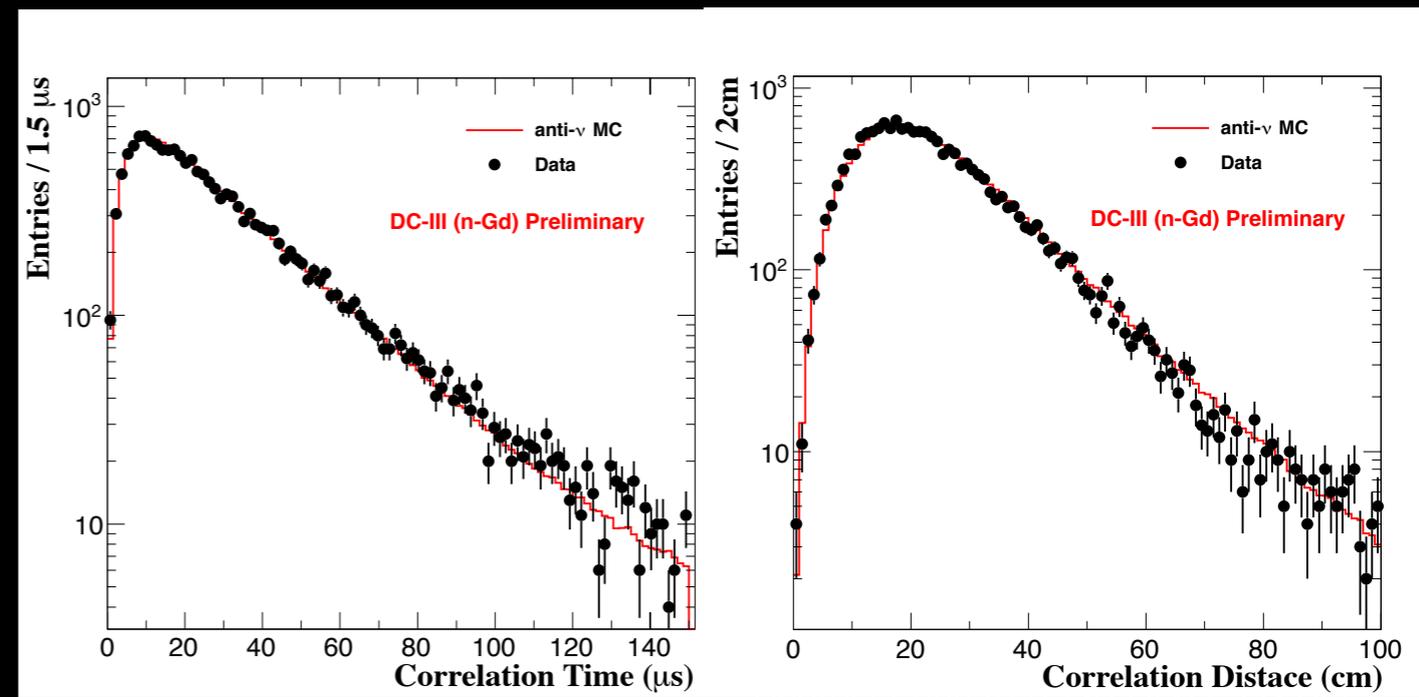
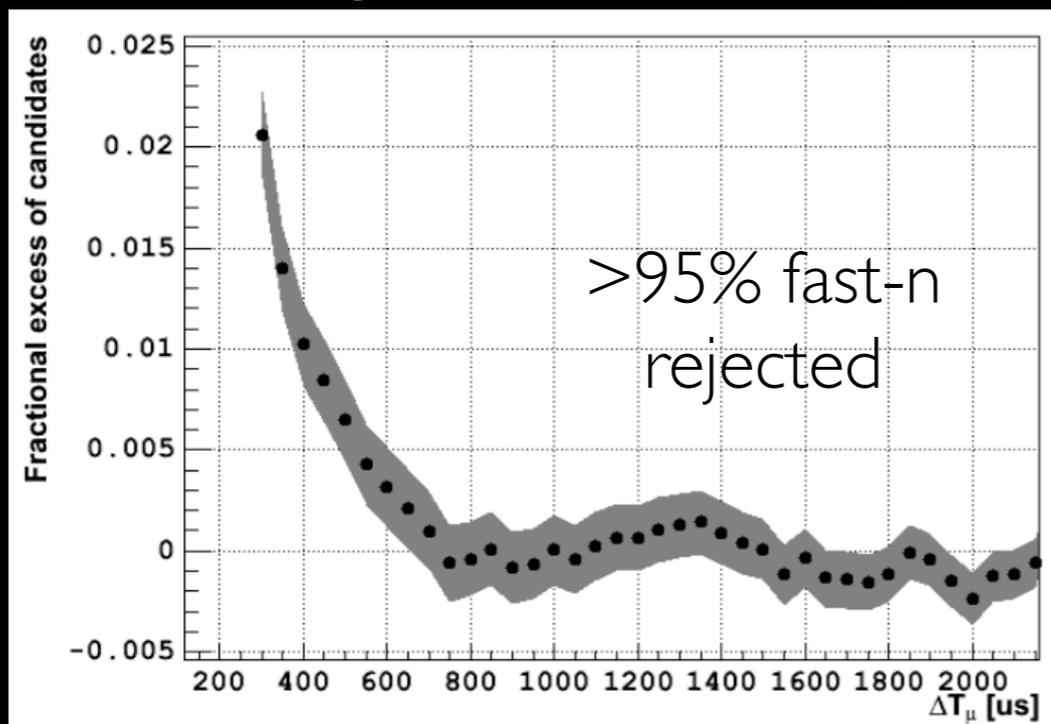
systematics	single detector (SD) (%)	multi-detector (MD) (%)
$\delta(\text{detection})$	$\sim 2.0$ ( $\sim 0.4^{\text{DC-IV}}$ )	$\sim 0.2$ ( $\sim 0.3^{\text{DC-IV}}$ ) (identical detectors)
$\delta(\text{flux})$	$\sim 3.0$ ( $\sim 1.7^{\text{DC-IV}}$ via Bugey4)	$\leq 0.5$ ( $< 0.1^{\text{DC-IV}}$ ) ( $\sim$ iso-flux site)
$\delta(\text{background})$	$\leq 1.0$ ( $\sim 0.4^{\text{DC-IV}}$ ) (active BG rejection)	$\leq 1.0$ ( $\sim 0.4^{\text{DC-IV}}$ ) (active BG rejection)

**DC challenge: statistics** (systematics are too good)

BG & detection systematics dominated by statistics too (this analysis)

let's get those neutrinos...

— event selection —

after- $\mu$  correlated IBDsIBD time/space ( $e^+ \sim n$ ) correlation

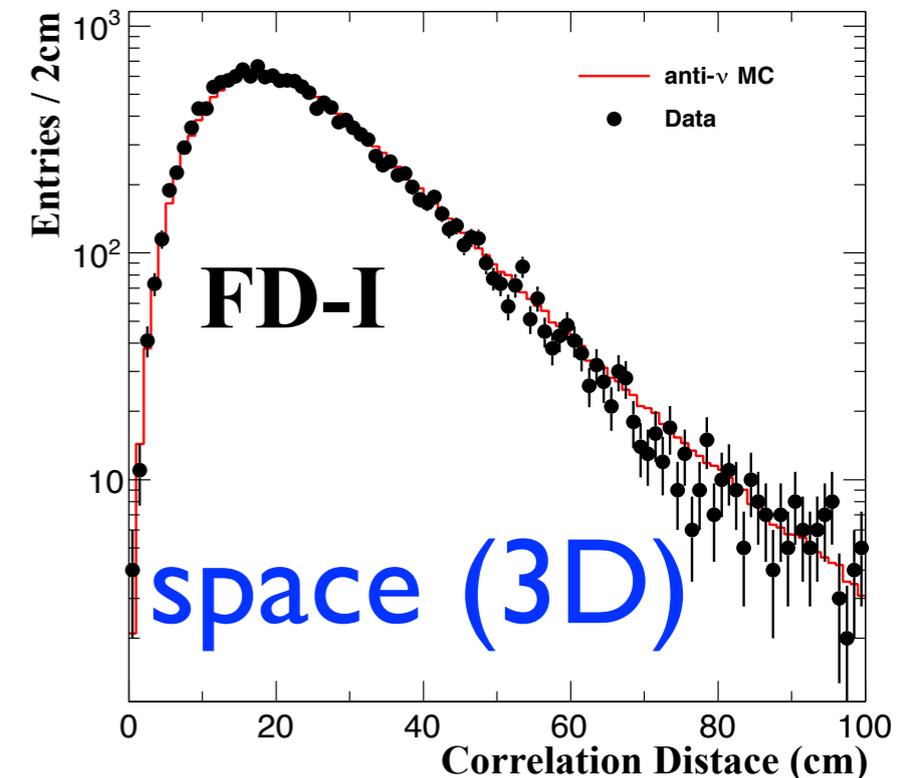
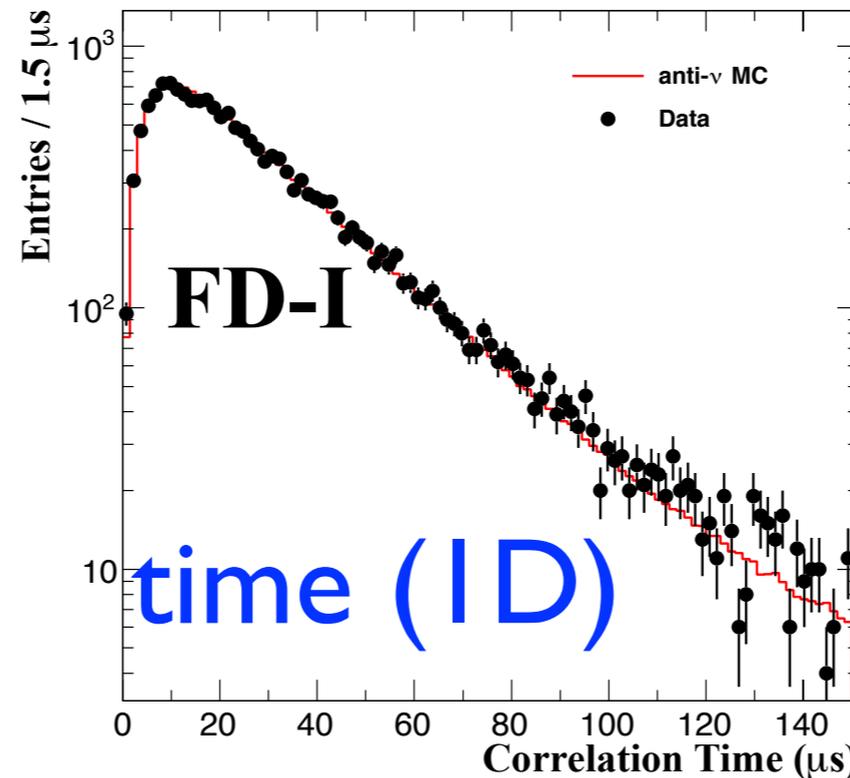
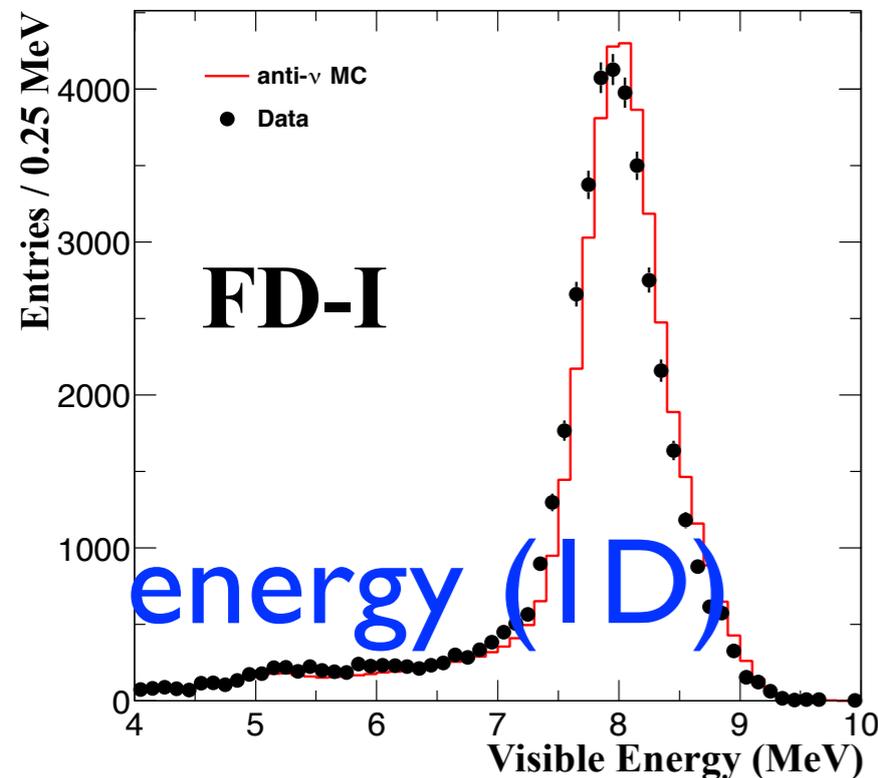
# IBD coincidence: 5D definition

stunning data-MC agreement

Delayed signal energy  
 $4 < E_{\text{vis}} < 10 \text{ MeV}$

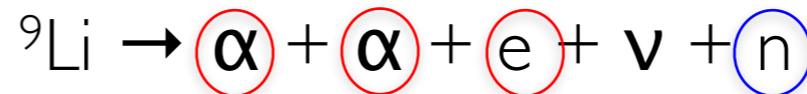
Correlation time  
 $0.5 < \Delta T < 150 \mu\text{sec}$

Correlation distance  
 $\Delta R < 100 \text{ cm}$



⇒ Remaining BG...

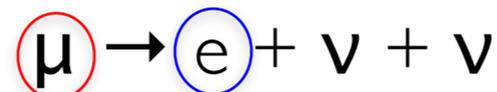
Cosmogenic  $\beta$ -n emitter:



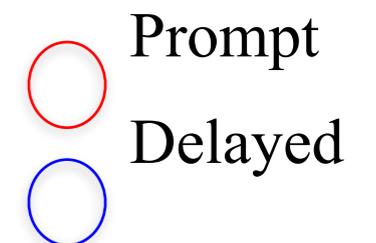
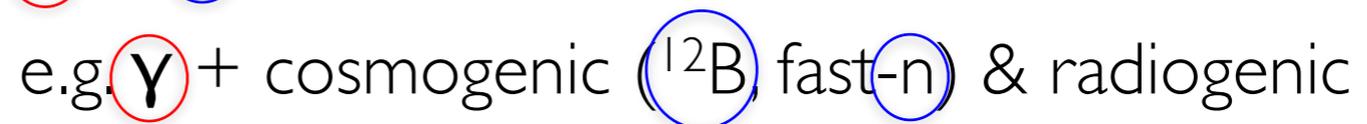
Fast neutron:



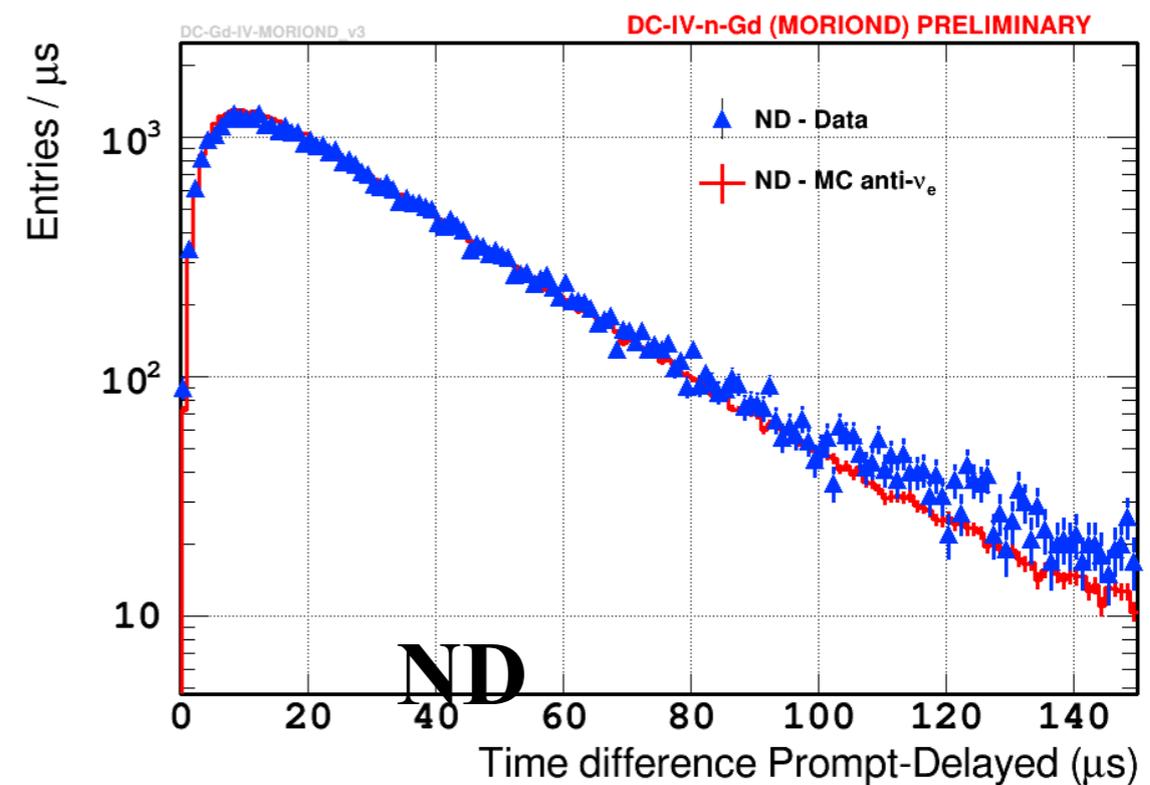
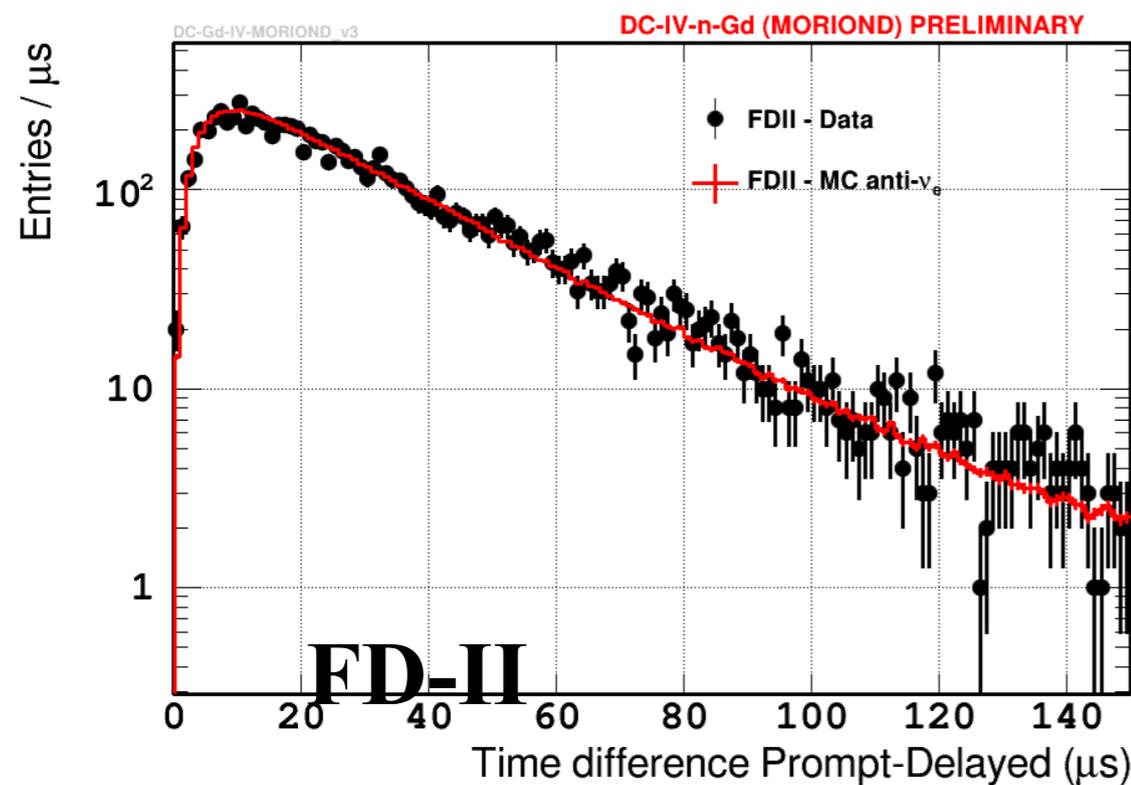
Stop- $\mu$ :



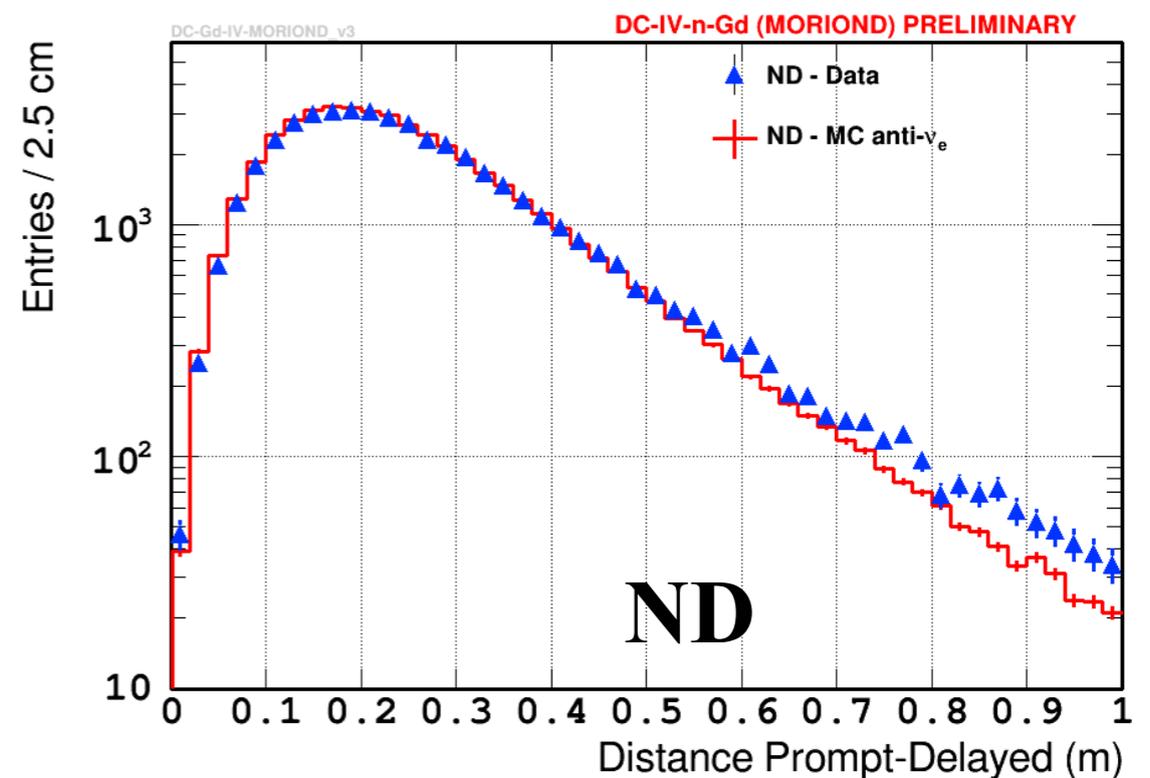
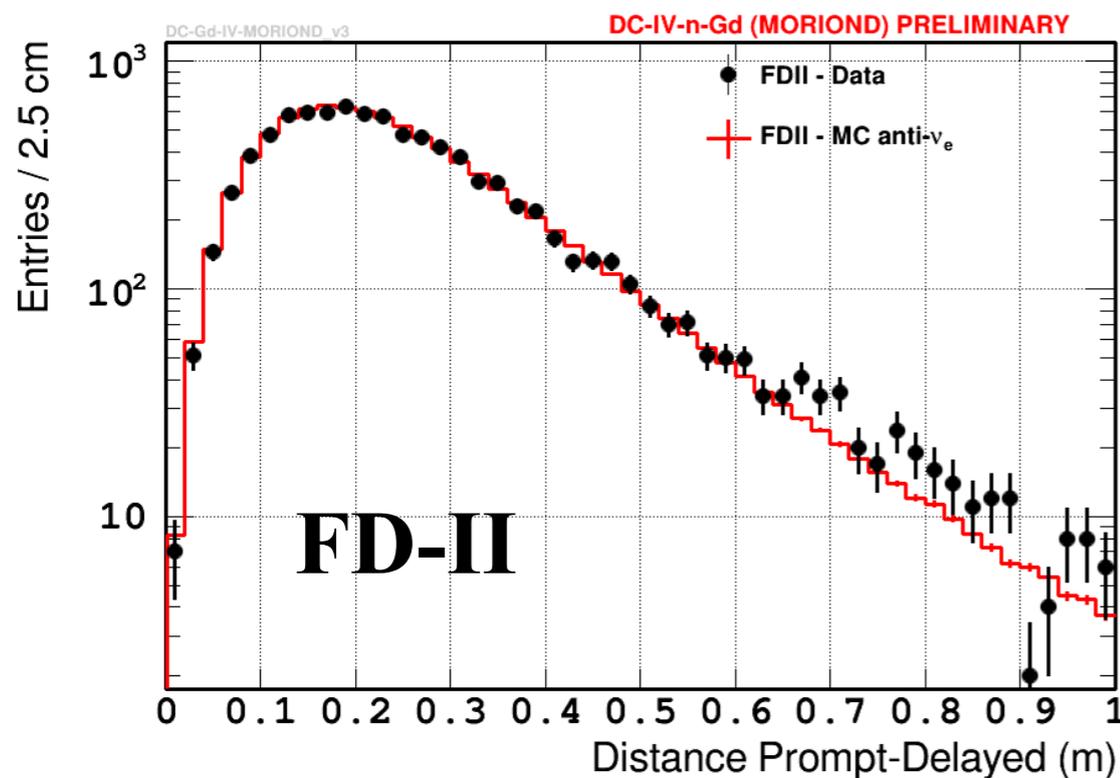
Accidental coincidence:



## Correlation prompt-delay in time (1D)...

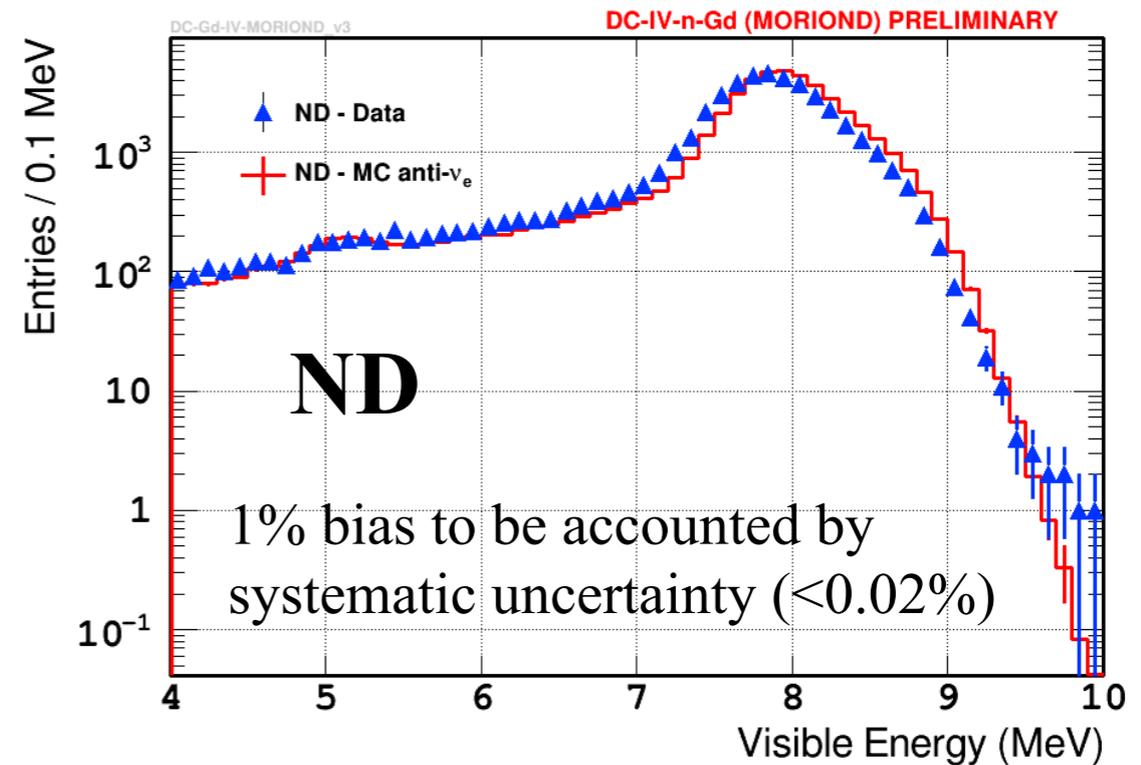
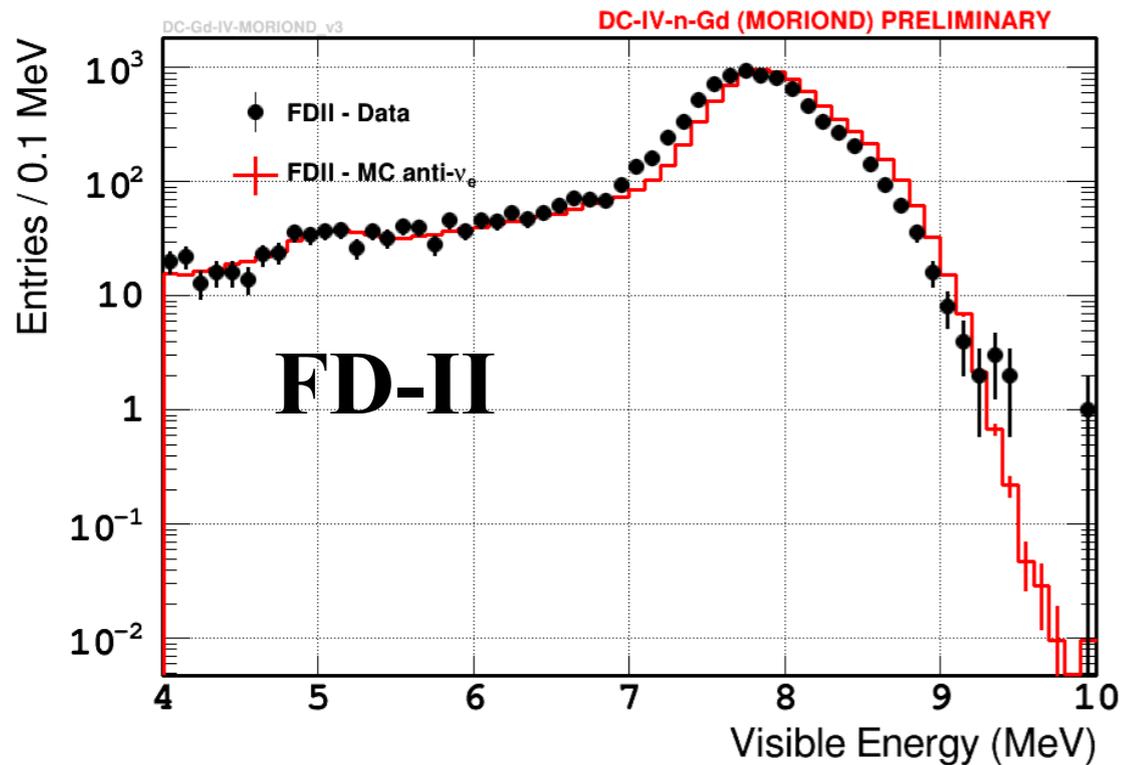


## Correlation prompt-delay in space (3D)...

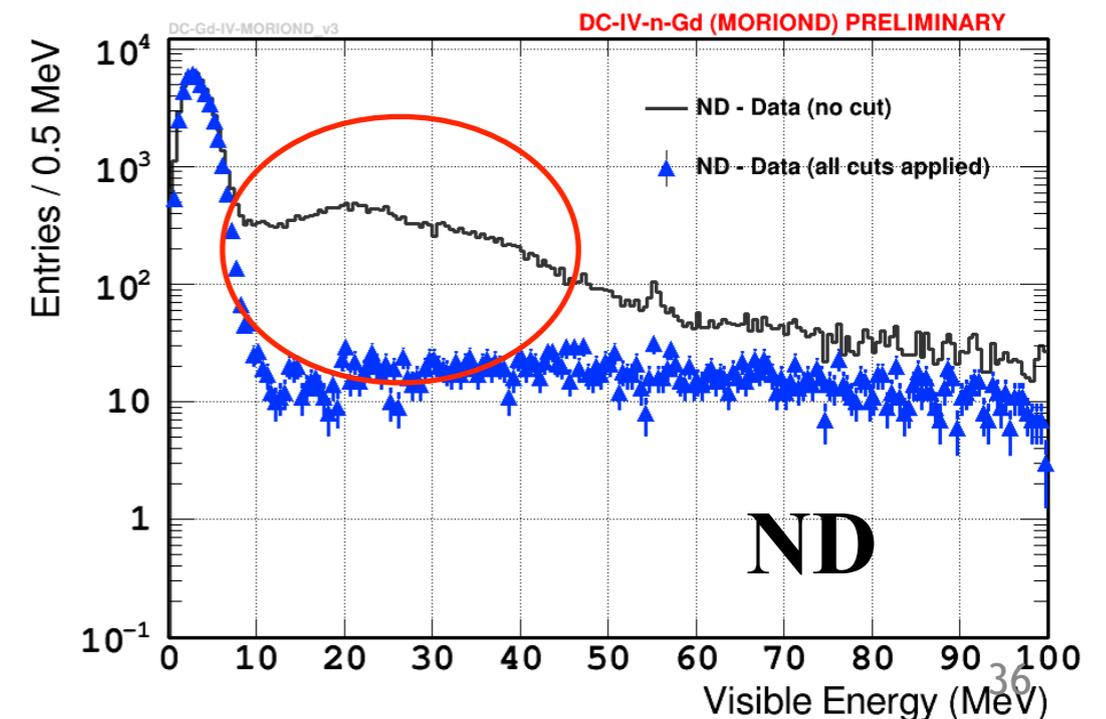


# IBD selection

## Delayed energy



## Prompt energy



contamination of liquid scintillator (leak?) in ND buffer causes buffer stop- $\mu$  background  
 $\Rightarrow$  **almost all such backgrounds are rejected**

by new selections based on

- Energy dependent MaxQ/TotQ cut
- Likelihood at chimney vs. vertex (CPS veto)

# Detection systematics

Double Chooz Preliminary

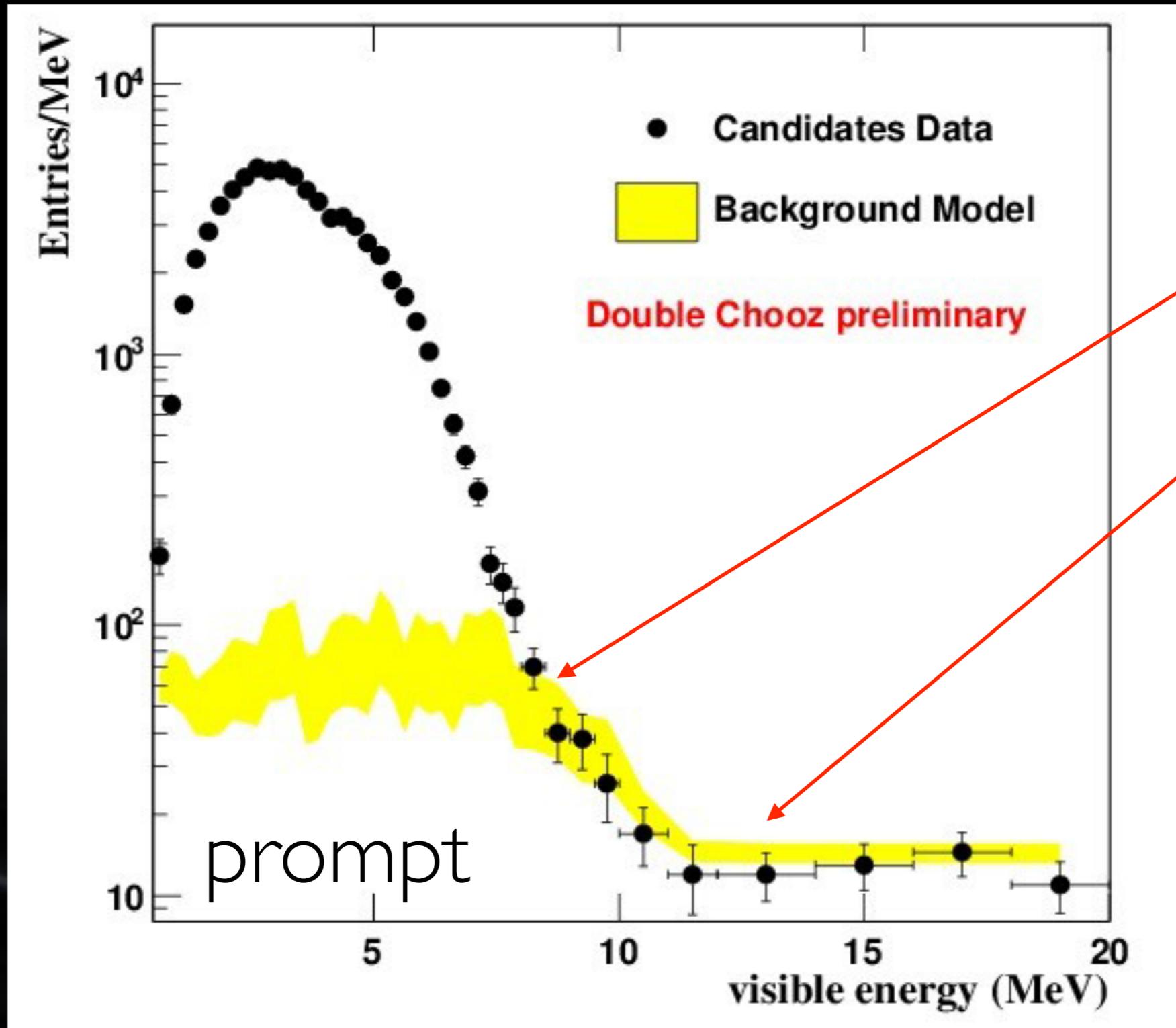
	FD-I	FD-II	ND
BG vetoes (%)	0.11 (0.11)	0.09 (0.09)	0.02 (0.02)
Gd fraction (%)	0.25 (0.14)	0.26 (0.15)	0.28 (0.19)
IBD selection (%)	0.21 (0.21)	0.16 (0.16)	0.07 (0.07)
Spill in/out (%)	0.27 (0)	0.27 (0)	0.27 (0)
Proton number (%)	0.30 (0)	0.30 (0)	0.30 (0)
<b>Total (%)</b>	<b>0.49 (0.26)</b>	<b>0.47 (0.22)</b>	<b>0.38 (0.15)</b>

Numbers in parentheses are uncorrelated uncertainties in multi-detectors analysis (FD-I, FD-II and ND)

**ND and FD identical** within statistical dominated errors (**0.29%**)



# BACKGROUNDS



• **<sup>9</sup>Li (+ a little <sup>8</sup>He)**

(dominant & knowledge @ 20%)

• **fast-n (+ little stopped- $\mu$ s)**

(still visible & knowledge @ <10%)

• **stopping- $\mu$** : mostly rejected

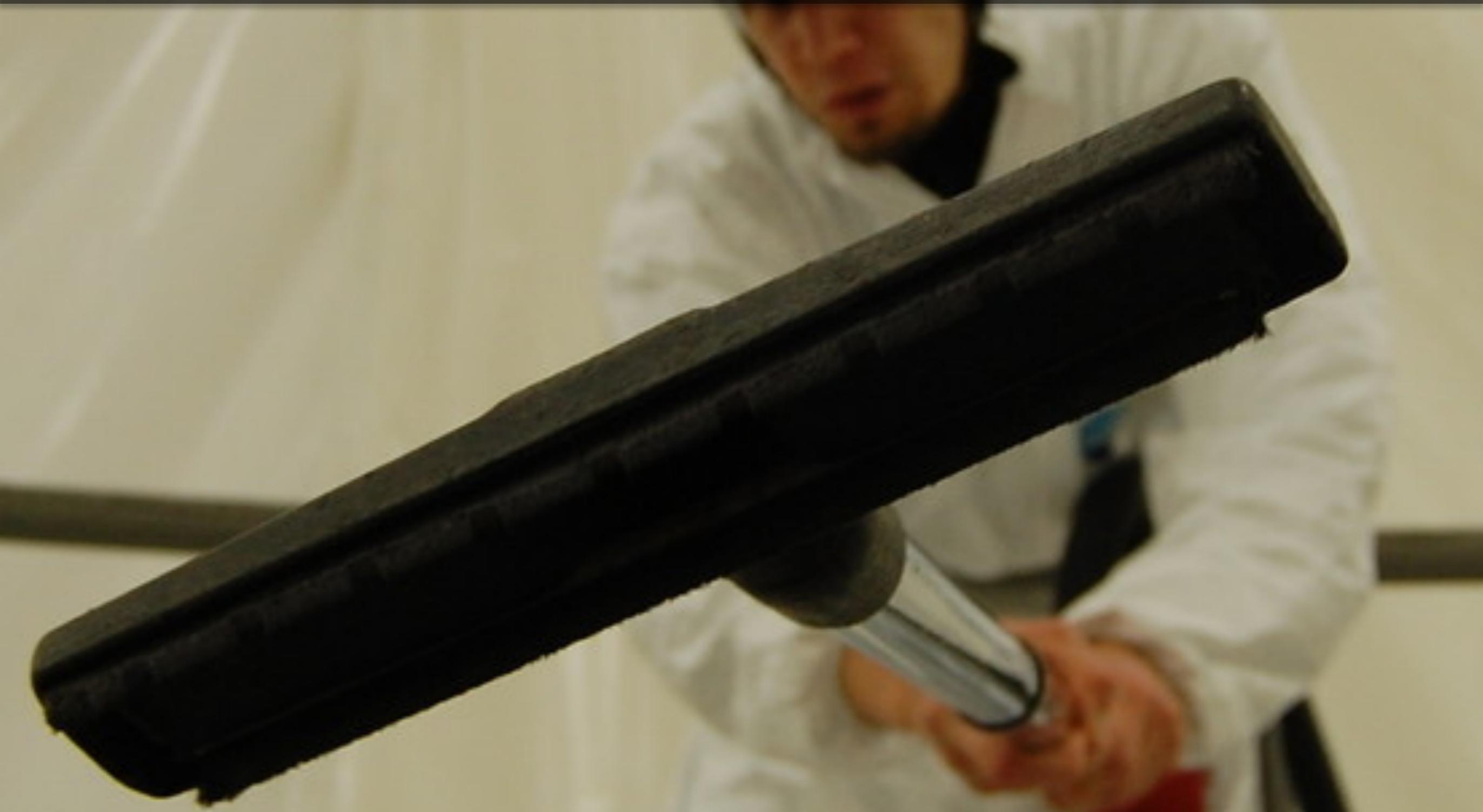
• **other BGs...**

• accidentals

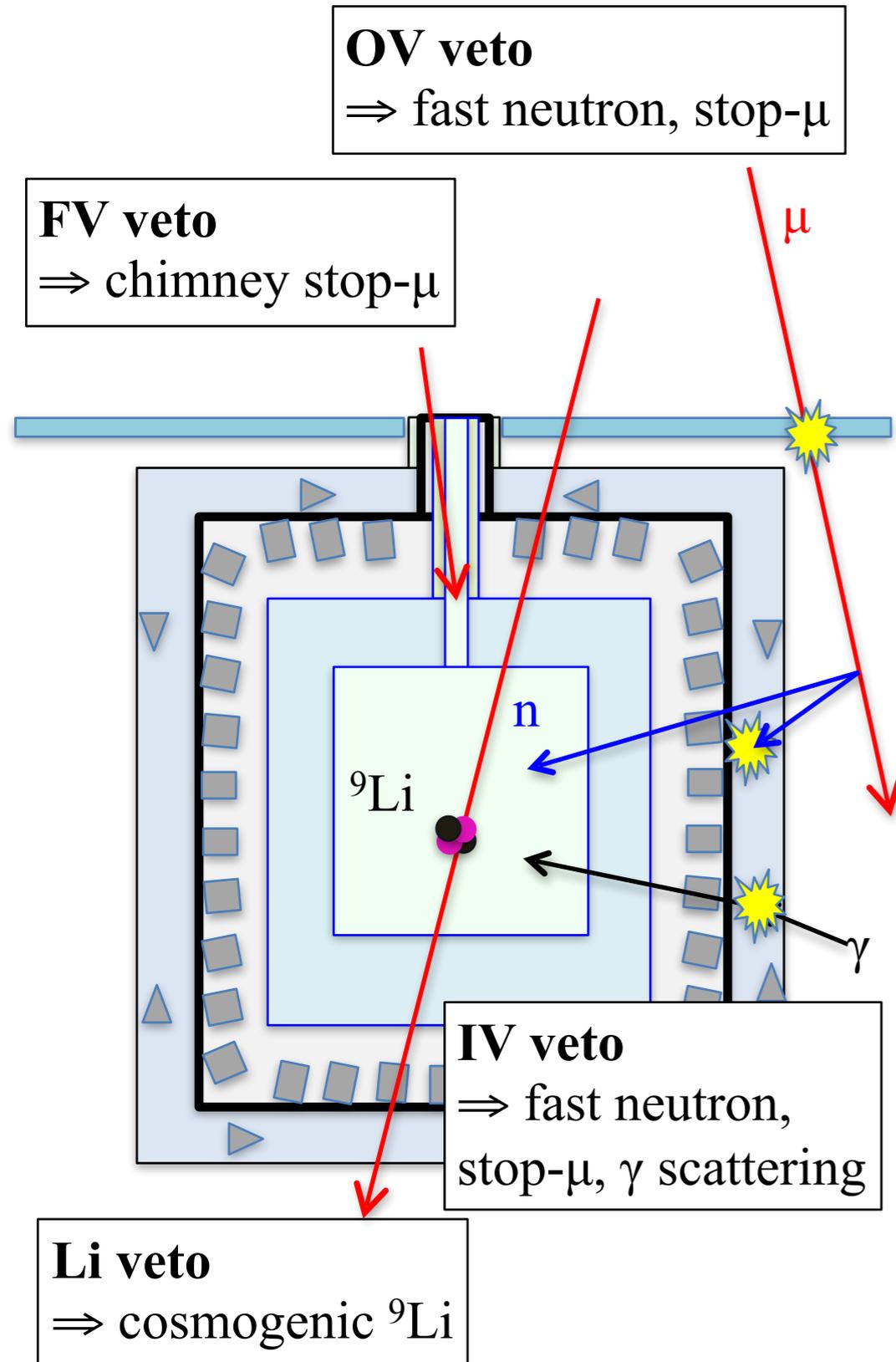
•  $^{13}\text{C}(\alpha, n)^{16}\text{O}$

•  $^{12}\text{B}$  related

active BG rejection...



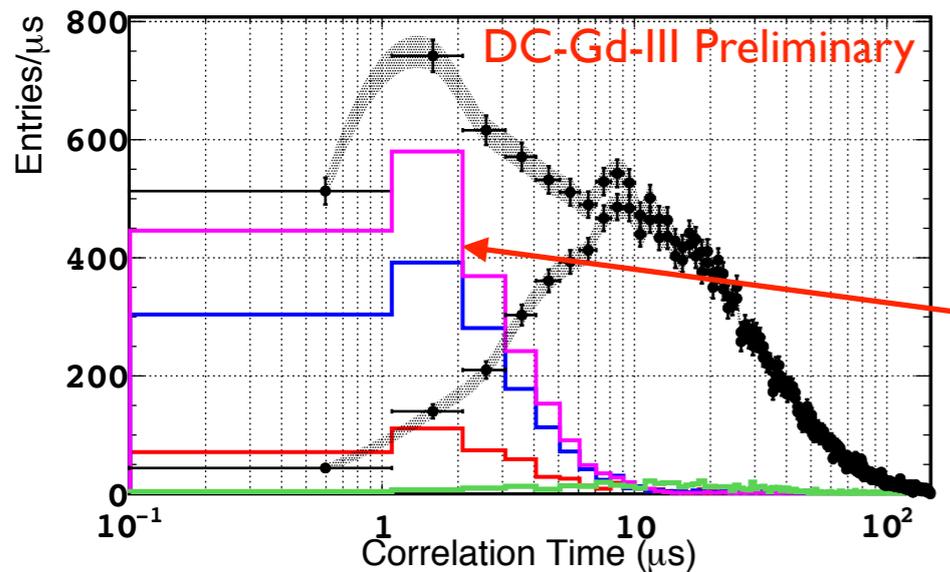
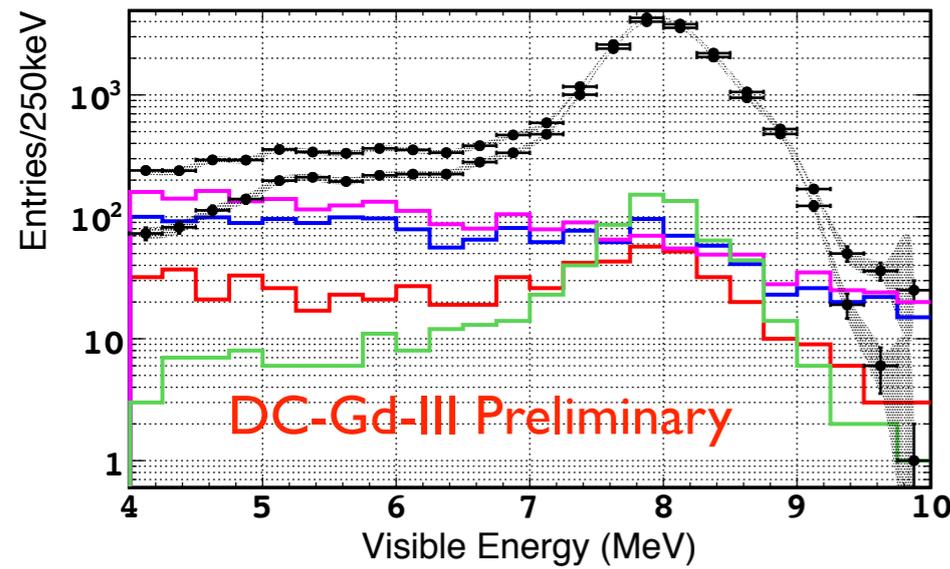
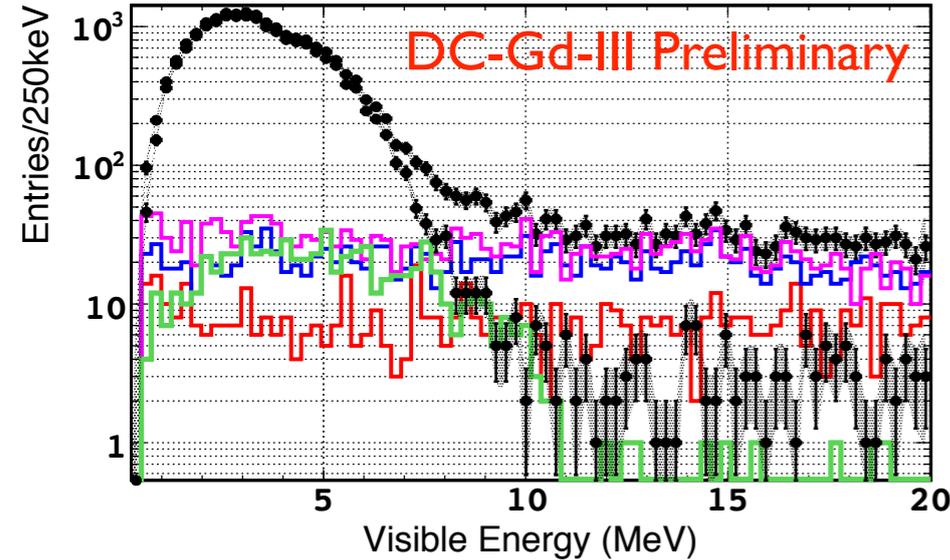
# Background vetoes



Cut	Information used	Target of cut
$\mu$ veto	1ms veto after $\mu$	$\mu$ , cosmogenic
Multiplicity	unicity condition	multiple-n
FV veto	vertex likelihood	chimney stop- $\mu$
IV veto	IV activity	fast n, stop- $\mu$ , $\gamma$ scattering
OV veto	OV activity	fast n, stop- $\mu$
Li veto	Li-likelihood	cosmogenic
LN cut	PMT hit pattern & time	light emission from PMT
(CPS veto)	chimney likelihood	stop- $\mu$
(Qratio)	Max Q/Tot. Q	ND buffer stop- $\mu$

(only applied in multi-detector analysis)

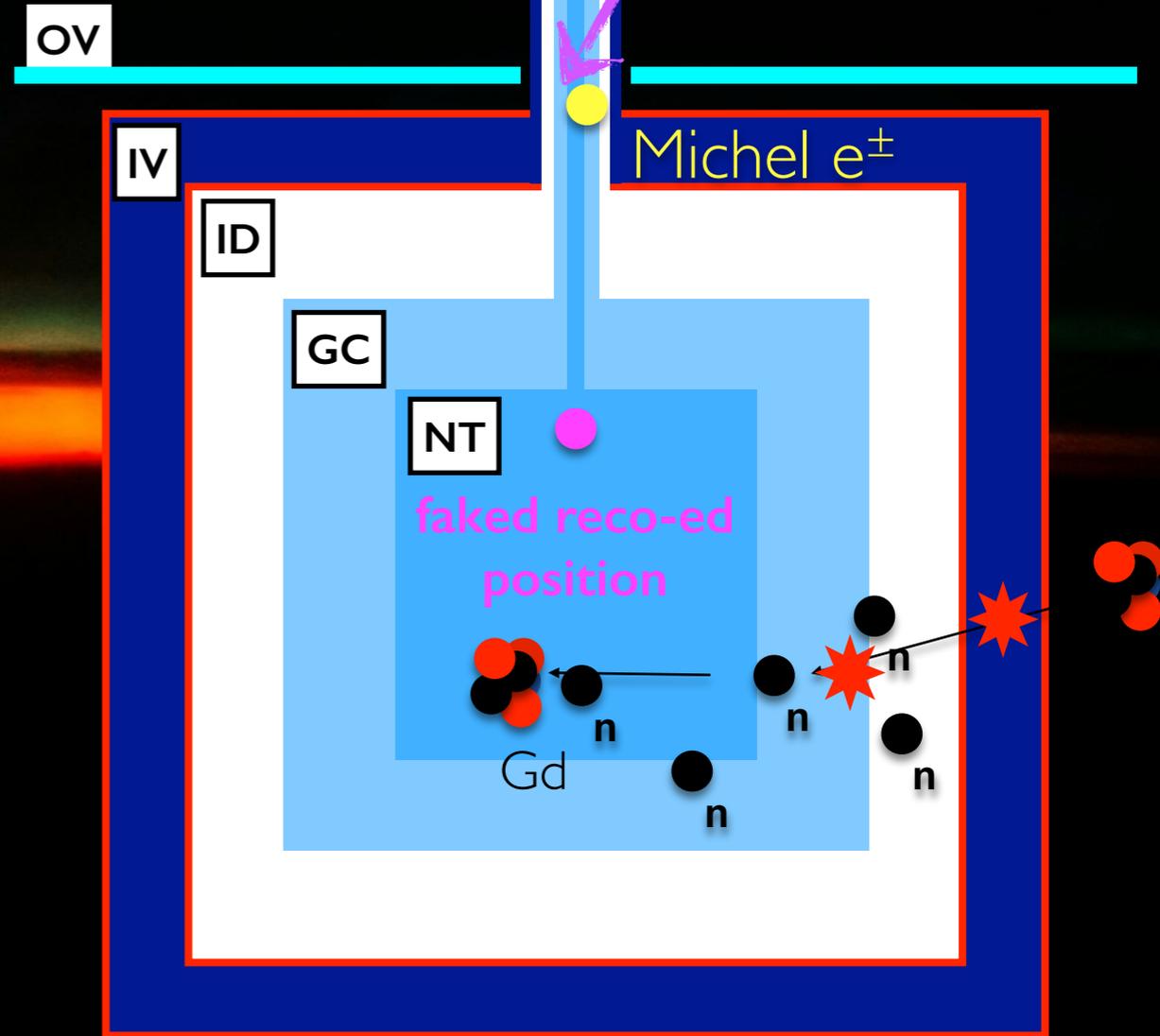
# our vetoes in action (one by one)...



Gd-III (vetoes OFF)  
 FVV veto  
 Li+He veto  
 OV veto  
 IV veto  
 Gd-III (vetoes ON)

## BG rejection...

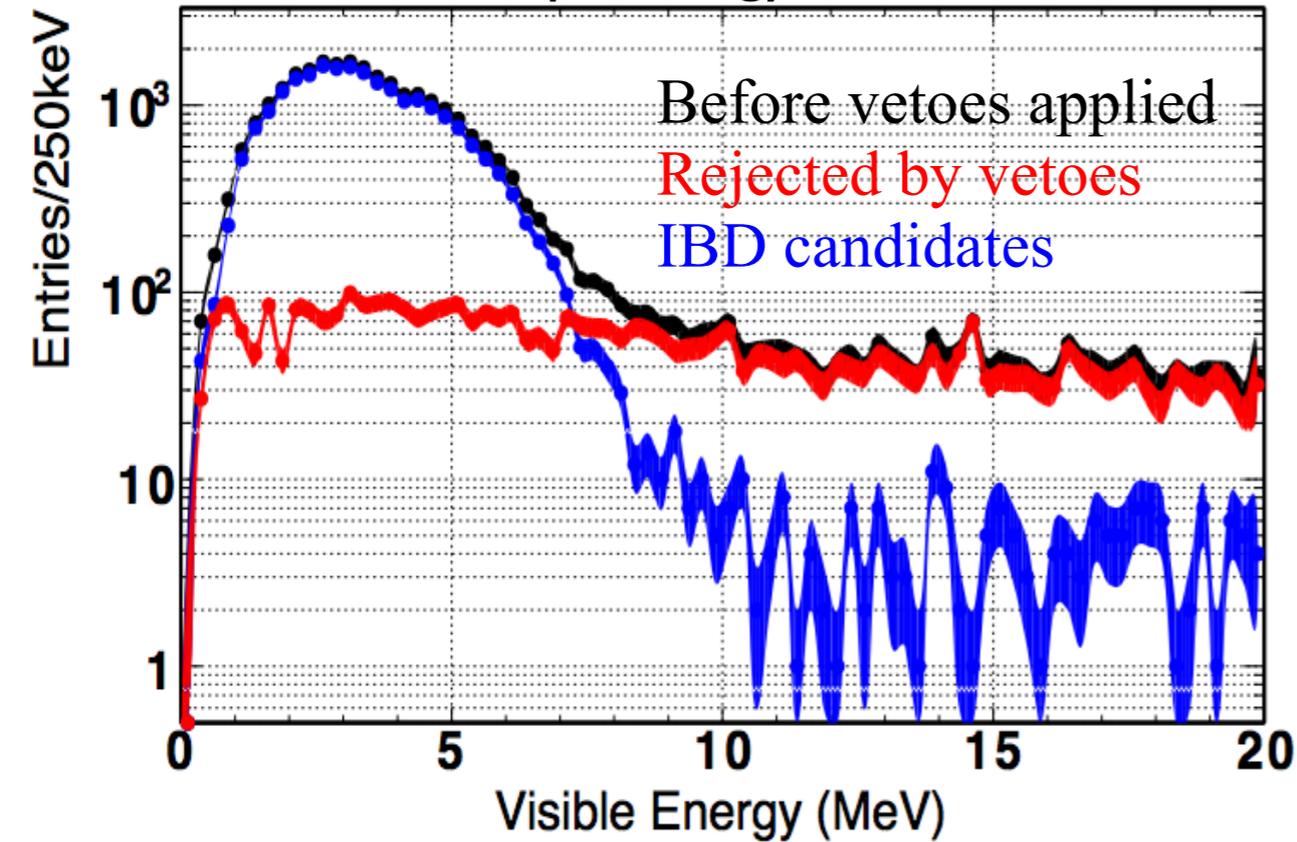
- Li/He → ~50% rejection
- fast-n + stop-μ → ~10x rejection
- accidental → >10x rejection



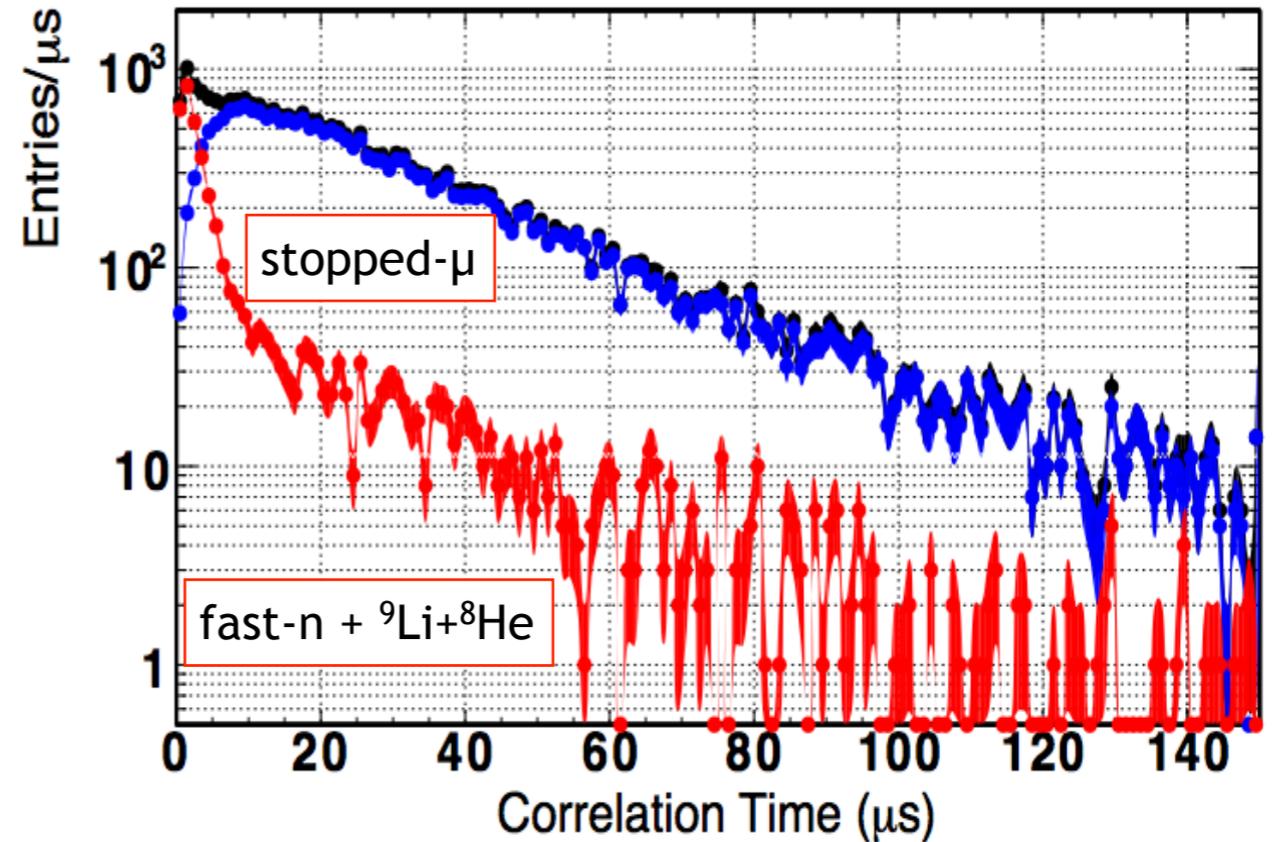
**stopped-μ largest BG**  
 (if not vetoed → ~fully rejected)

# Background vetoes

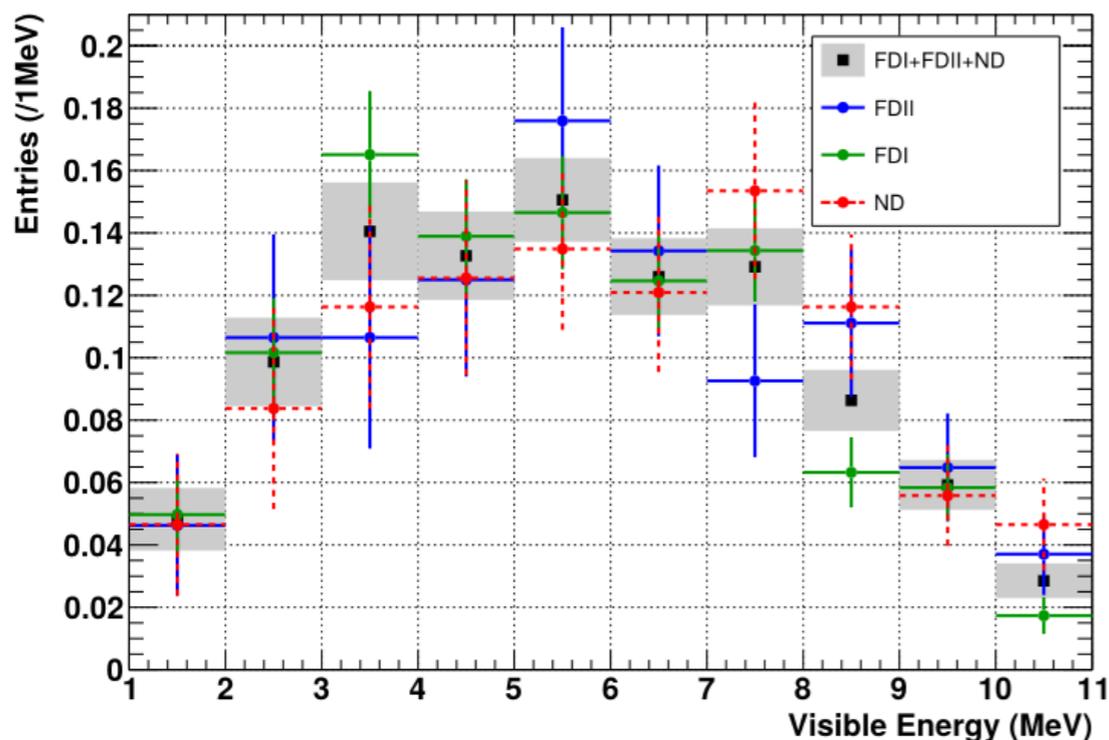
Prompt energy



Correlation time

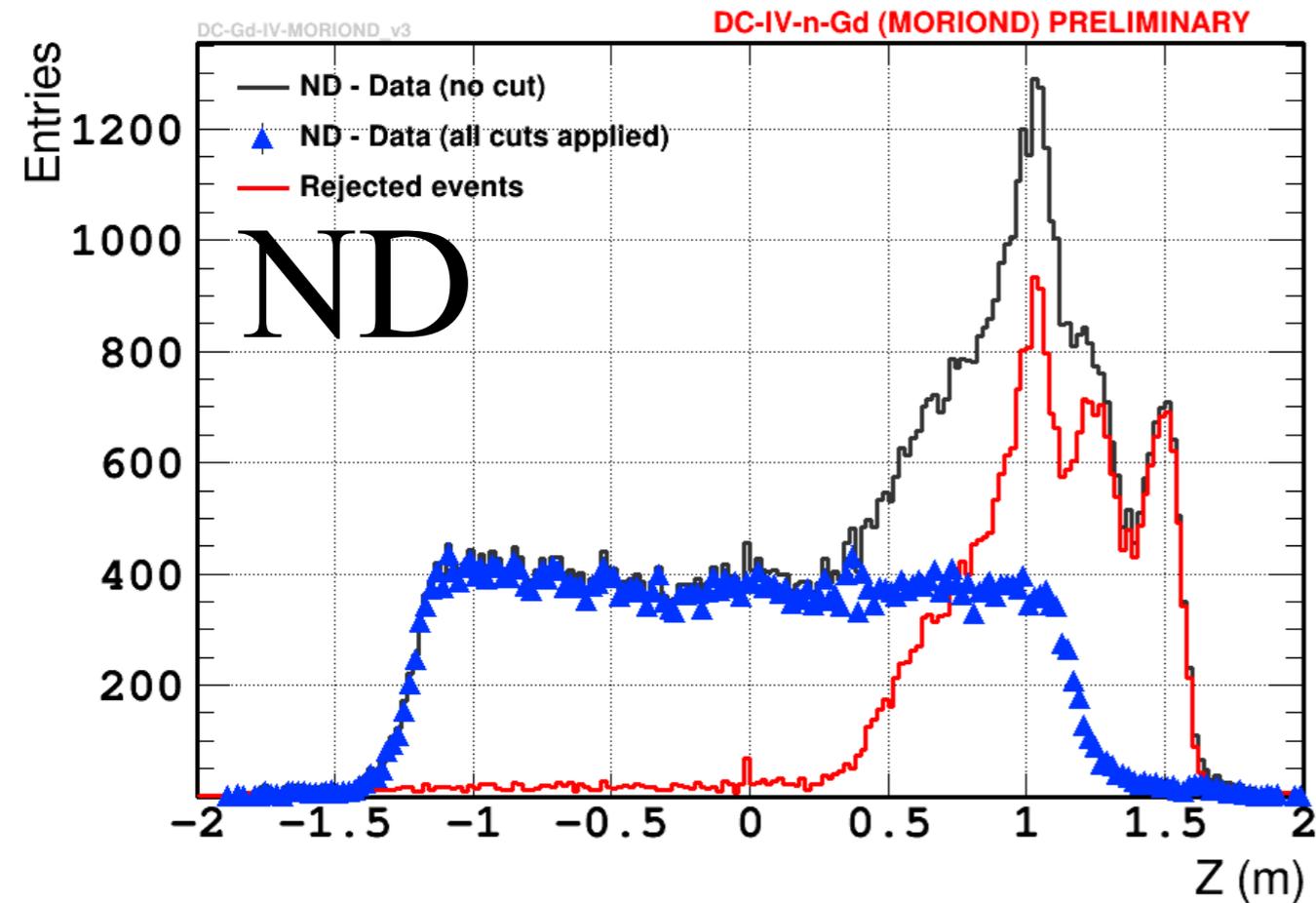


prompt energy (rejected by  ${}^9\text{Li}$  veto)

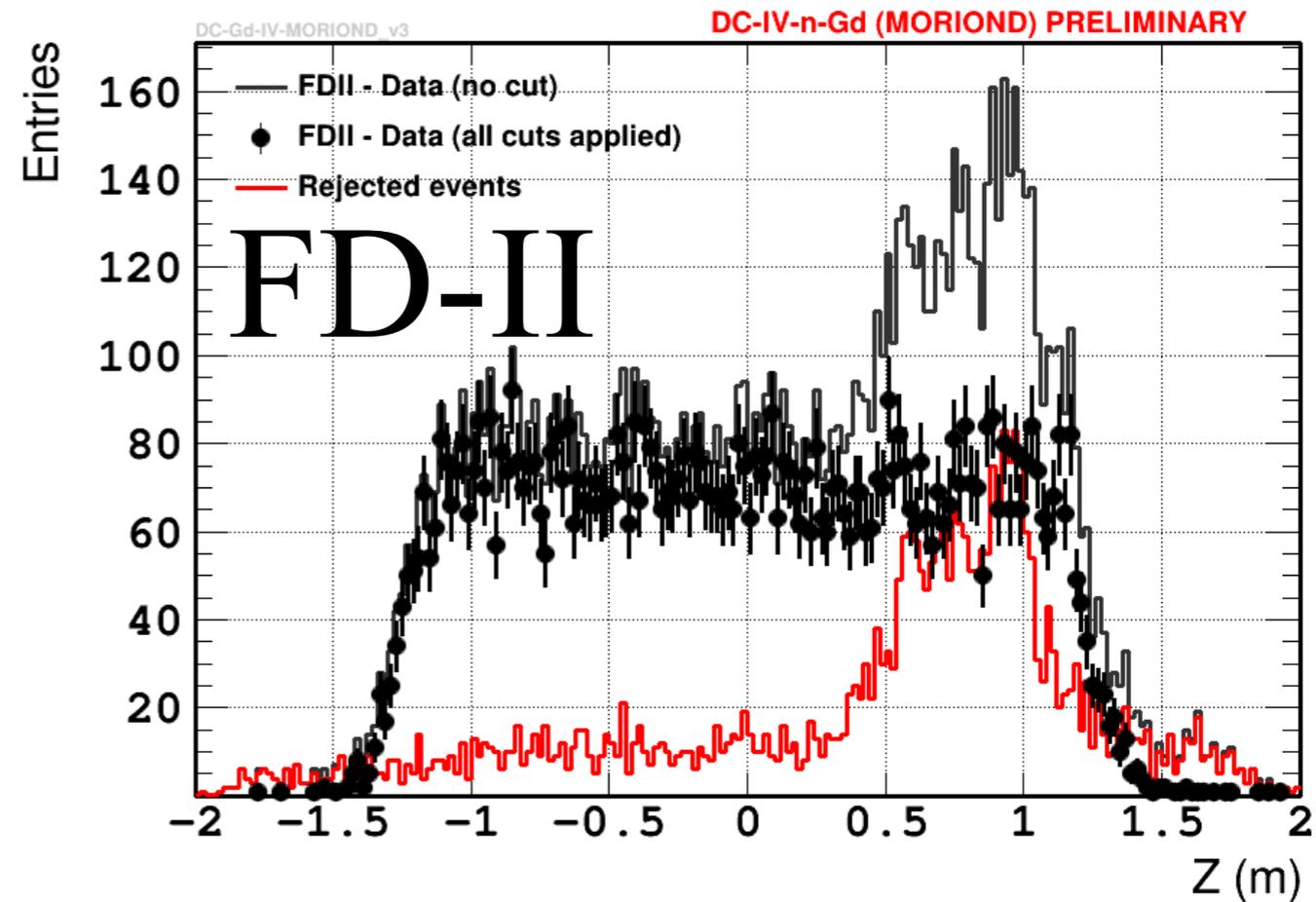


- Significant reduction of background: stop- $\mu$ , fast n,  ${}^9\text{Li}$ , natural radioactivity  
 $\Rightarrow$  Rejected (tagged) events are used to evaluate residual background
- IBD inefficiency:  $< 1\%$  (besides  $\mu$  veto)

# stopped- $\mu$ contamination



inefficiency:  $97.53 \pm 0.02$  (all vetoes)  
rejection power  $> 10x$



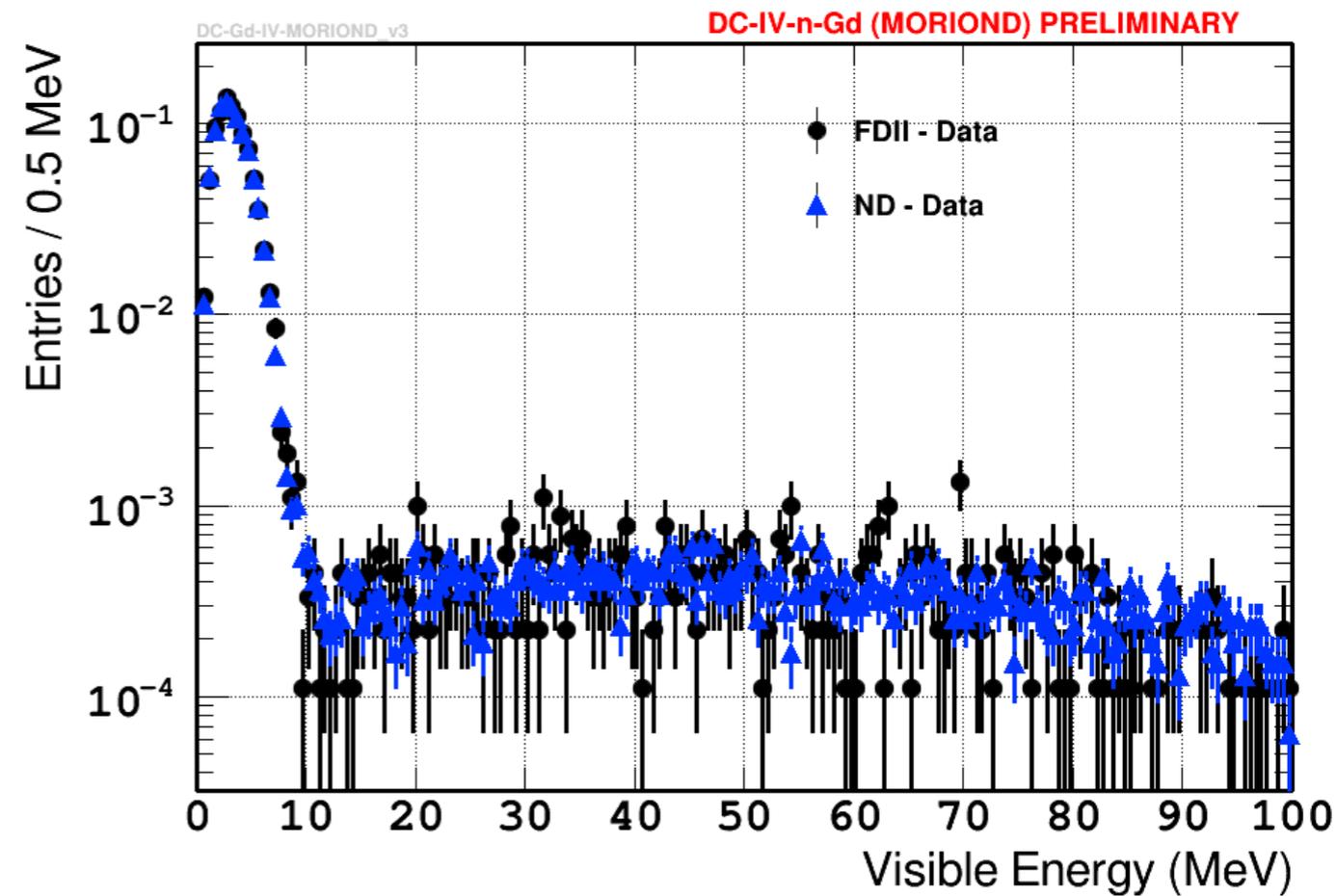
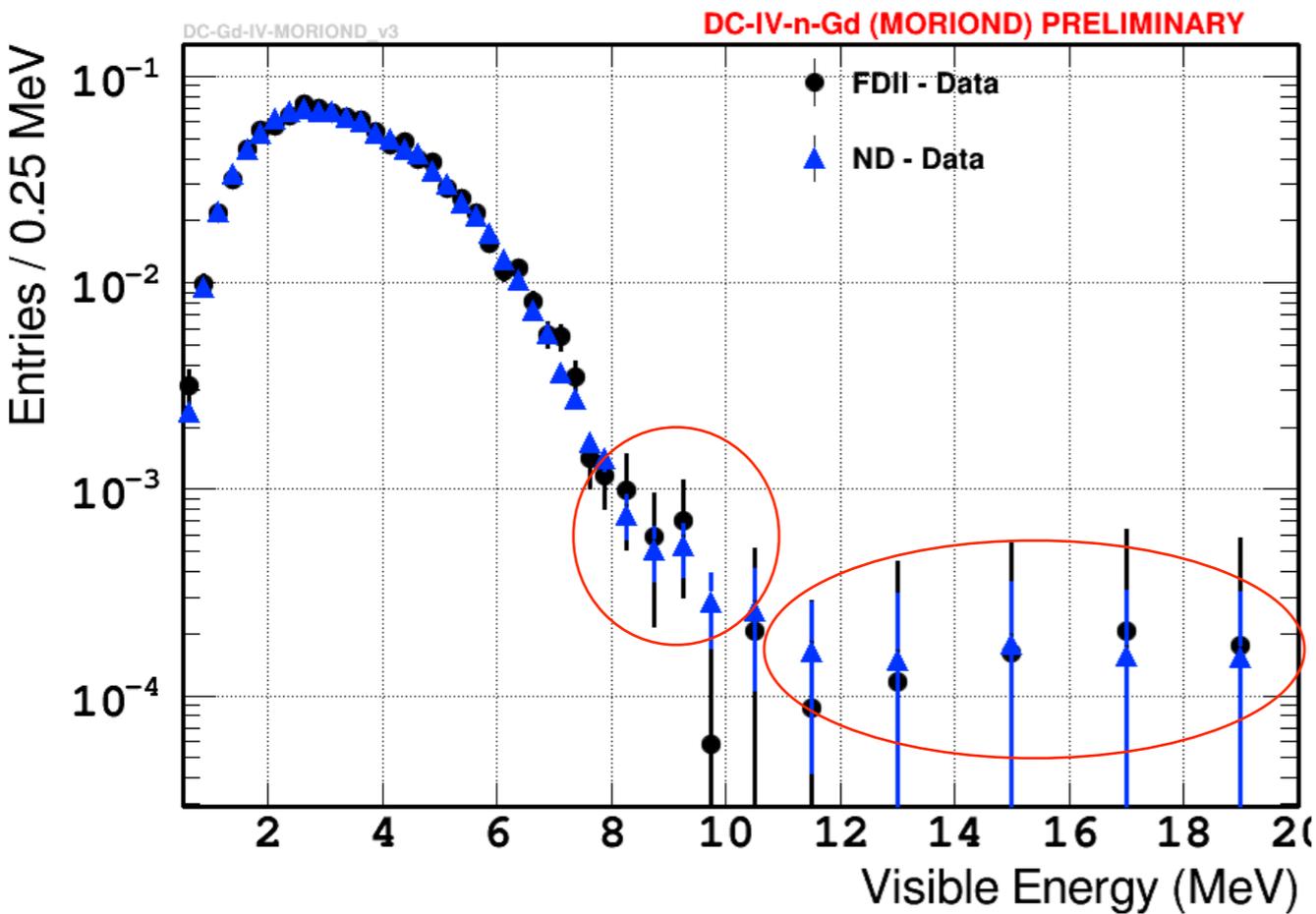
inefficiency:  $97.27 \pm 0.09$  (all vetoes)  
rejection power  $\sim 10x$

(validated with reactor-OFF data)

**vetoes are treated as totally uncorrelated** (same cuts though)

# after cuts: FD vs ND

relatively normalise  $\rightarrow$  hide  $\theta_{13}$  disappearance

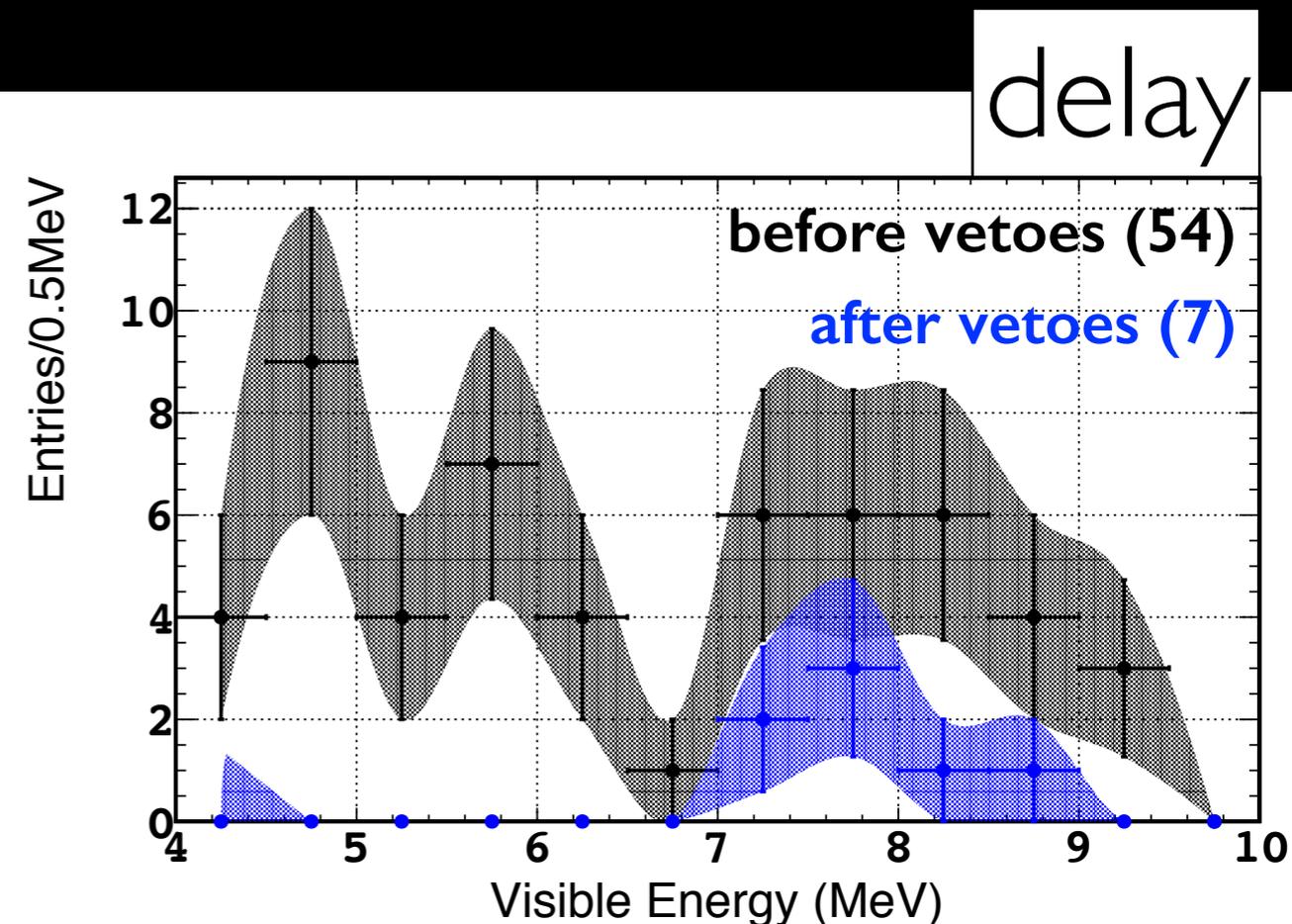
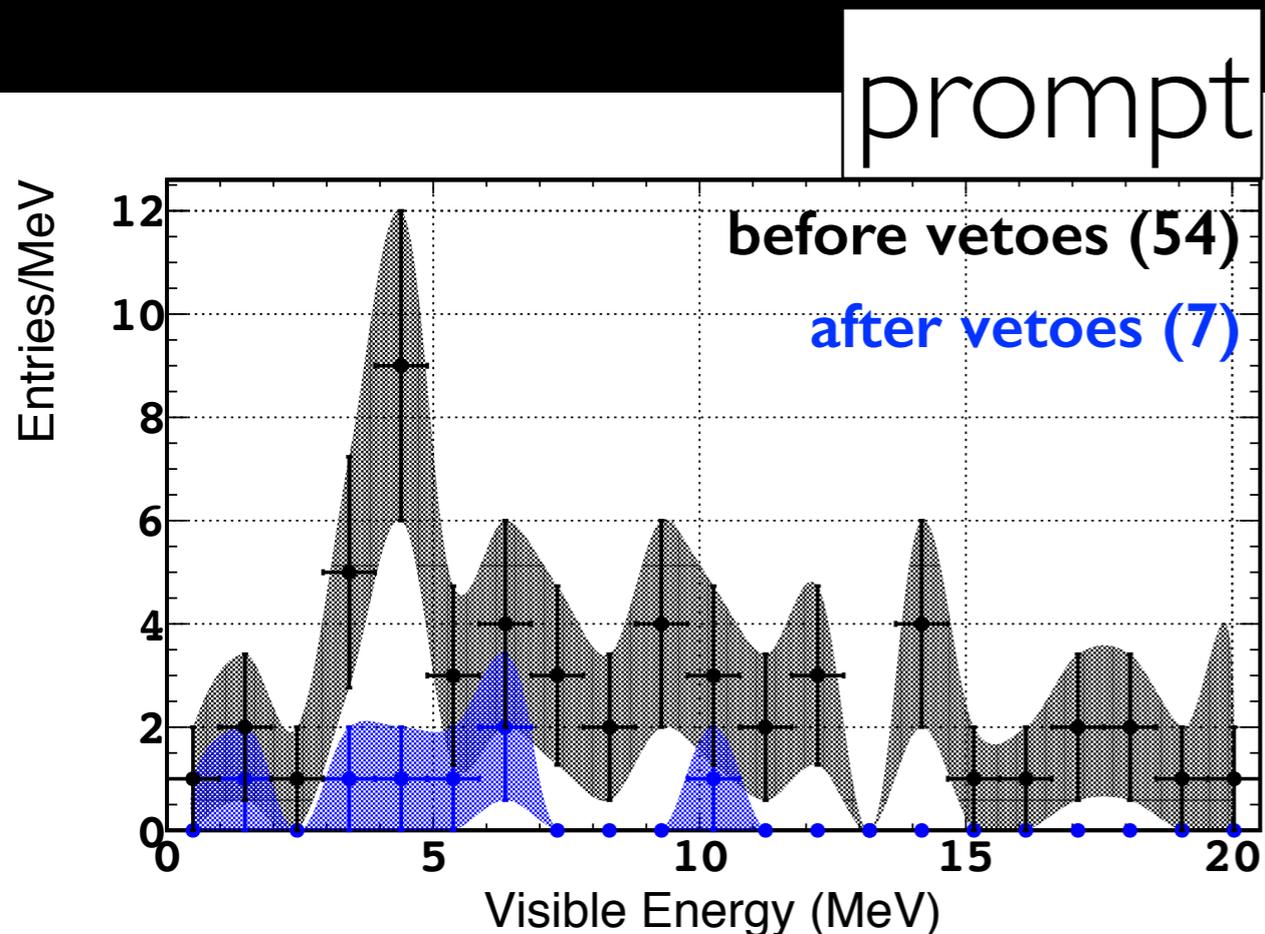


remarkable consistency (signal & BG are different) across FD & ND full dynamics

@ND ~7x more IBDs than @FD

reactor OFF data (only FD)...





**2x reactor-OFF data: powerful information before/after veto evolution**  
(scrutinising a few event-wise BG-only)

1 week → **poor stats** (spectral info fluctuations dominated) → inconclusive

**$P(\text{rejection}) = (7.7 \pm 3.1) @ \text{Gd-III}$**

(in agreement with  $(9.9 \pm 1.0)$  estimated between  $[12, 20]$  MeV)

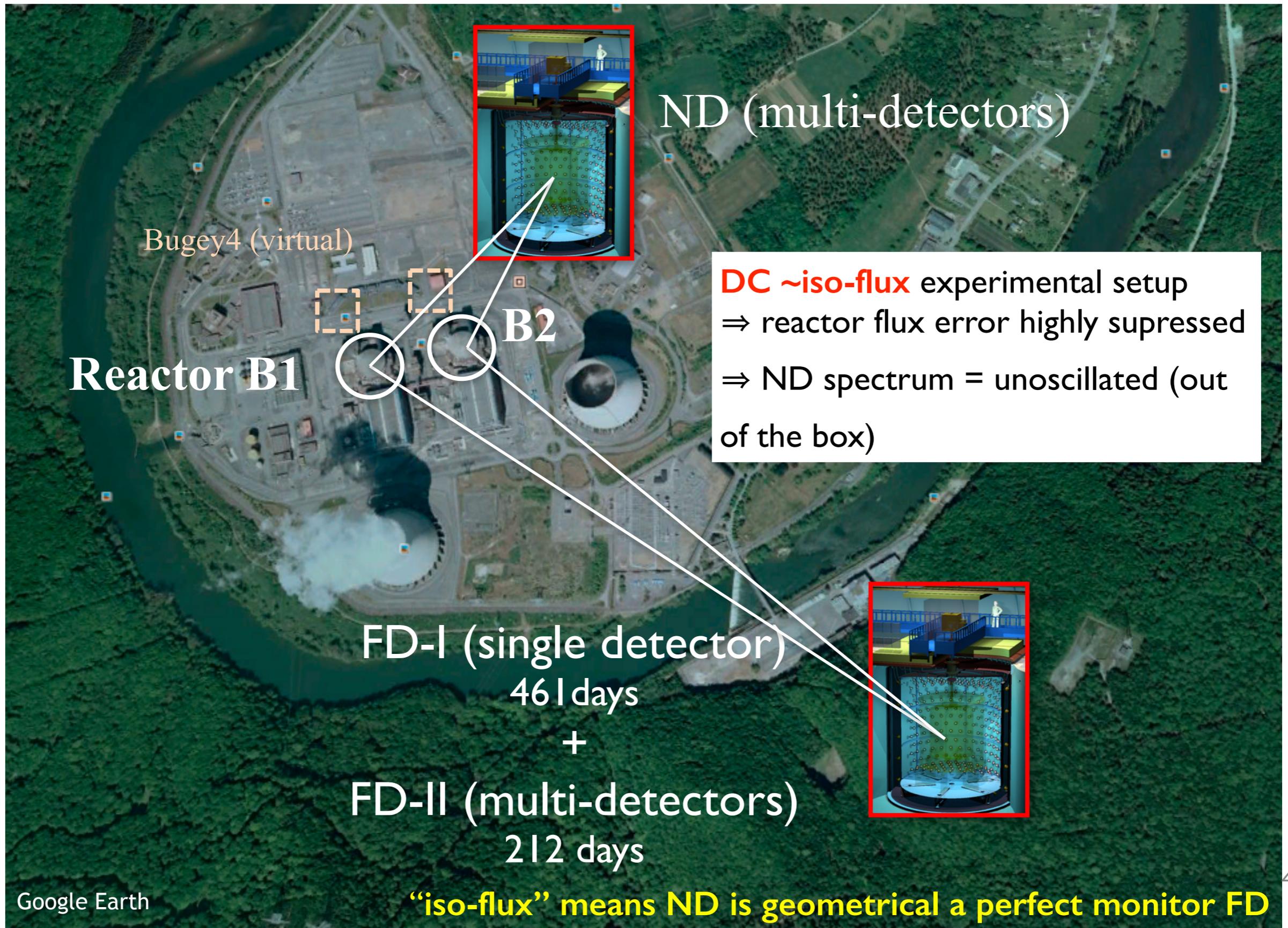
# Signal and background

Double Chooz Preliminary

	FD-I	reactor-off	FD-II	ND
Live-time (d) (after $\mu$ veto)	460.93	7.24	212.21	150.76
IBD prediction ( $d^{-1}$ )	$38.04 \pm 0.67$	$0.217 \pm 0.065$	$40.39 \pm 0.69$	$280.5 \pm 4.7$
Accidental BG ( $d^{-1}$ )	$0.070 \pm 0.003$		$0.106 \pm 0.002$	$0.344 \pm 0.002$
Fast-n + stop- $\mu$ ( $d^{-1}$ )	$0.586 \pm 0.061$			$3.42 \pm 0.23$
Cosmogenic ( $d^{-1}$ )	$0.97^{+0.41}_{-0.16} *$			$(5.01 \pm 1.43) *$
<b>Total prediction (<math>d^{-1}</math>)</b>	<b><math>39.63 \pm 0.73</math></b>	<b><math>1.85 \pm 0.30</math></b>	<b><math>42.06 \pm 0.75</math></b>	<b><math>289.3 \pm 4.9</math></b>
<b>IBD candidates (<math>d^{-1}</math>)</b> (number of events)	<b>37.64</b> (17351)	<b>0.97</b> (7)	<b>40.29</b> (8551)	<b>293.4</b> (44233)

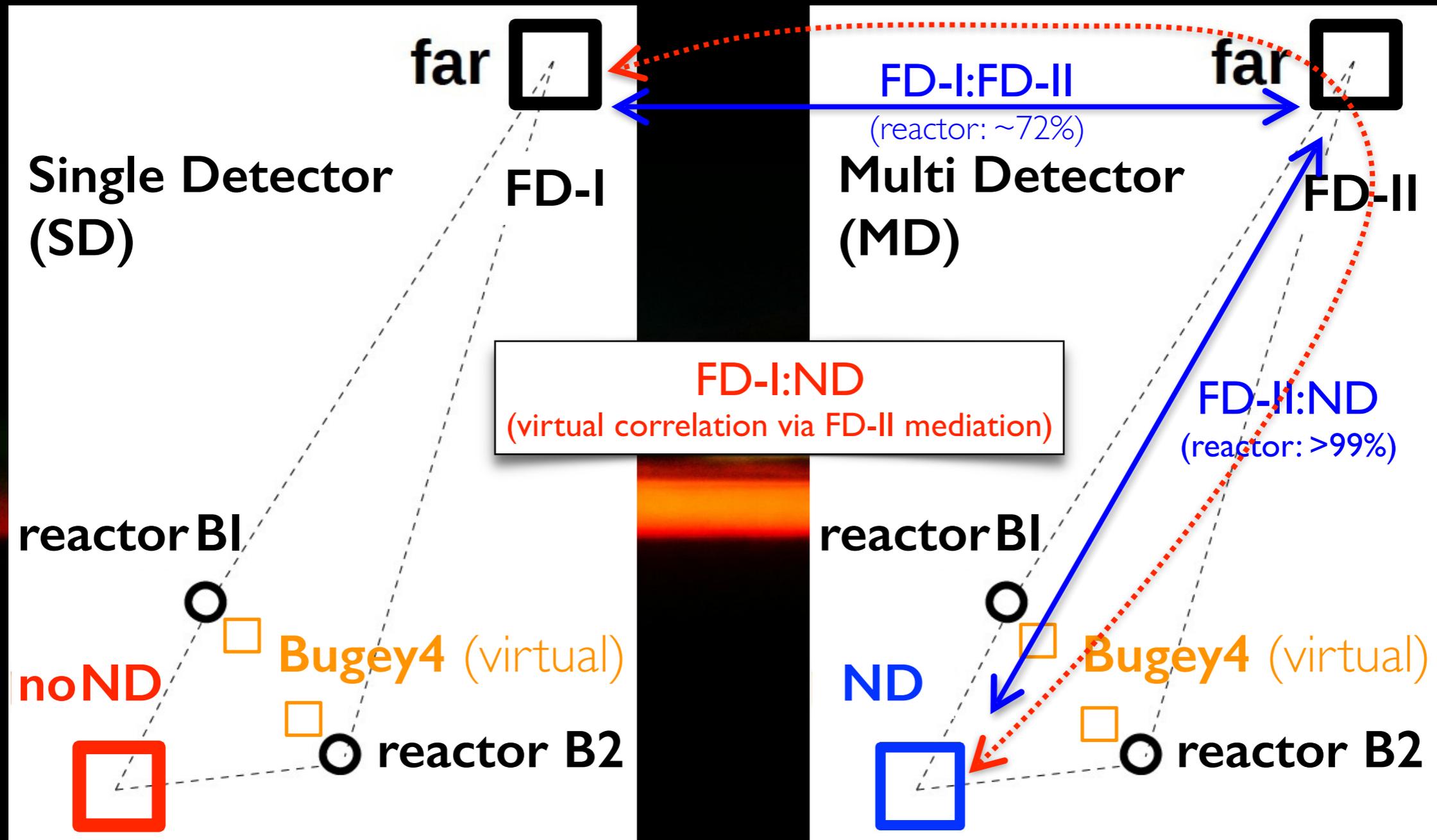
- **all BG shapes measured with data** (no MC used)
- cosmogenic BG ( ${}^9\text{Li} + {}^8\text{He}$ ): rate not used as input to rate+shape fit (independent)  
 $\Rightarrow$  rates are constrained directly in fit via rate+shape information (data spectrum)

# Multi-detectors analysis (today)



(exposure used: ~48months)

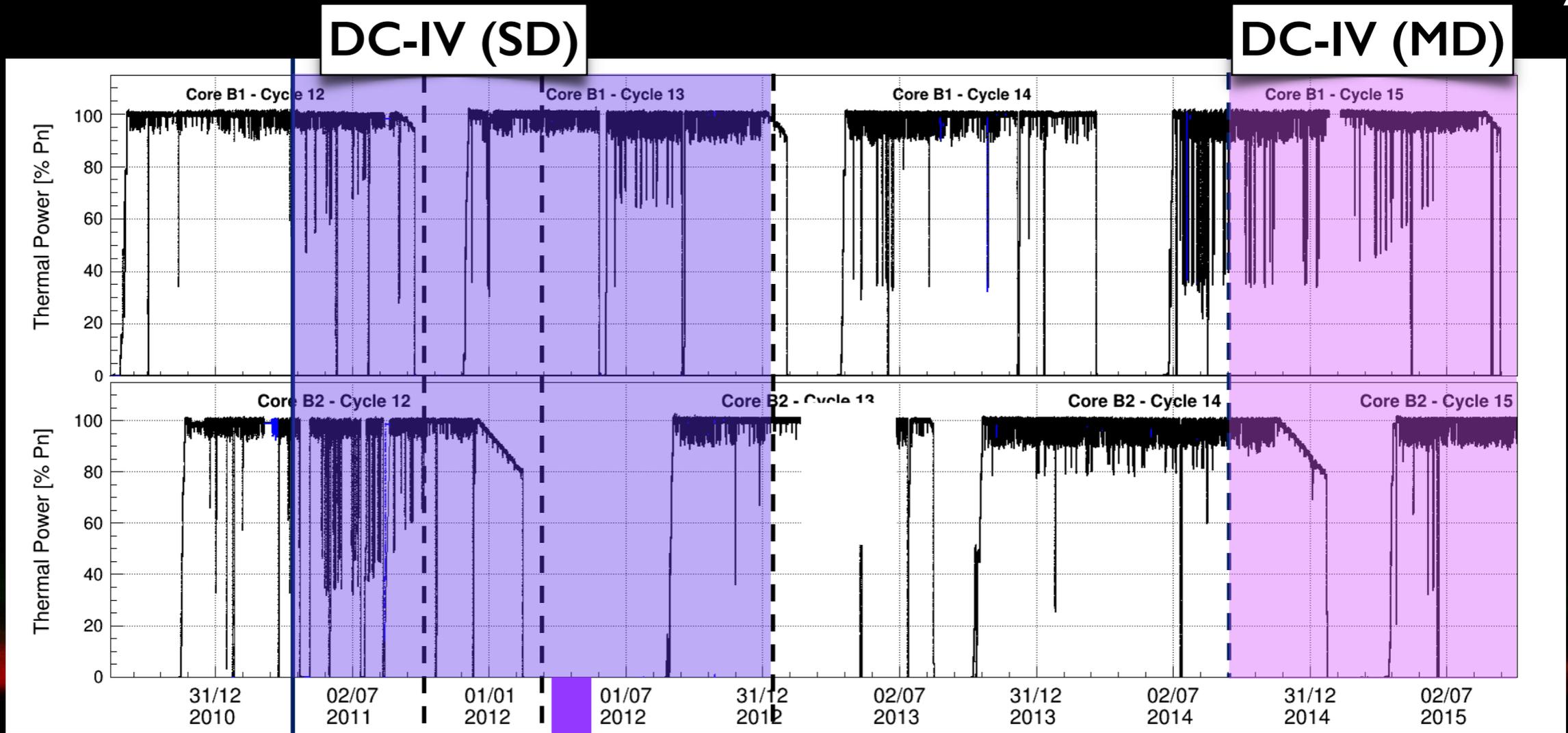
(exposure used: ~9months @ Moriond)



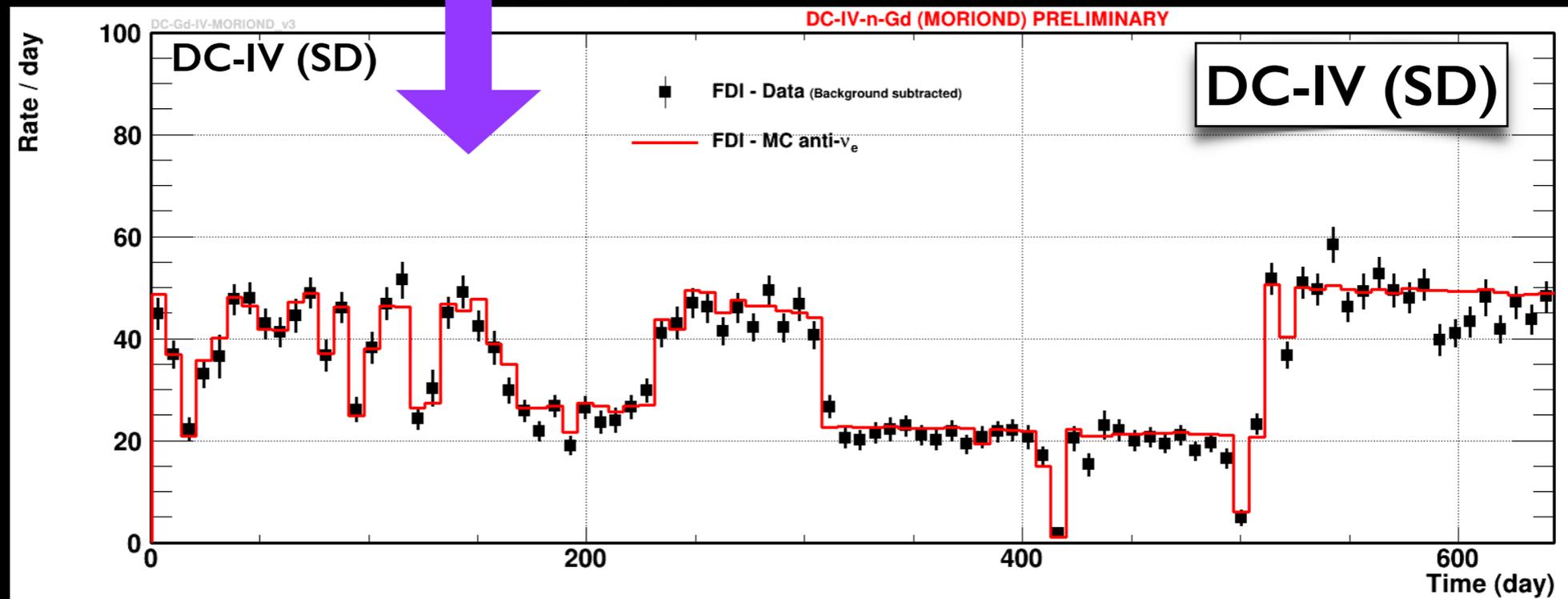
- FD-I  $\approx$  FD-II (same)  $\rightarrow$  strongly correlated BG & response
  - precious extra Li+He BG & **Reactor-OFF** info
  - FD-II hardware upgraded (2x more gain & all PMTs ON)
- **indirect ND monitored:** new error  $\sim 0.9\%$  (1.7% before)

- FD-II **directly ND monitored**
- correlation FD-II with FD-I and with ND
  - unknown inter-reactor error correlation
  - **use value maximising  $\sigma[\sin^2(2\theta_{13})]$**

reactor power information



measured IBD's



# Reactor flux uncertainties

Double Chooz Preliminary

without Bugey4  
total error ~3.0%

	FD-I (%)	FD-II (%)	ND (%)	
Bugey4	1.40	1.40	1.40	} <b>Correlated</b> across FD-I, FD-II and ND
Energy per fission	0.16	0.16	0.16	
Spectrum $\oplus \sigma_{IBD}$	0.20	0.20	0.20	
Baselines	< 0.01	< 0.01	0.01	} <b>Uncorrelated</b> $\Rightarrow$ suppressed with two detectors (in parallel operation)
Fission fraction ( $\alpha_k$ )	0.82	0.74	0.73	
Thermal power ( $P_{th}$ )	0.44	0.44	0.44	
<b>Total</b>	<b>1.70</b>	<b>1.66</b>	<b>1.66</b>	
$\rho(\text{FD-I:FD-II})$ (error)	<b>0.72 (0.90% relative)</b>		—	
$\rho(\text{FD-II:ND})$ (error)	—	<b>&gt;0.99 (0.07% relative)</b>		

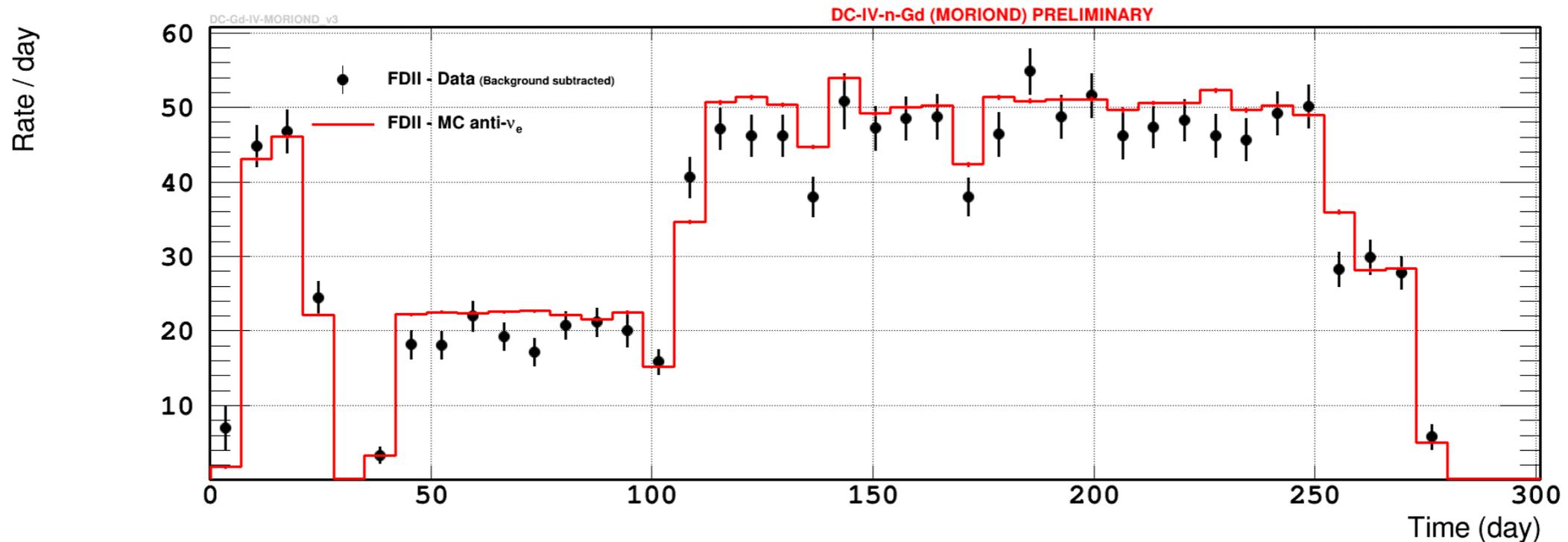
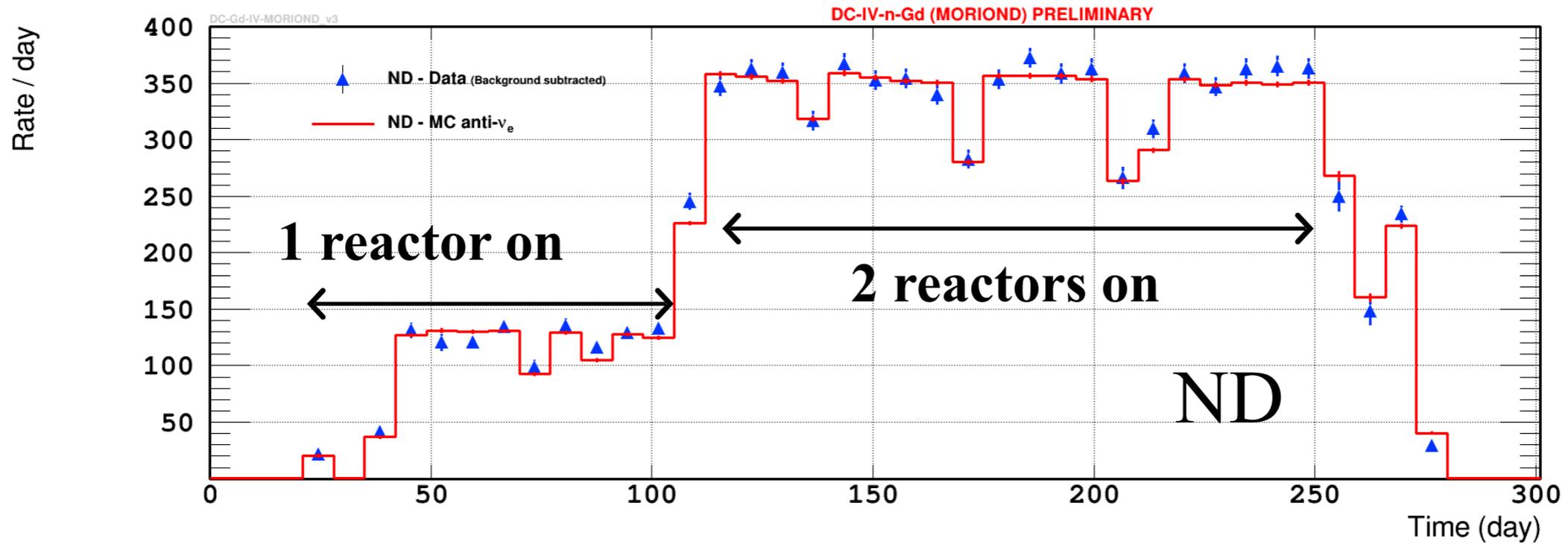
inter-reactor correlation for  $\alpha_k$  and  $P_{th}$ :  $\rho_{B1/B2} = 0.78$   
(most conservative assumption with current data set)

- reactor flux uncertainty suppressed to **< 0.1%** in multi-detector analysis thanks to nearly **iso-flux** experimental setup in DC

# IBD rate vs. time

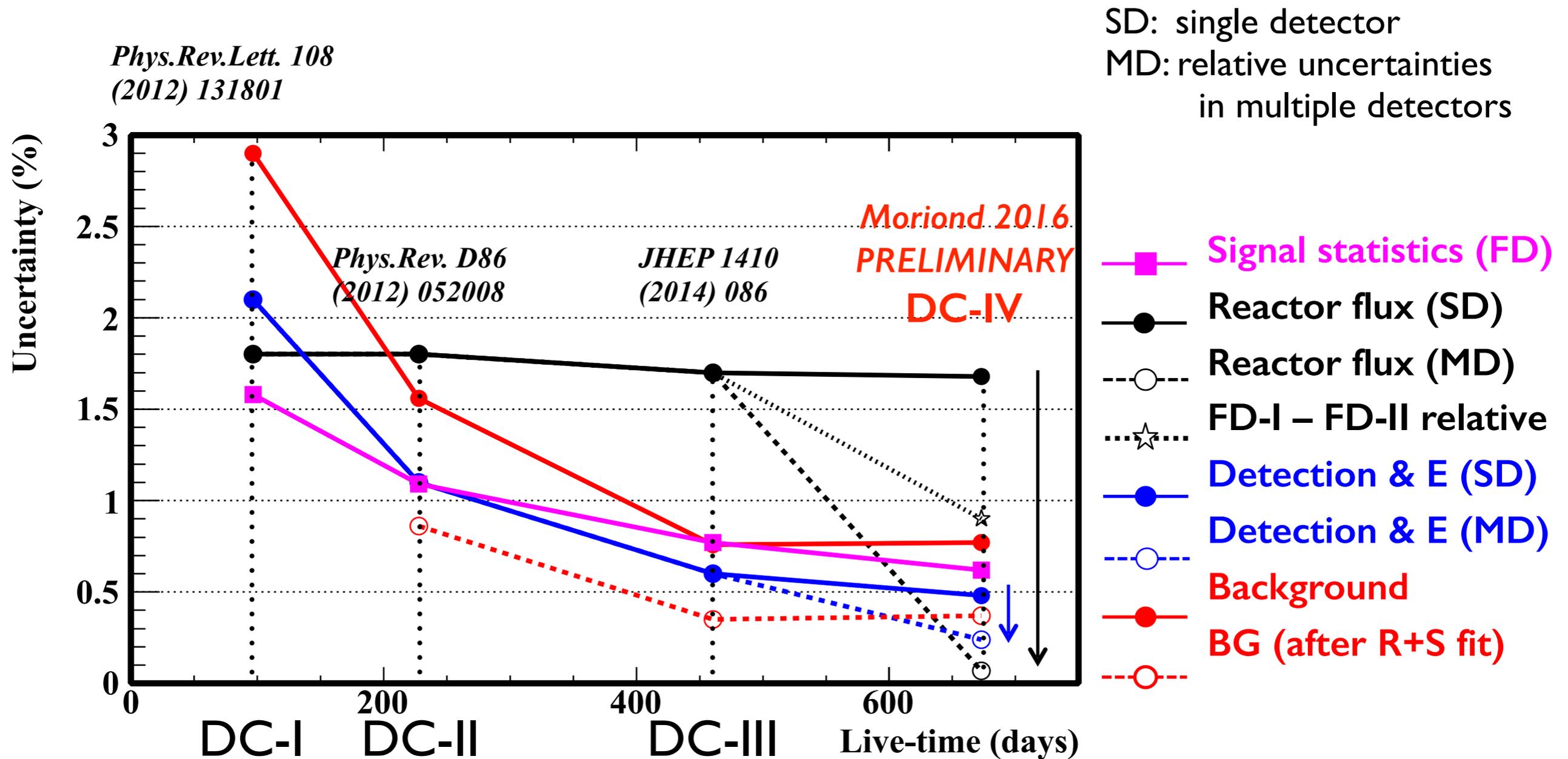
Background subtracted

Double Chooz Preliminary



ND is ~perfect monitor of FD (rate & shape): ~iso-flux configuration

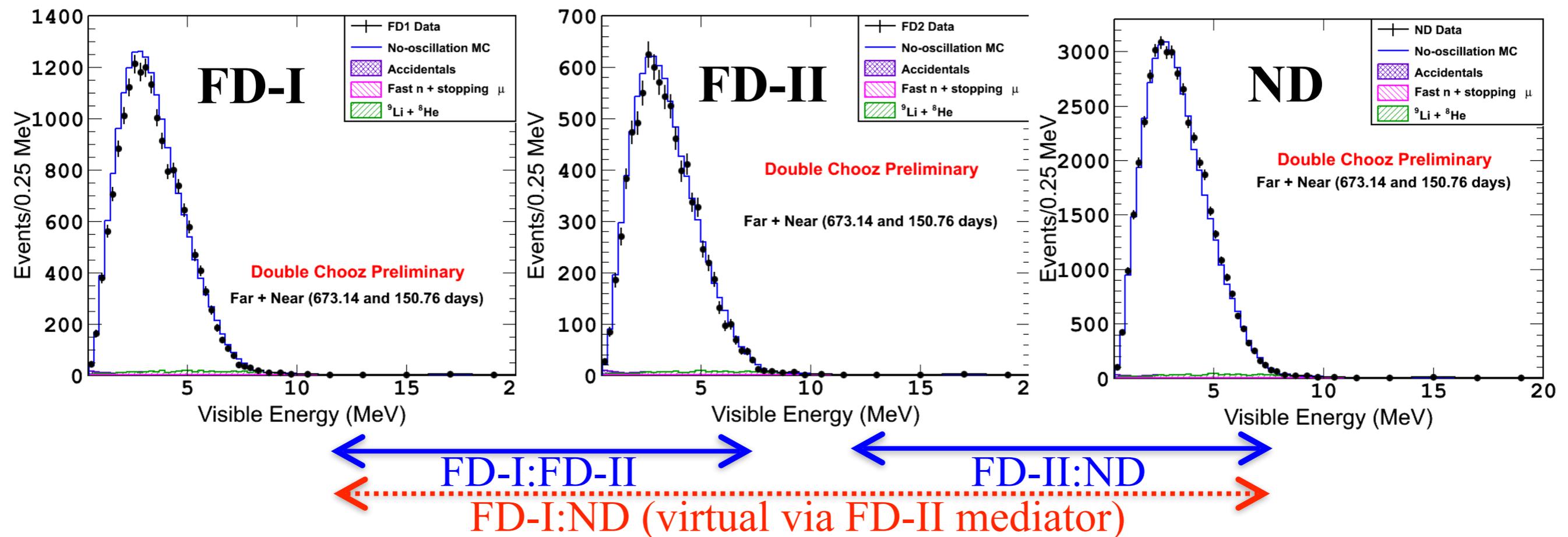
# uncertainties in multi-detectors analysis



- systematic errors suppressed with two detectors and in rate+shape fit  
 $\Rightarrow$  All systematic uncertainties below  $< 0.4\%$  (after R+S fit)
- current precision (9 months ND) is limited by the statistical uncertainty

# Extraction of $\theta_{13}$ by oscillation fit

- Compare FD-I, FD-II and ND data simultaneously to predictions
  - Background rate and shape estimated by data (Li rate not constrained)
  - Observed data in reactor off as separate term  $\Rightarrow$  BG constraint



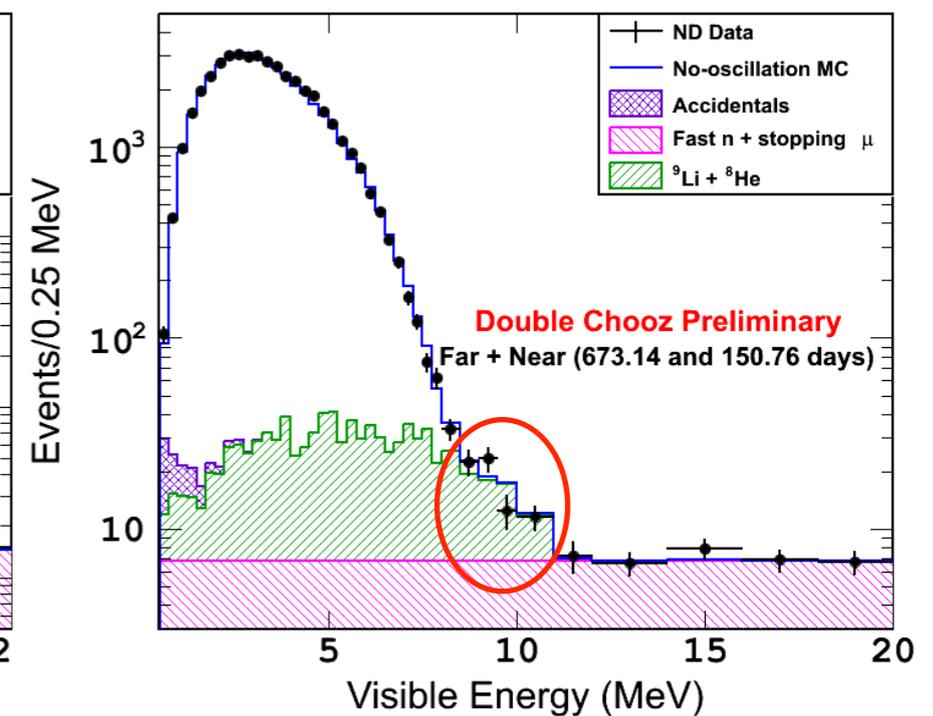
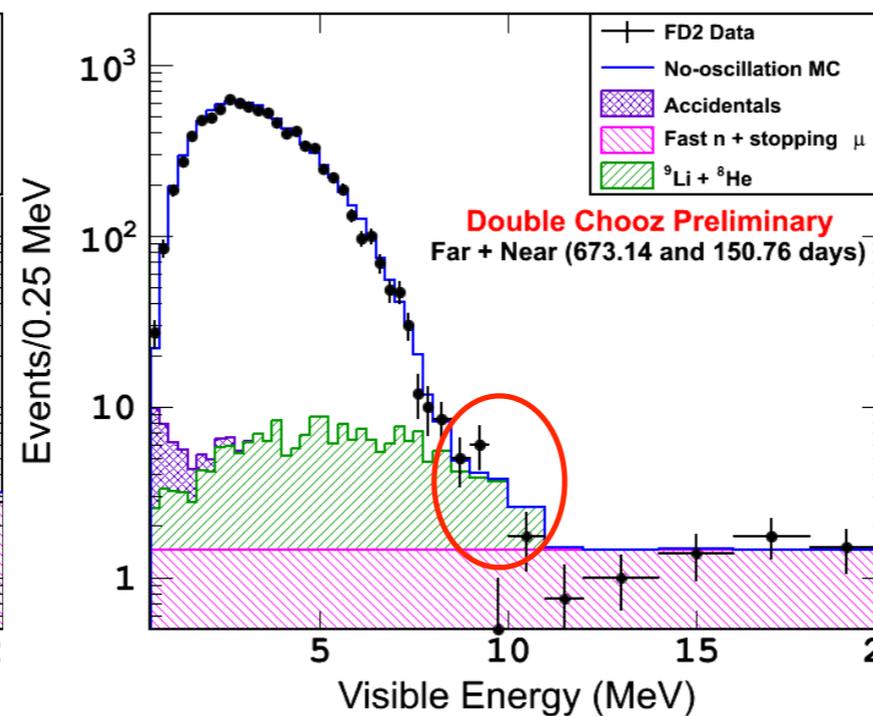
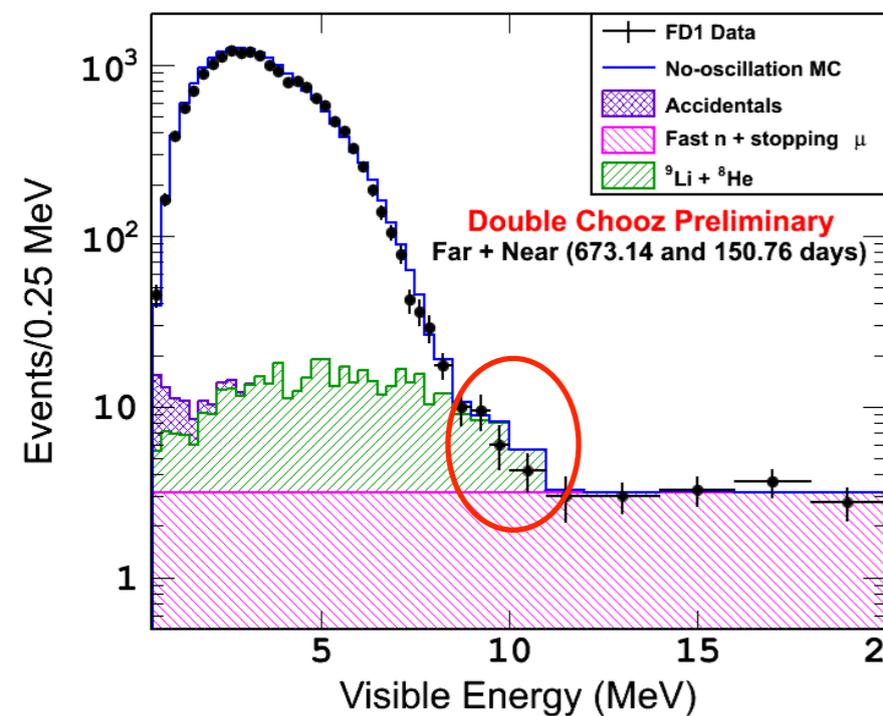
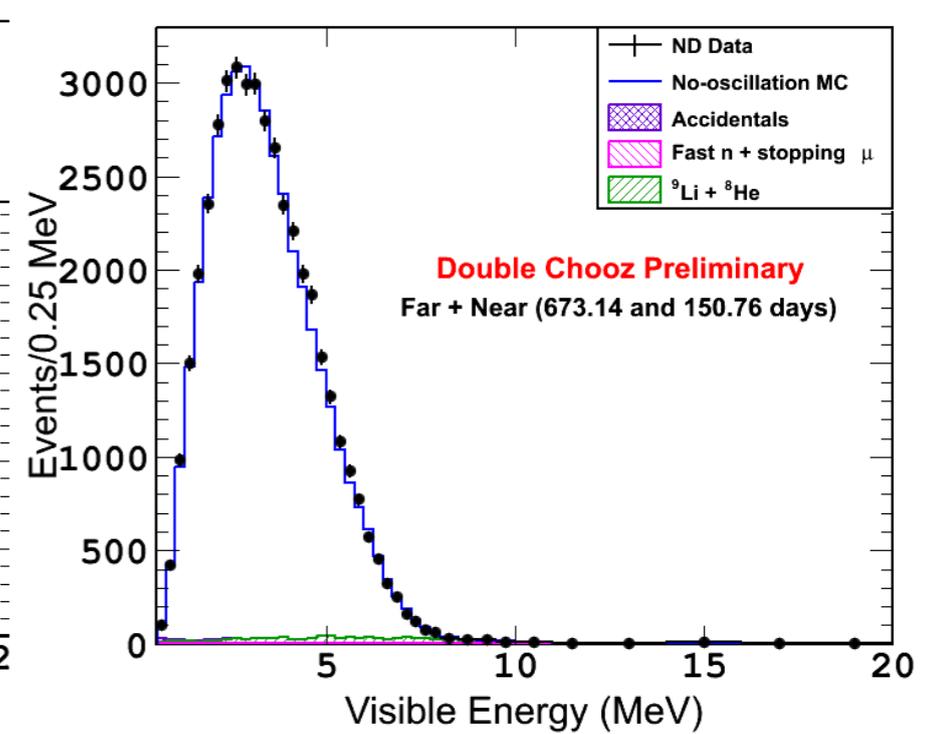
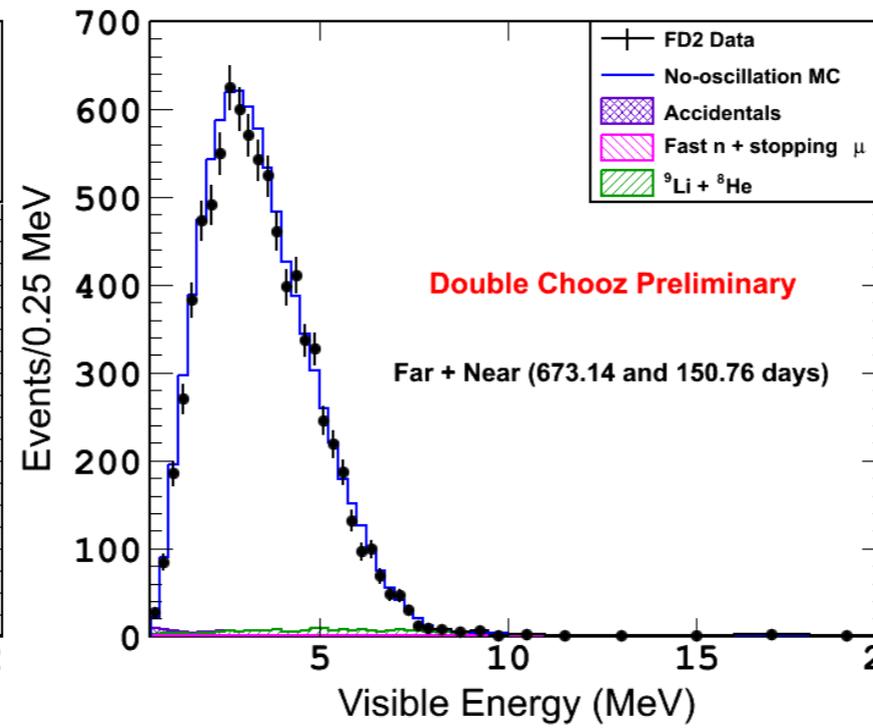
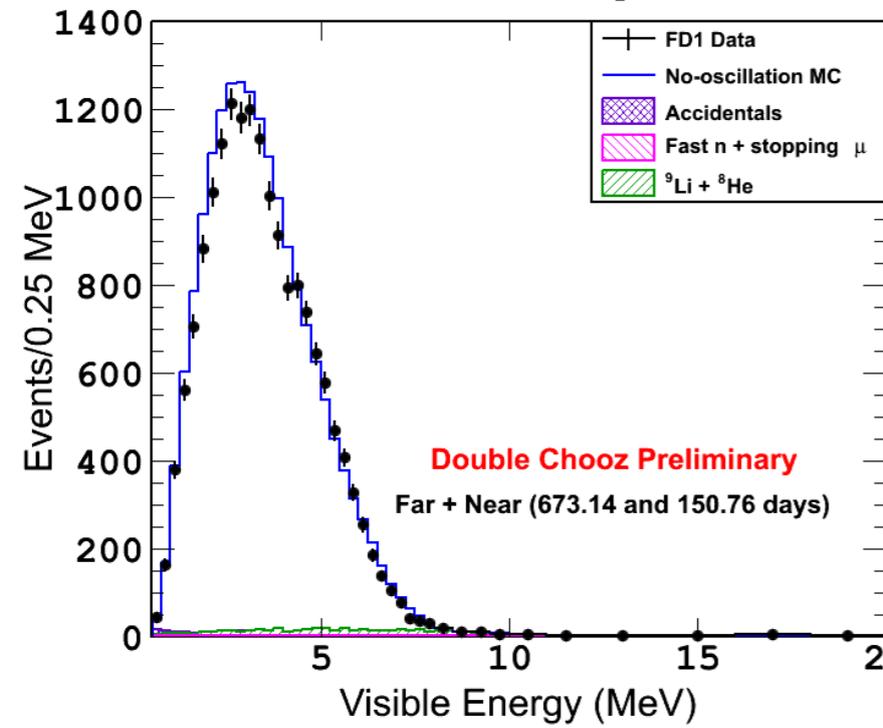
- correlation of systematic uncertainties are taken into account
- energy: **non-linearity uncorrected across all detectors** ( $\rightarrow$ conservative)
- many cross-checks (data-only, BG constraint etc) & independent  $\chi^2$  and likelihood fits ( $\rightarrow$ different implementation of systematics)

# prompt energy spectrum

FD-I  
460.93 days

FD-II  
212.21 days

ND  
150.76 days



**SD fit:** R+S most precise measurement  ${}^8\text{Li}+{}^9\text{He}$  &  $\theta_{13}$  (for each)

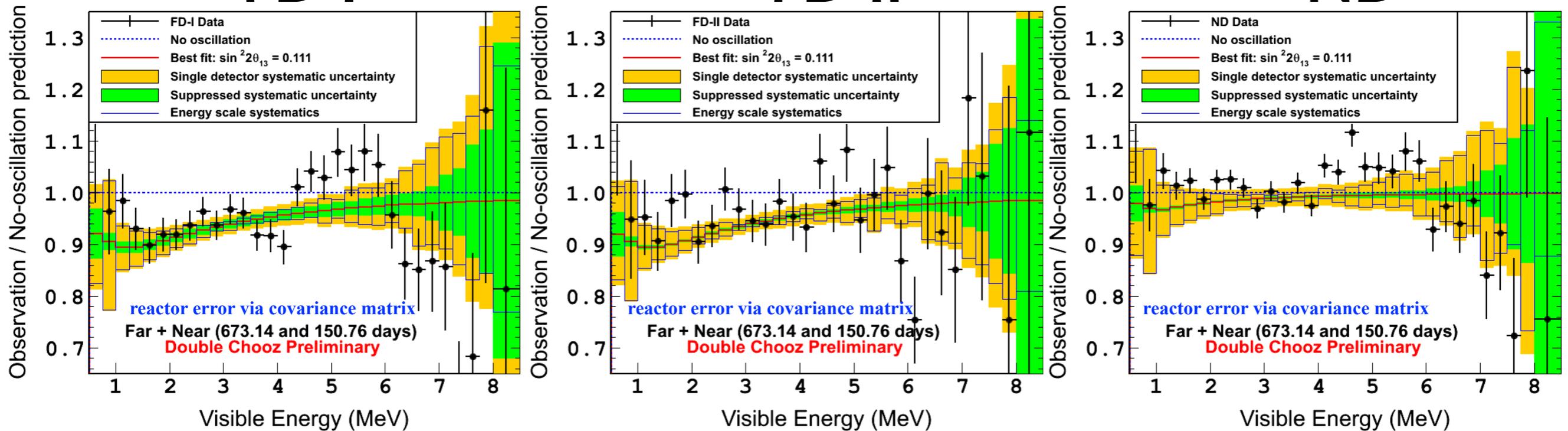
# Fit results

Double Chooz Preliminary

## FD-I

## FD-II

## ND

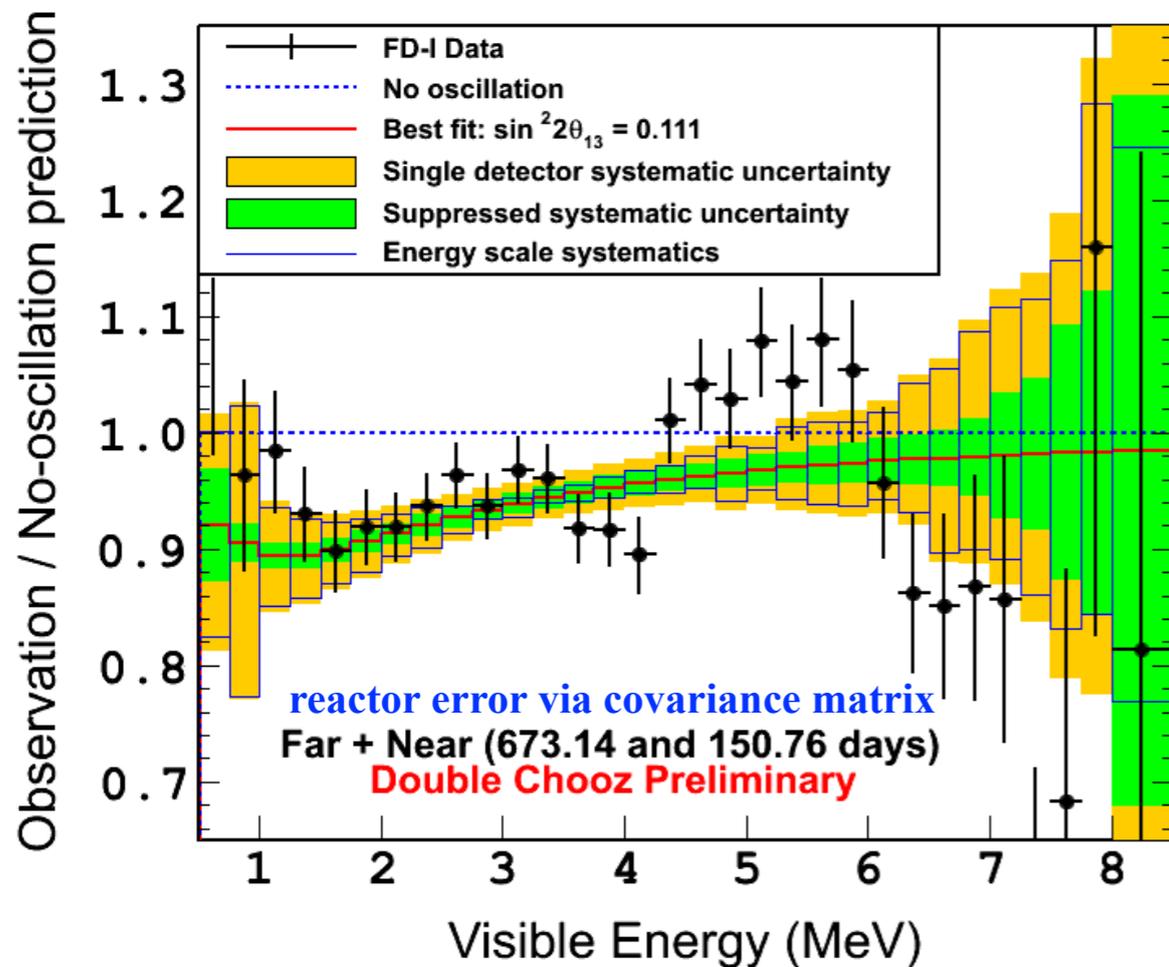


**best-fit:  $\sin^2 2\theta_{13} = 0.111 \pm 0.018$**  (stat.+syst.) ( $\chi^2/\text{dof} = 128.8/120$ )

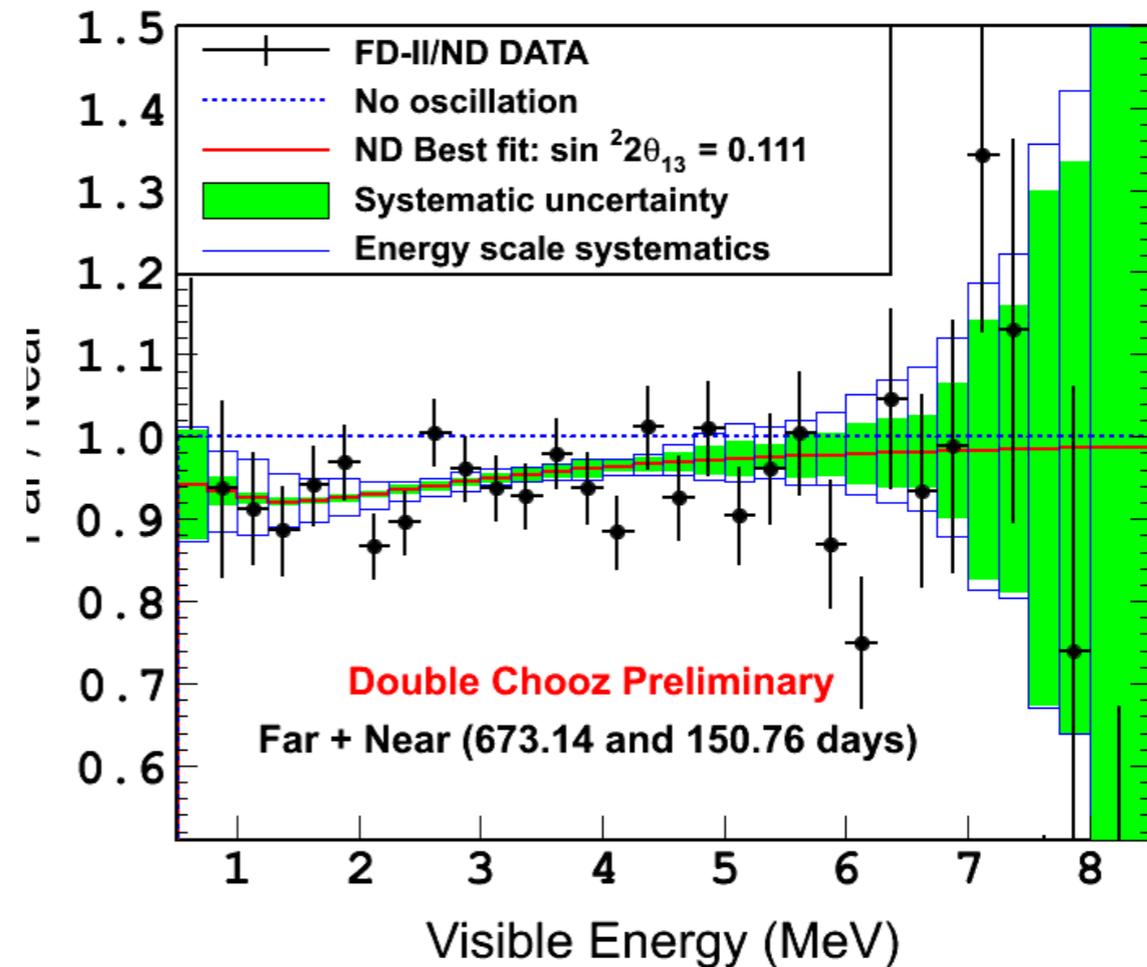
- marginalised over latest MINOS'  $|\Delta m_{23}^2| = [2.28-2.46] \times 10^{-3} \text{eV}^2$  (Normal Ordering)
- non-zero  $\theta_{13}$  observation at  $5.8\sigma$  C.L.
- cosmogenic  ${}^9\text{Li}+{}^8\text{He}$  BG:  $0.75 \pm 0.14 \text{ d}^{-1}$  (FD),  $4.89 \pm 0.78 \text{ d}^{-1}$  (ND)
- fast-n (stop- $\mu$ ) BG:  $0.535 \pm 0.035 \text{ d}^{-1}$  (FD),  $3.53 \pm 0.16 \text{ d}^{-1}$  (ND)
- **energy non-linearity:** consistent across all detector as well as calibration data (non-trivial)

# FD-I $\oplus$ FD-II / ND ratio

## FD-I data/prediction



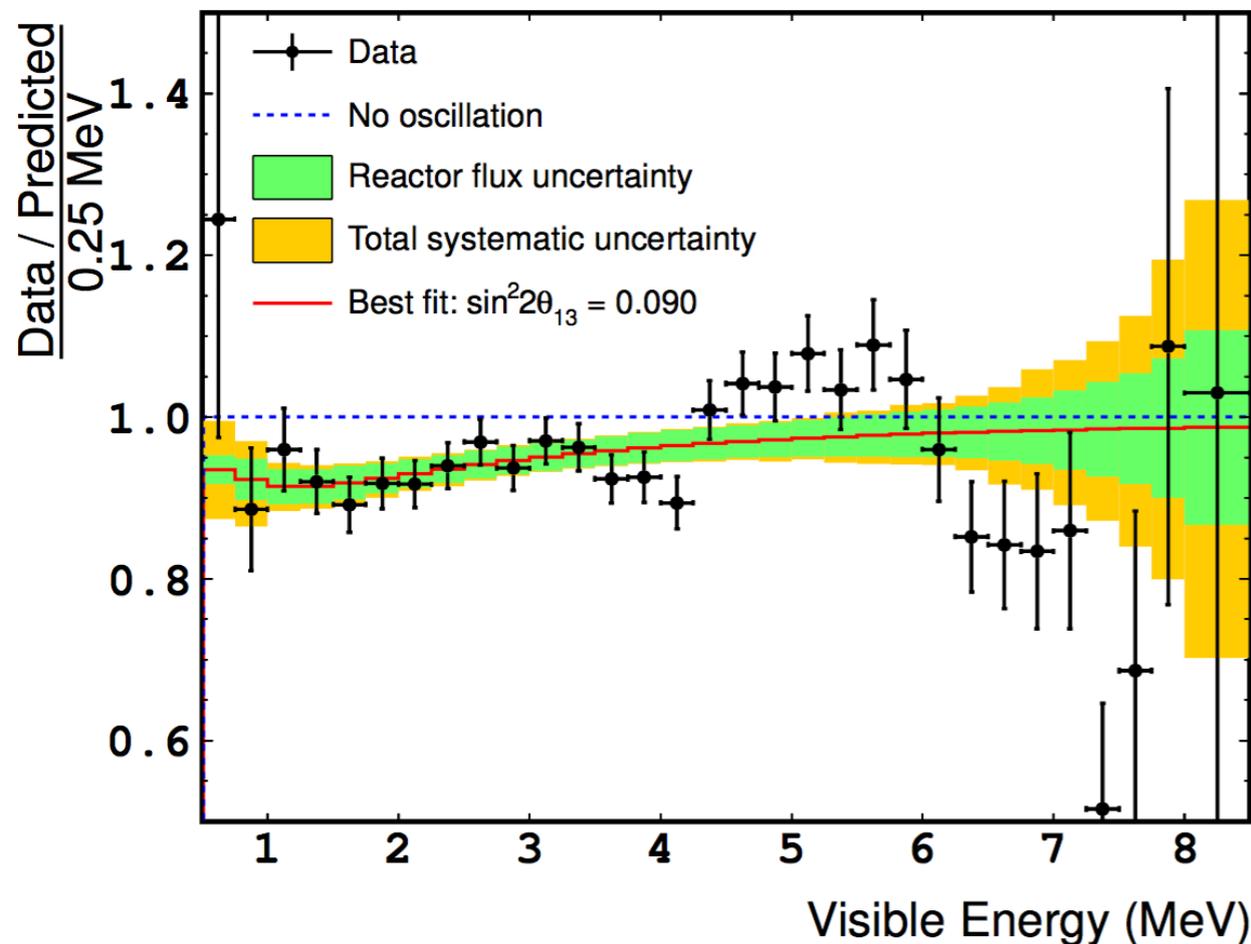
## FD-II data/ ND data



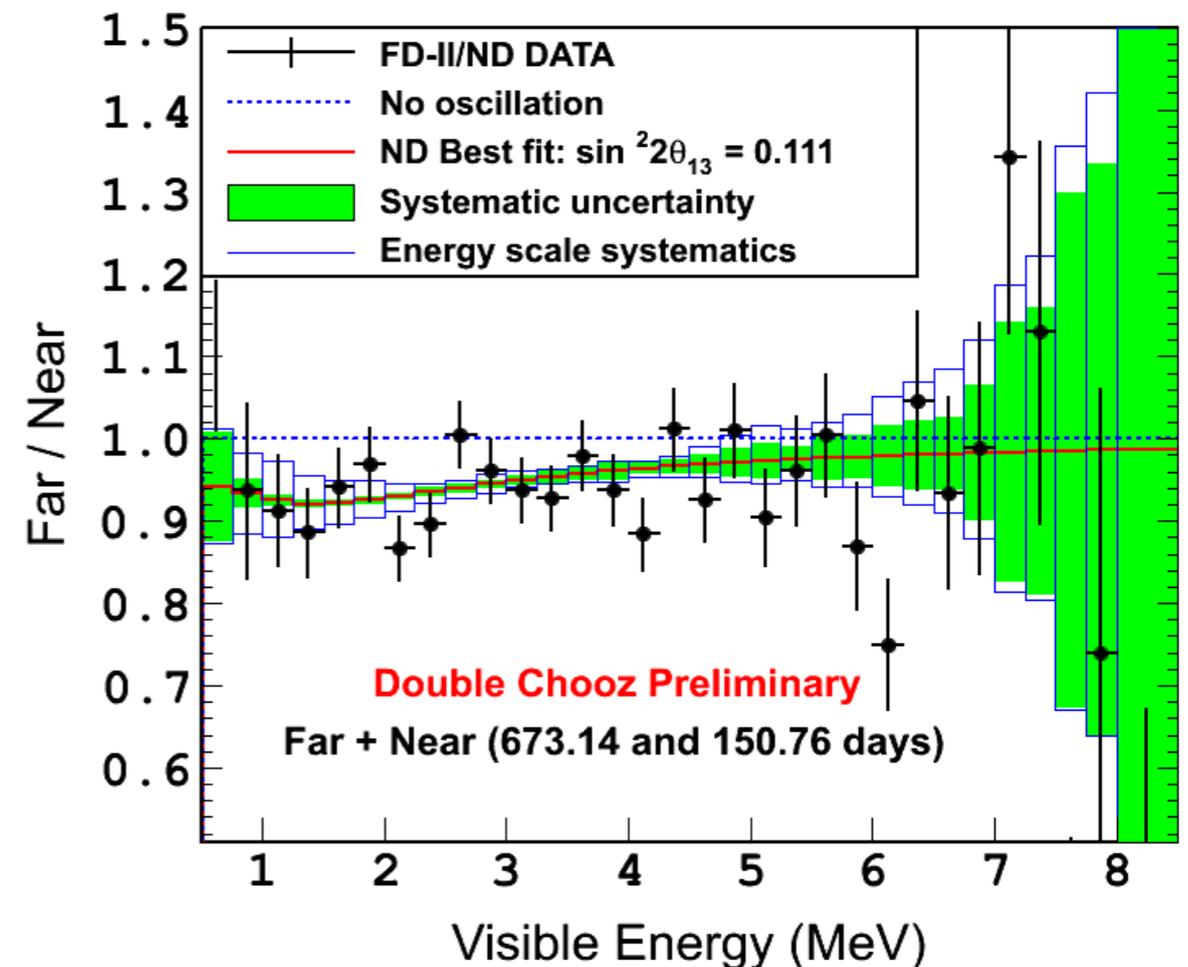
- Systematic uncertainties suppressed in FD-ND relative comparison
- Currently energy uncertainties are assumed to be uncorrelated across detectors (conservative approach)
- ⇔ strong correlation expected with the same scintillator and electronics

# DC-Gd-III (SD) vs DC-Gd-IV (MD)

FD-I data/prediction MC (Gd-III)

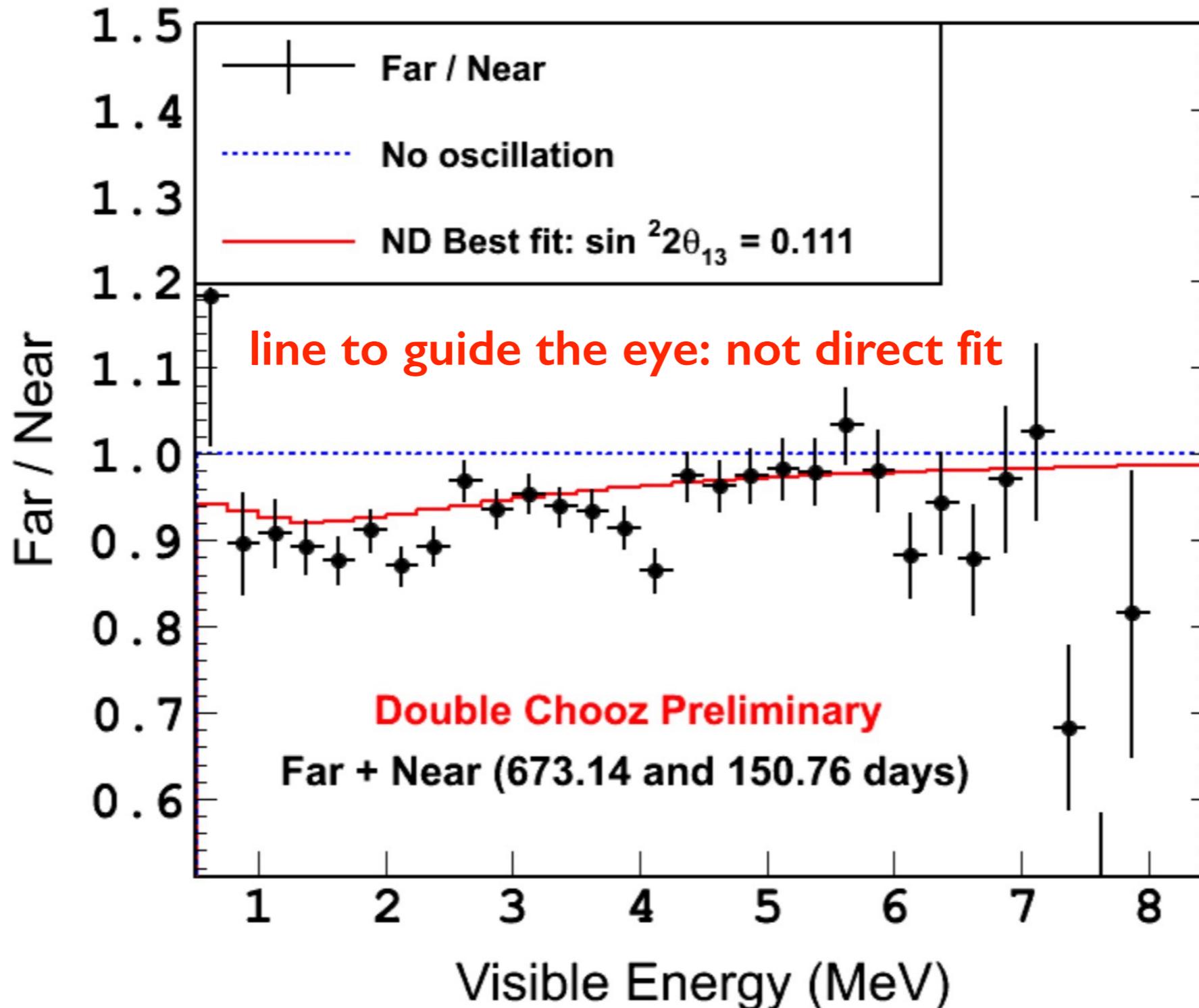


FD-II data/ ND data (Gd-IV)



- **DC-IV:** remarkable systematics suppression across ND:FD (flux: 0.07%)
  - very conservative treatment of energy-linearity: uncorrelated all detectors
  - lack of statistics in FD-II (9months only)
- **DC-III:** very precise single-detector analysis (detection $\oplus$ energy:  $\sim 0.6\%$ )

# (FD-I ⊕ FD-II) / ND ratio



[consistency guideline](#) with [additional FD-I](#) (higher stats → shape)

in terms of  $\sin^2(2\theta_{13})$ ...

$\Delta(\text{DC:DBY}) \sim +35\%$   
[significance  $\sim 1.6\sigma$ ]

$$\sin^2(2\theta_{13})^{\text{DC-IV-M}} = (0.111 \pm 0.018)$$

$$\sin^2(2\theta_{13})^{\text{DB4}} = (0.084 \pm 0.005)$$

in terms of  $\sin^2(\theta_{13})$ ...

$$\sin^2(\theta_{13})^{\text{DC-IV-M}} = (0.0286 \pm 0.0048)$$

$$\sin^2(\theta_{13})^{\text{DYB4}} = (0.0215 \pm 0.0013)$$

**implications...?**

# Double Chooz $\theta_{13}$ in the world

**Double Chooz**  
JHEP 1410, 086 (2014)

**Preliminary (Moriond)**

**Daya Bay**  
PRL 115, 111802 (2015)

**RENO**  
Preliminary (arXiv:1511.05849)

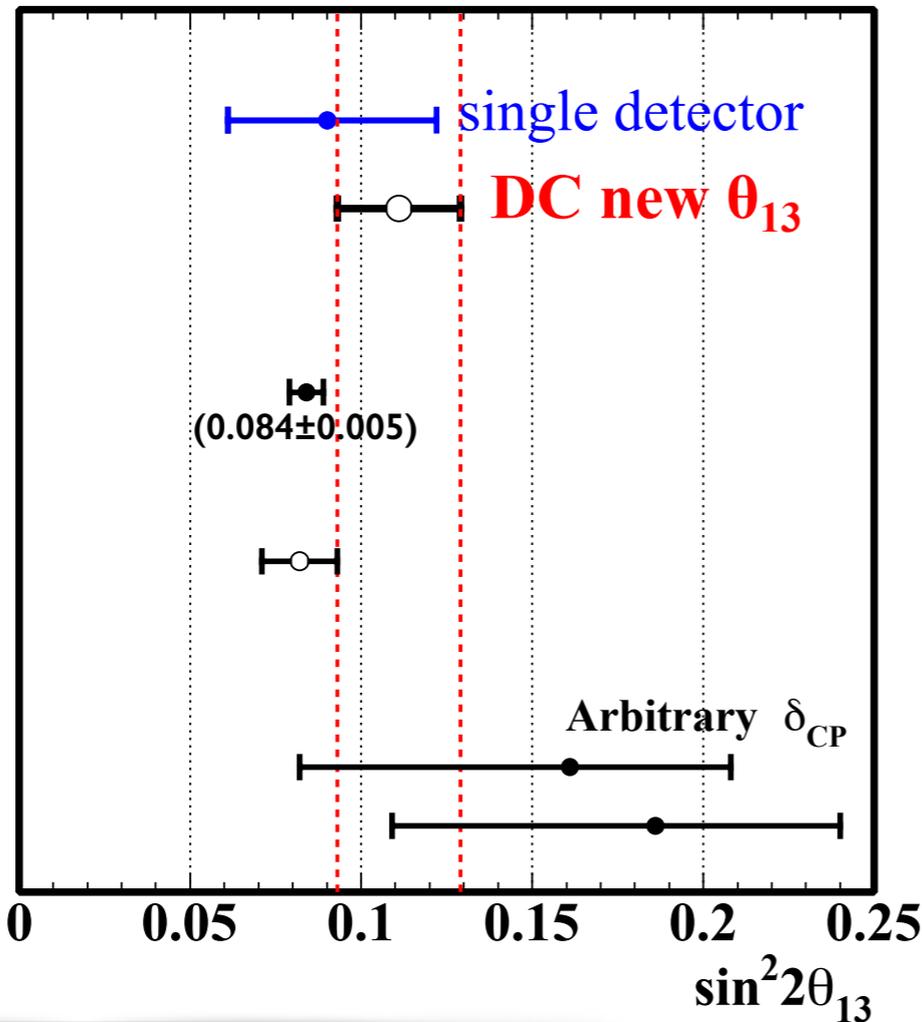
**T2K**  
PRD 91, 072010 (2015)

● published

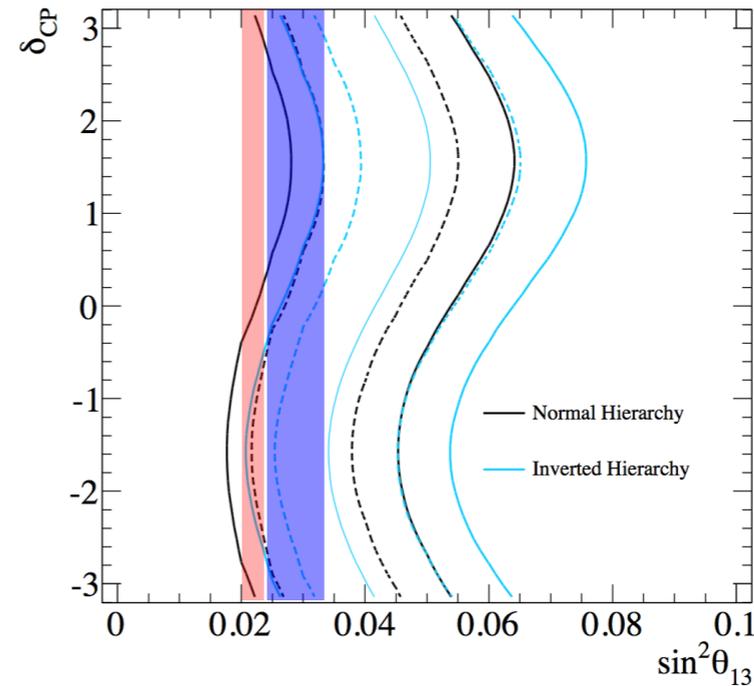
○ preliminary

$\Delta m_{32}^2 > 0$   
 $\Delta m_{32}^2 < 0$

World  $\theta_{13}$  comparison



**@Moriond-2016: DYB Gd-n+H-n: (0.082±0.004)**



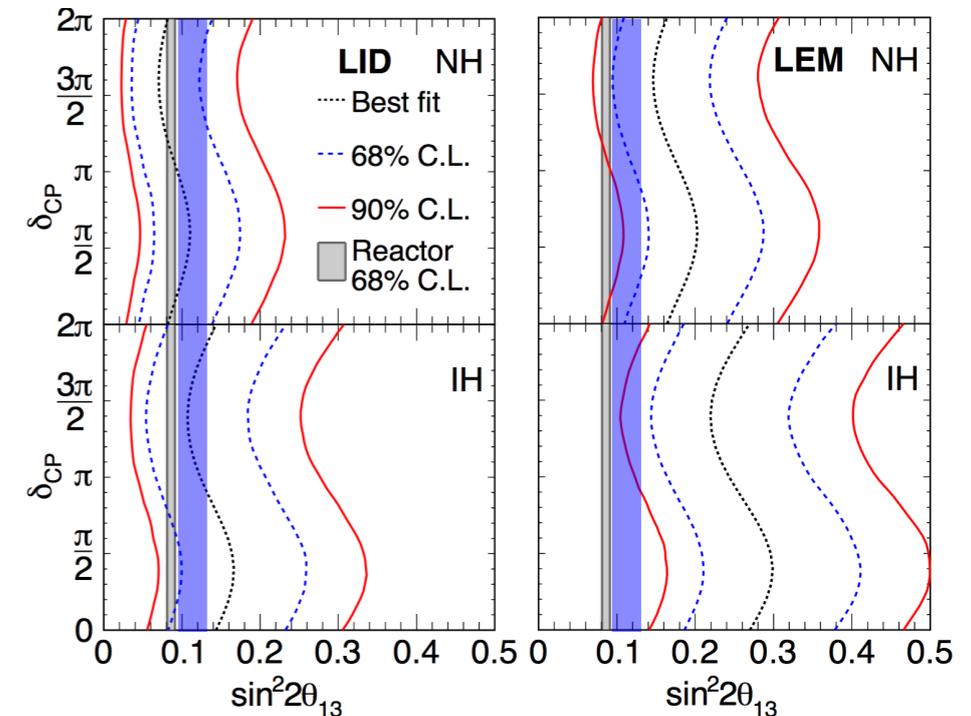
Reactor vs. T2K

PRD91 072010 (2015)

Double Chooz 1σ  
Daya Bay 1σ

Reactor vs. NOvA

arXivL1601.05522 (accepted by PRL)



- DC  $\theta_{13}$  is higher than other reactor  $\theta_{13}$  by  $\sim 35\%$  ( $1.4\sigma \rightarrow 1.6\sigma$  wrt DYB)
- Long baseline (T2K, NOvA) weakly favours higher  $\theta_{13}$  than reactor average
- Reactor  $\theta_{13}$  is key parameter to solve CP-violation and mass hierarchy

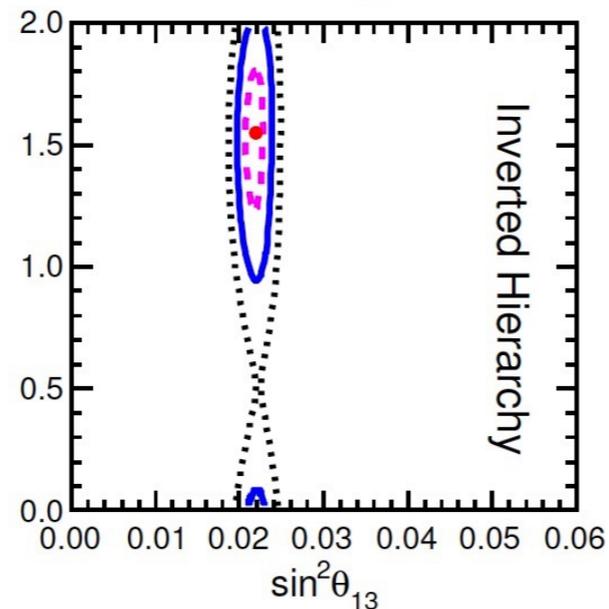
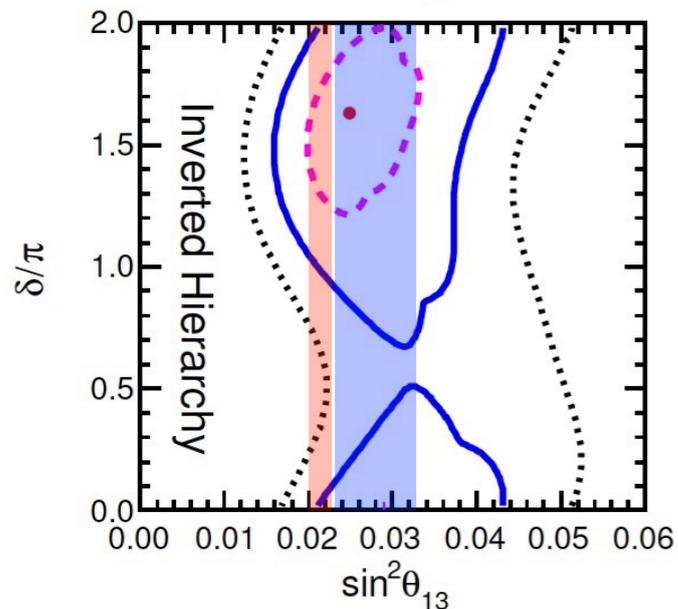
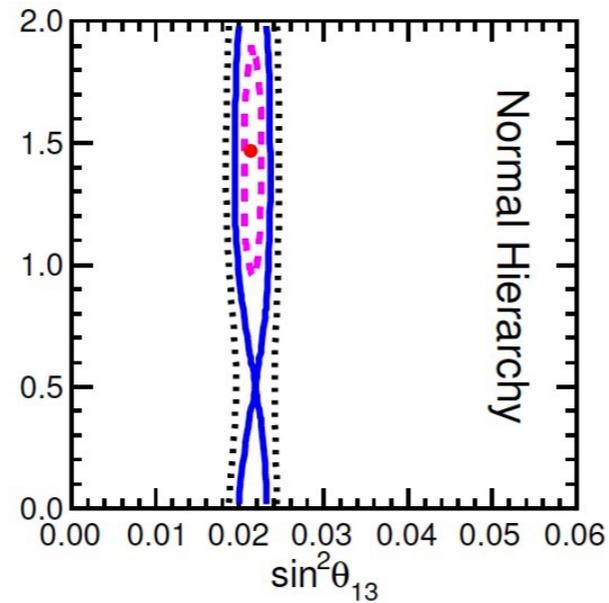
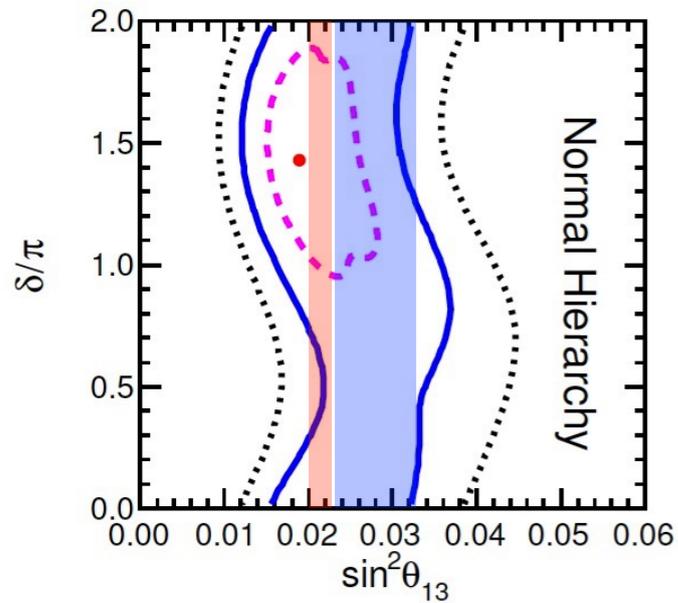
## DC-IV implications to the global picture...

DYB (last)  
DC-IV(Moriond)

reactors before Moriond

LBL Acc + Solar + KL

+ SBL Reactors

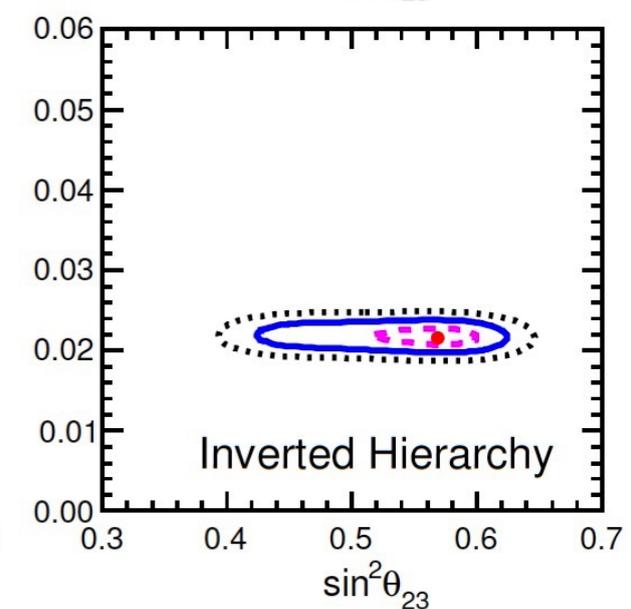
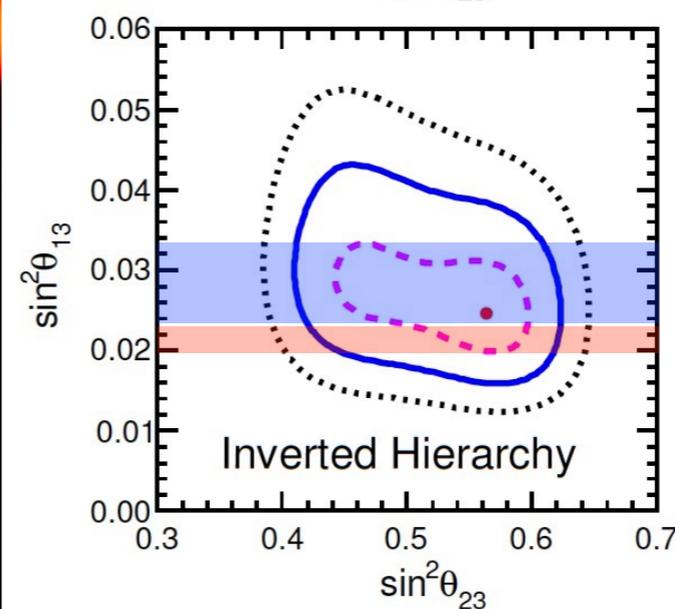
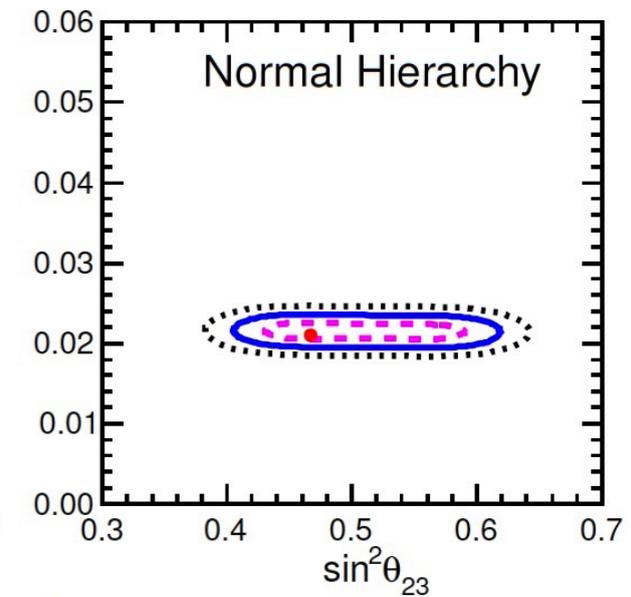
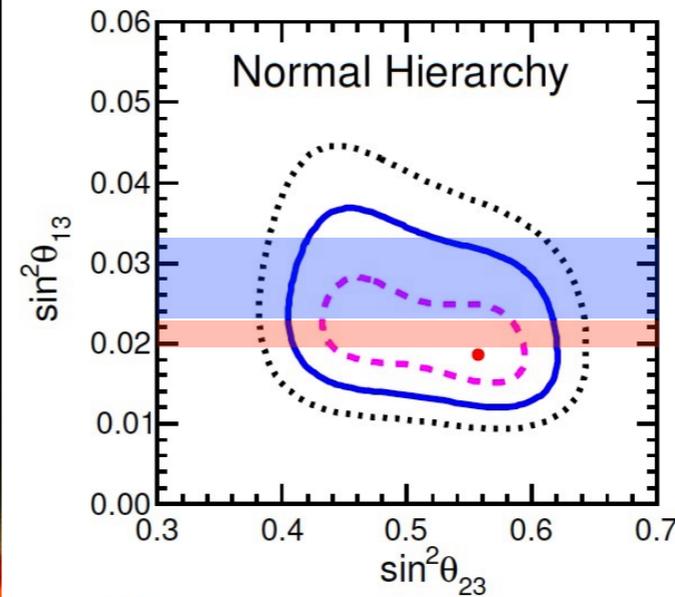


DYB (last)  
DC-IV(Moriond)

reactors before Moriond

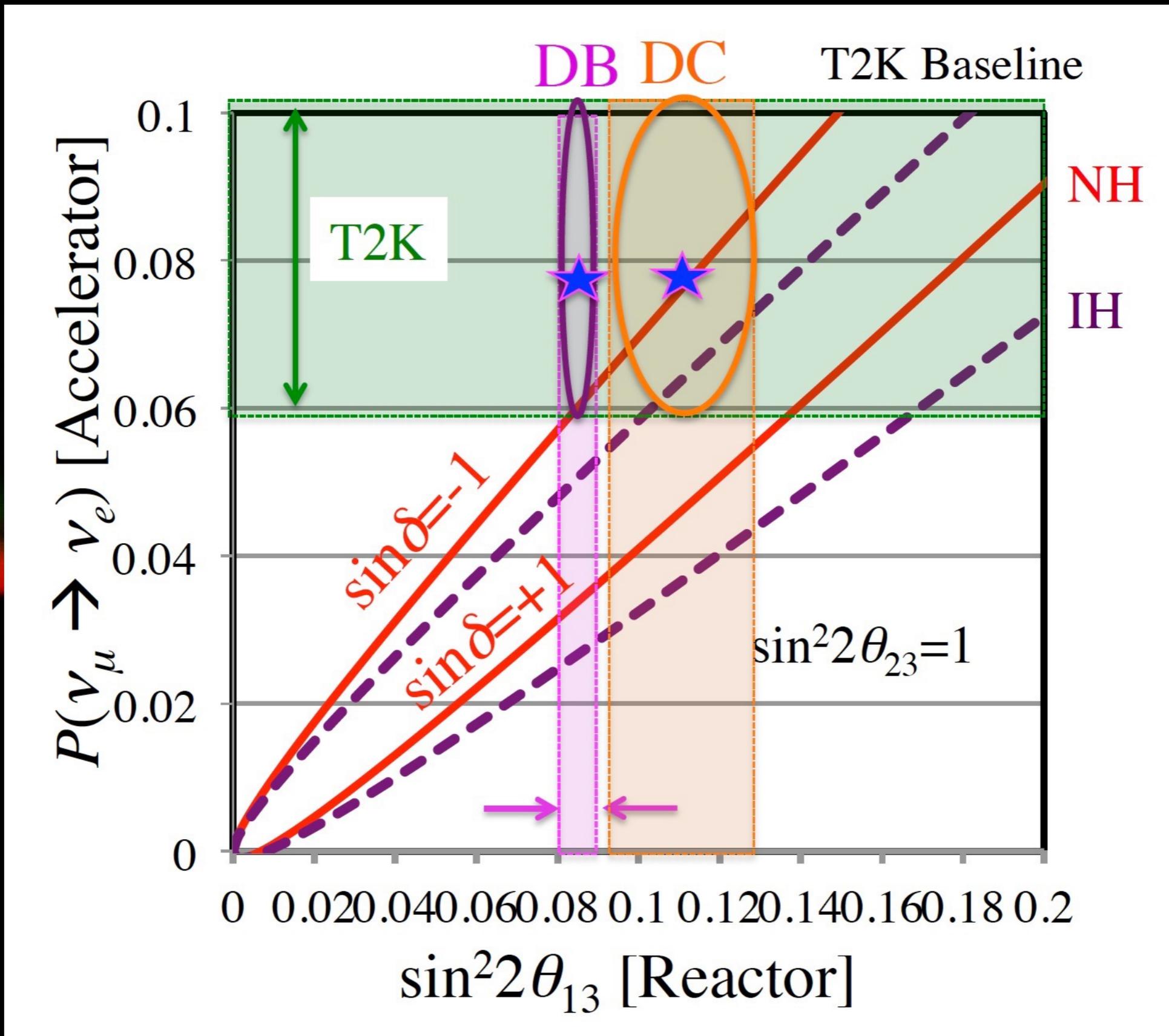
LBL Acc + Solar + KL

+ SBL Reactors

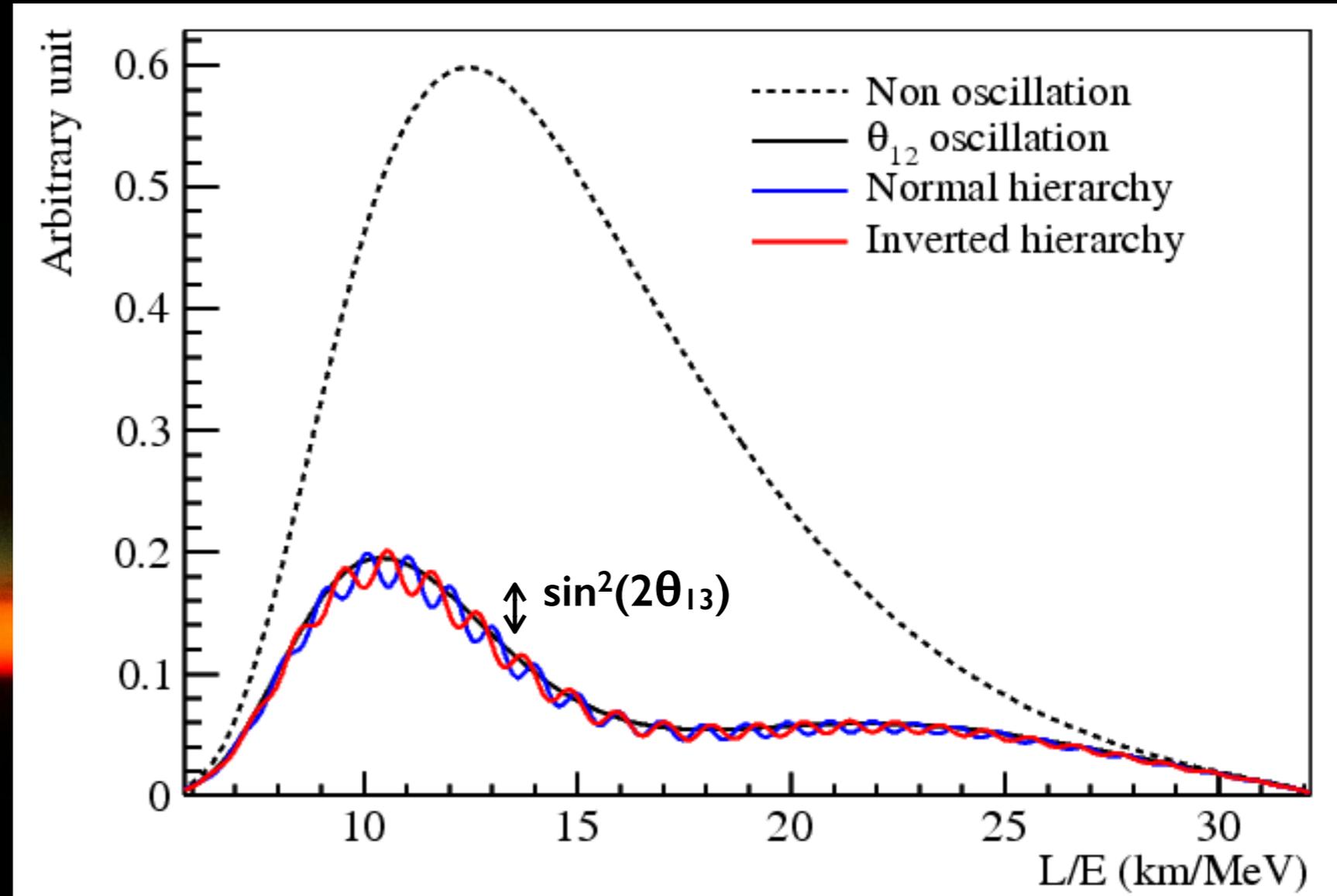
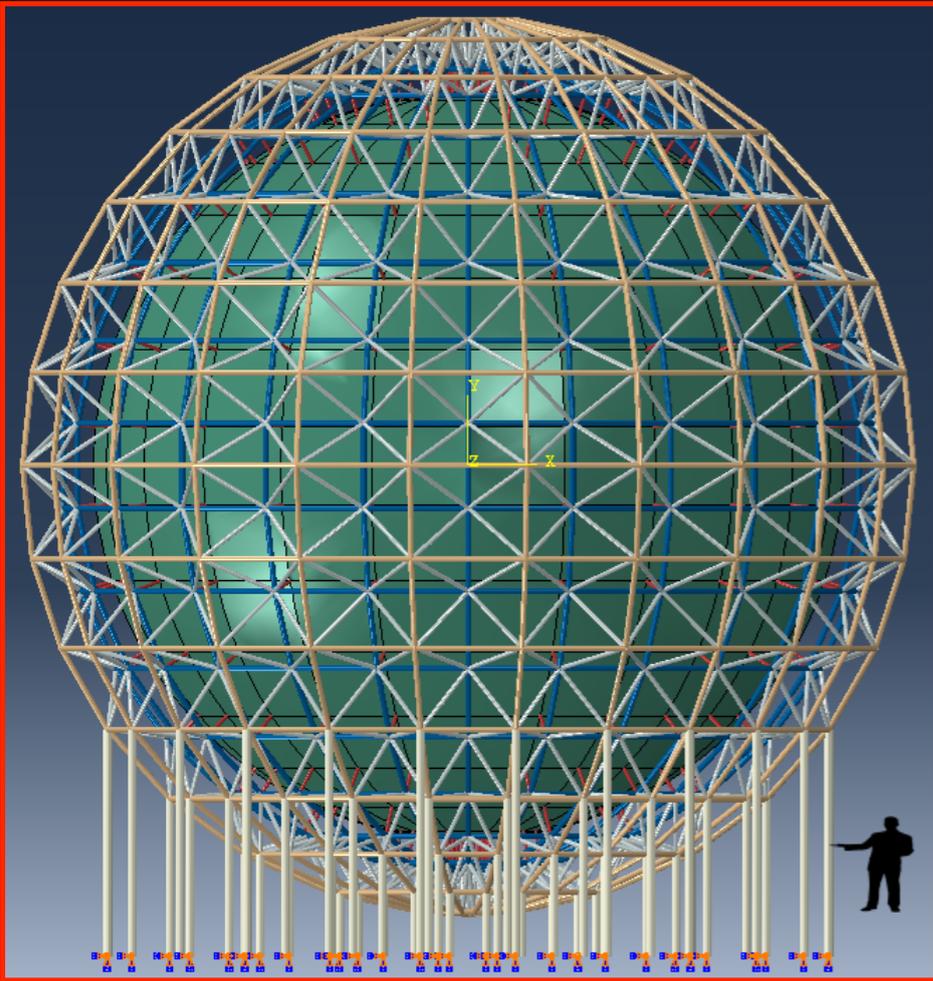


(caveat: drawn regions are very approximative  $\rightarrow$  overlaid)

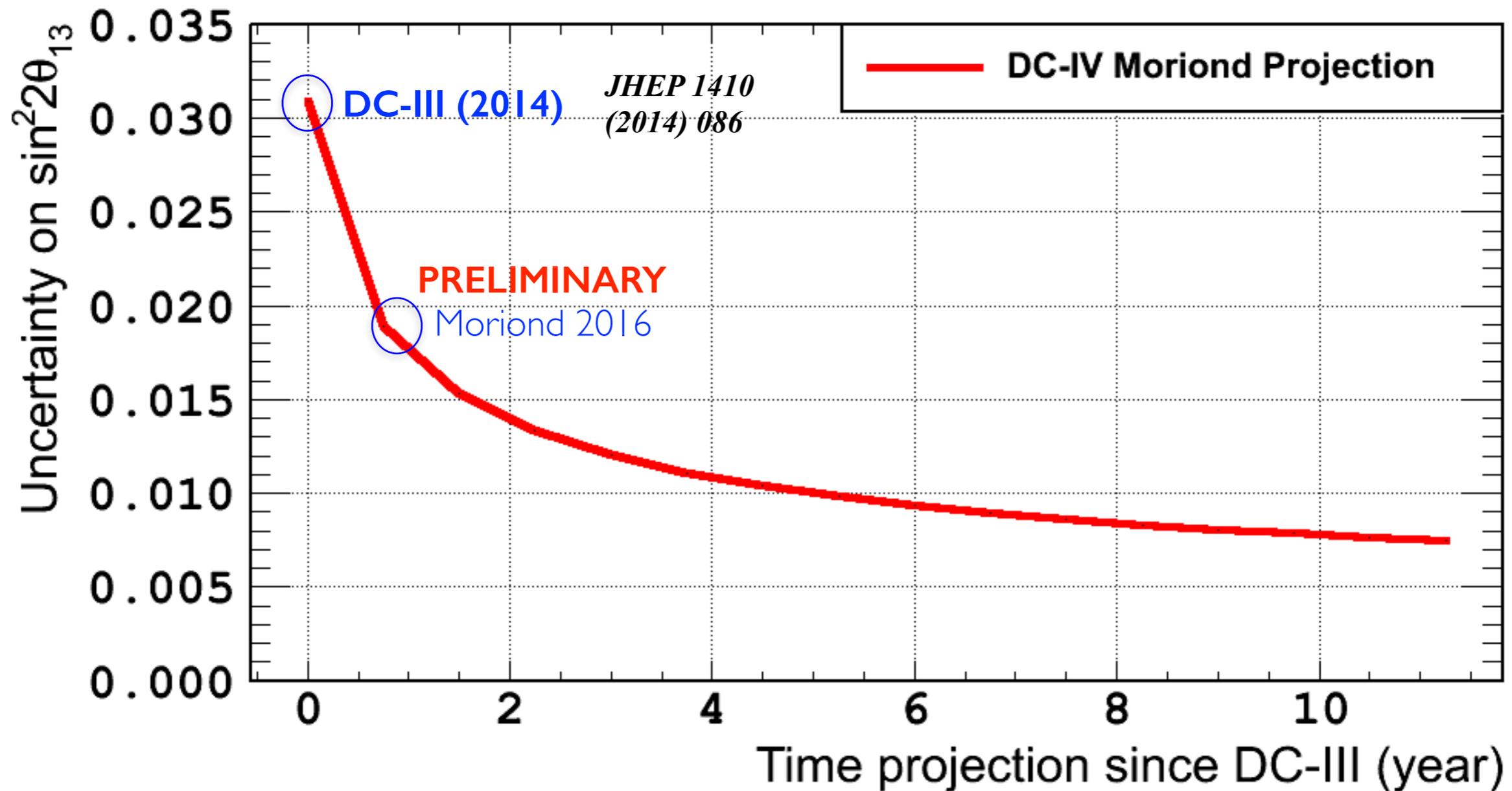
solution phase-space might change if  $\theta_{13}$  was larger (not yet)



courtesy by Suekane-san



**JUNO will benefit from larger  $\theta_{13}$ ... (if true)**



DC-IV can go for >10years without hitting systematics ( $\approx 0.005$ )  
 ( $\Rightarrow$  analysis is too good for detector's size)

— working on higher stats sample —

# Reactor- $\theta_{13}$

*(combining results)*

**Daya Bay  $\oplus$  Double Chooz  $\oplus$  RENO**

**1<sup>st</sup> workshop**  $\rightarrow$  Autumn 2016 (Seoul, South Korea)  
(systematics, results, etc)

**2<sup>nd</sup> workshop**  $\rightarrow$  later on (allow full statistics samples)  
(towards the combined  $\theta_{13}$ )

*[the best  $\theta_{13}$  value for a long while]*

**BUMP**



**AHEAD**

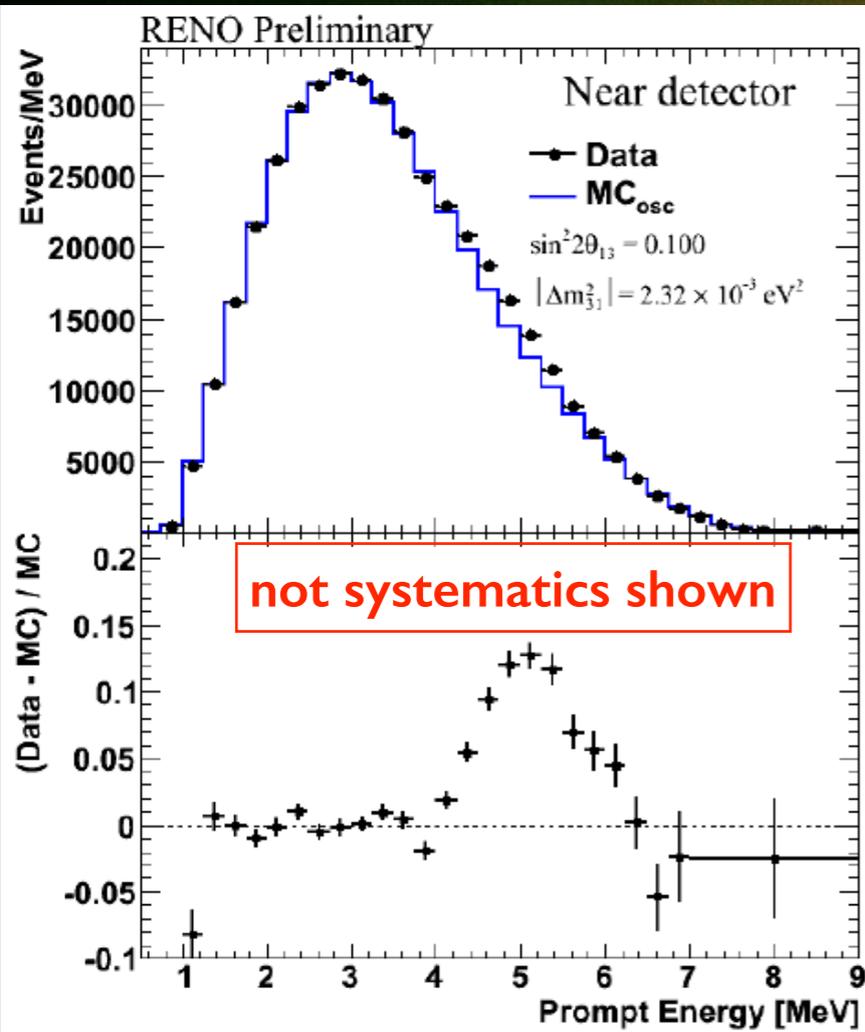
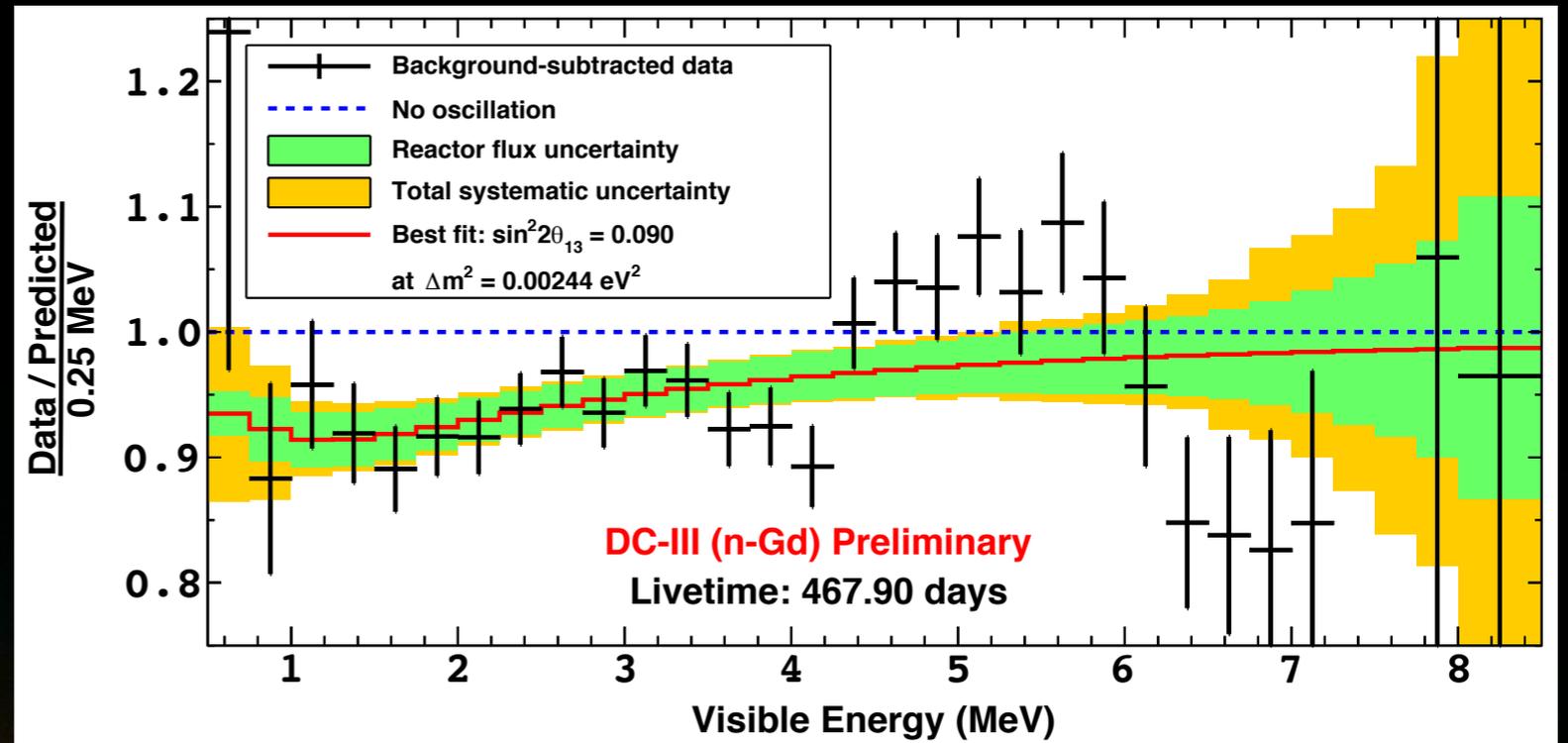
# $\theta_{13}$ experiments $\rightarrow$ high precision spectrum...

Double Chooz (May 2014)

$\sim 3.0\sigma$  ( $\sim 17k$  events @ FD)

$\sigma(\text{stat}): \sim 0.8\%$

$\sigma(\text{detection}): \sim 0.6\%$  (0.4%)



RENO (June 2014)

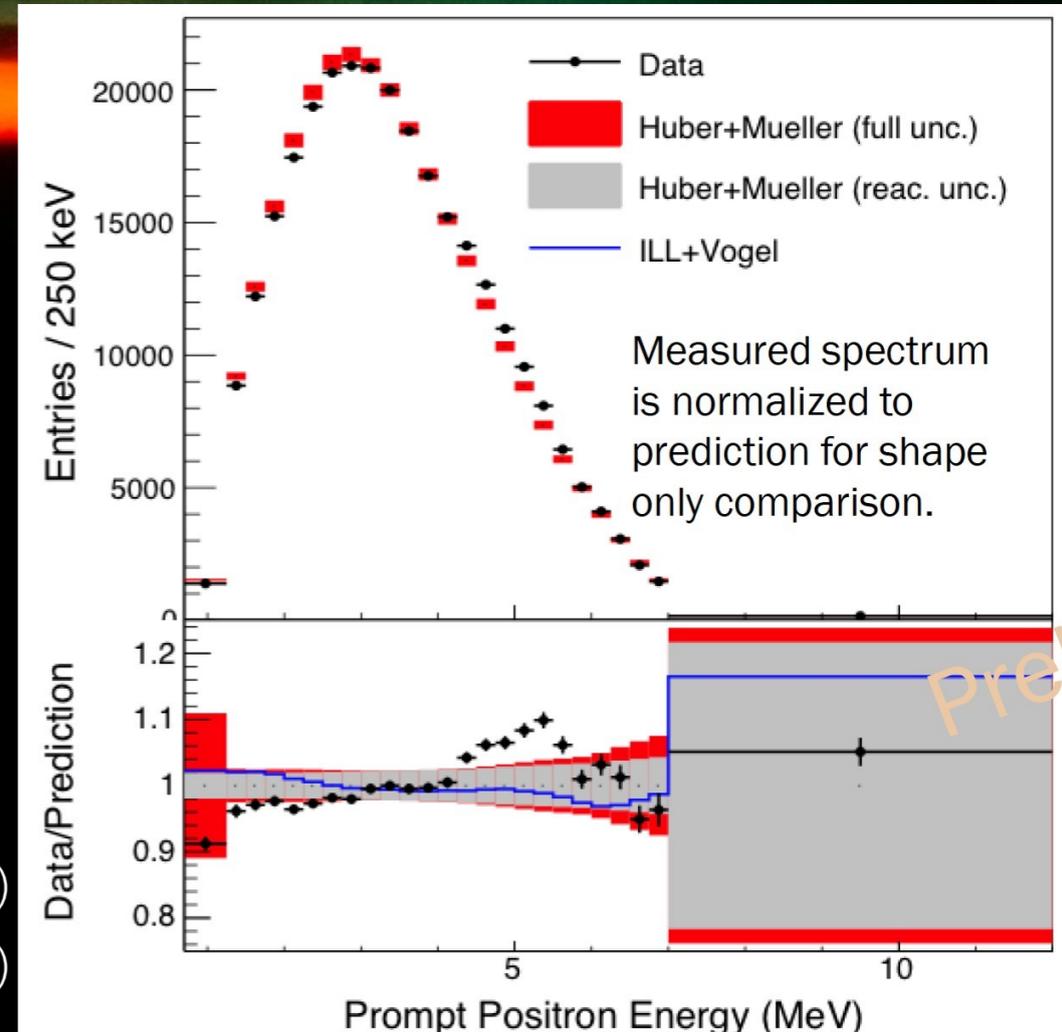
$\sim 3.6\sigma$  ( $\sim 500k$  events @ ND)

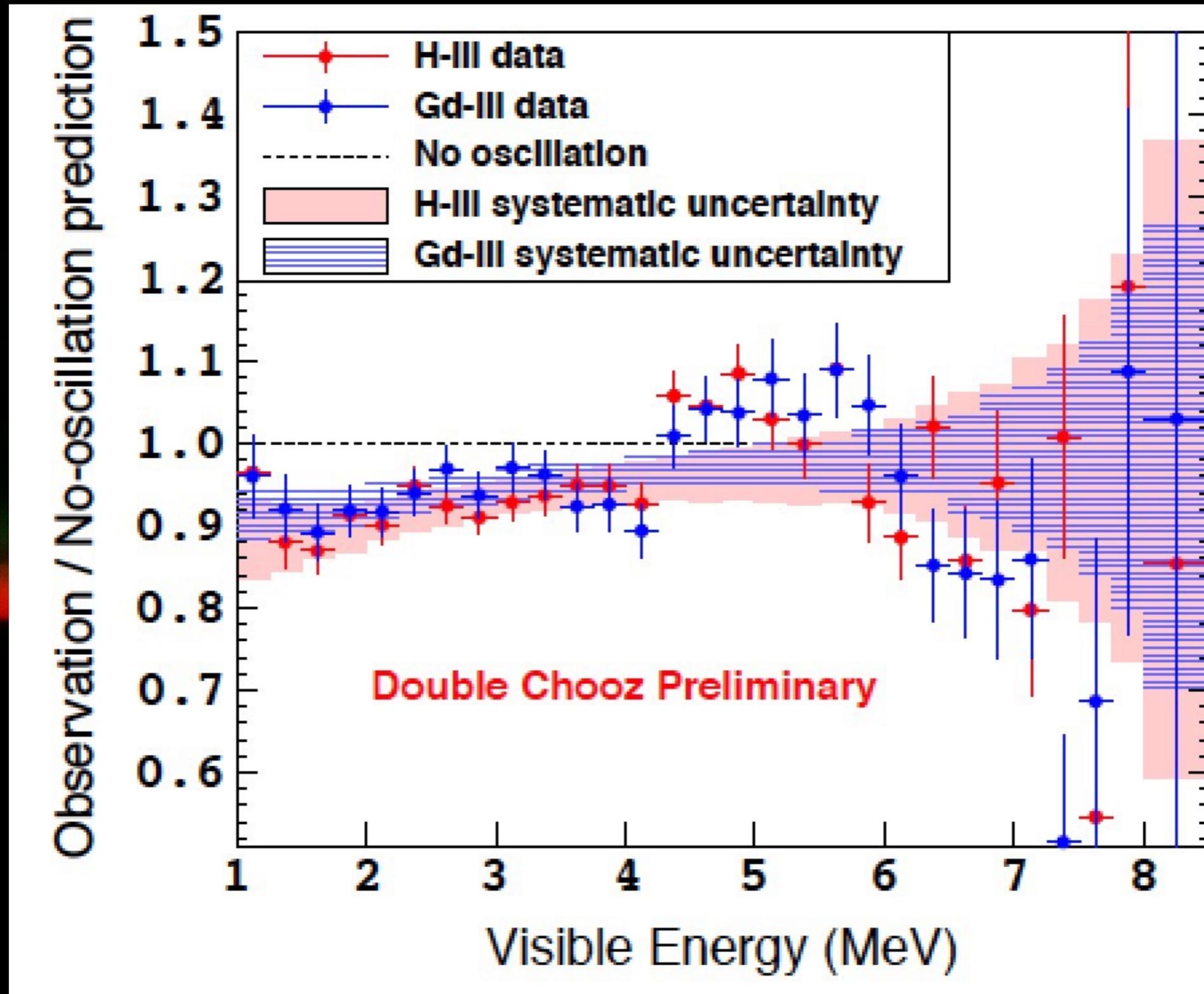
$\sigma(\text{stat}): \geq 0.2\%$

$\sigma(\text{detection}): \geq 2.0\%$

Daya Bay (July 2014)

$\sim 4.0\sigma$  ( $\sim 300k$  events @ 3xNDs)

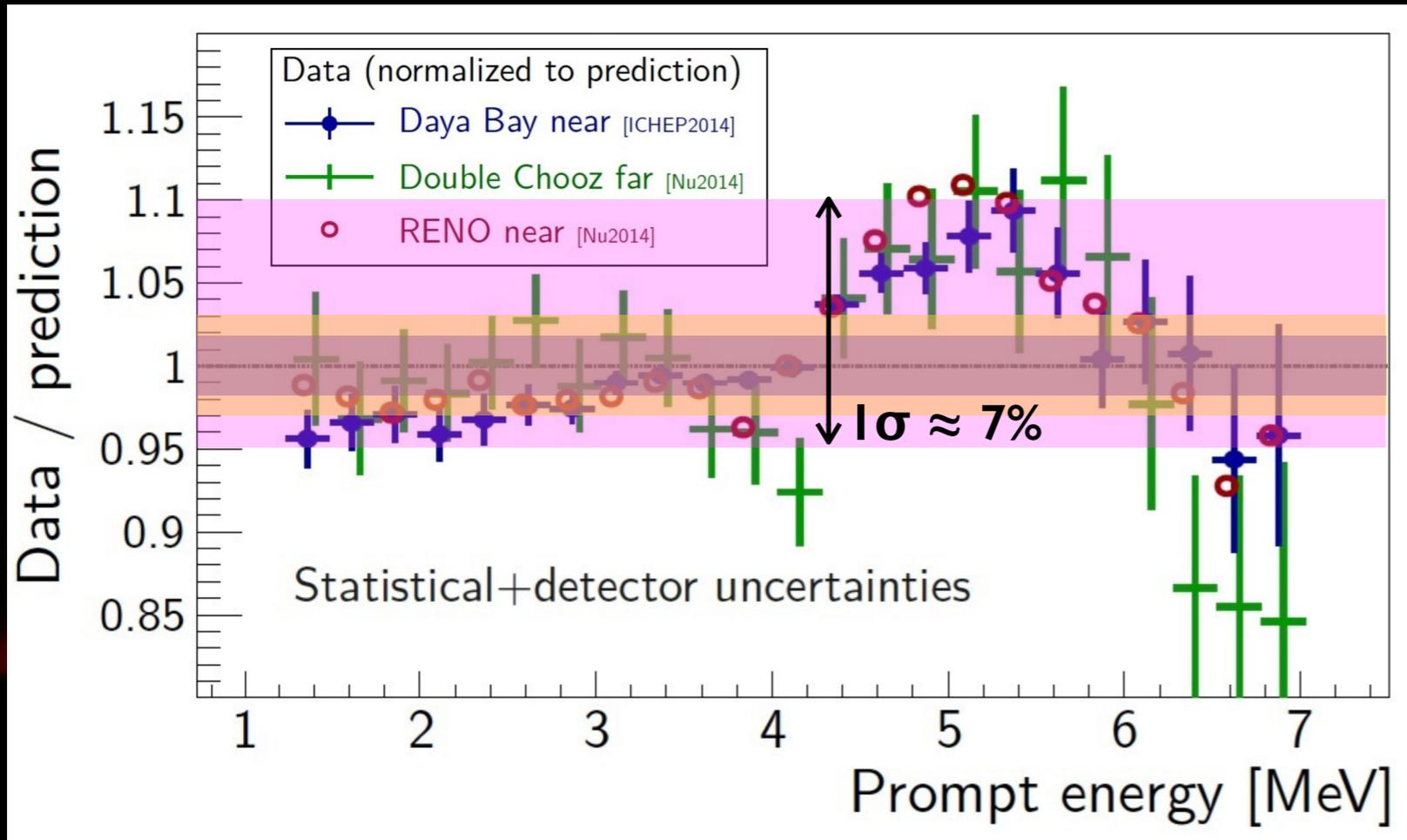




consistent result with both Gd-n and H-n IBDs (independent)  
 (but **different BGs, detection volume, capture mechanism**)

# $\theta_{13}$ (3 experiments) spectral distortion [4,8]MeV...

$1\sigma$  of  $\delta(\text{flux}) \rightarrow \pm 3\%$  (DB & RENO) &  $\pm 1.7\%$  (DC $\oplus$ Bugey4)



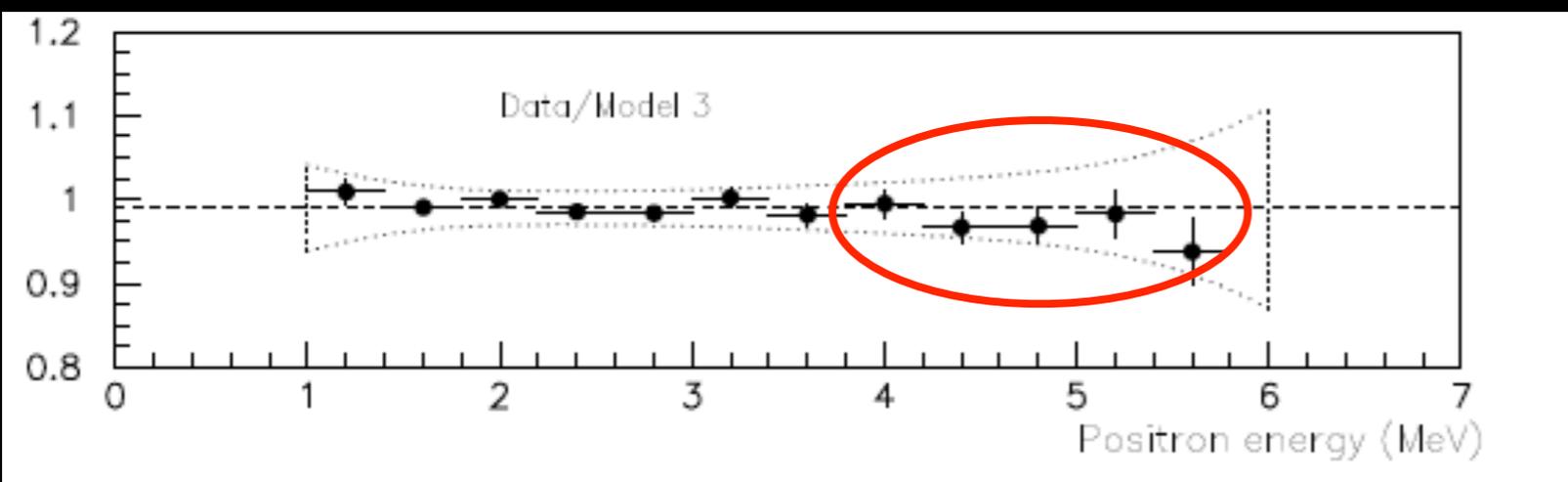
**3 different experiments in agreement**  
(not trivial: not the same fuel)

**$\delta(\text{flux})$  error is very likely to increase**  
(hard to believe otherwise)

$\Rightarrow$  **necessity for neutrino-sterile hypothesis (reactor-data): insignificant?**

**how to set new (reasonable) error?**

**...and what do we do with Bugey3?**  
(could we ever reconcile?)



# 5MeV distortion implies...

## (the problem)

- its origin might likely entail a complex explanation (nuclear, bias, etc...)
  - likely beyond any particle physics interest (not clear yet though)

## (applied neutrino physics)

- reactor neutrino physics precision surpasses reactor nuclear knowledge?
  - neutrino data use to understand better reactor physics (first time ever)

## (effective fact)

- incompatible with the  $\sim 3\%$  error (conversion method) → **error must be larger**
  - how to set a new error? (not easy)
  - **$\sim(3,7]\%$  error** to be used as effective reference (in the meantime)

## (particle physics implication)

- any fundamental physics conclusion/hypothesis relying on reactor-spectrum knowledge (error budget) must be taken with extreme caution — to say the least.

(my opinion)

# BUMP

**tip:** don't go too fast....!



# AHEAD

A joint Fermilab/SLAC publication

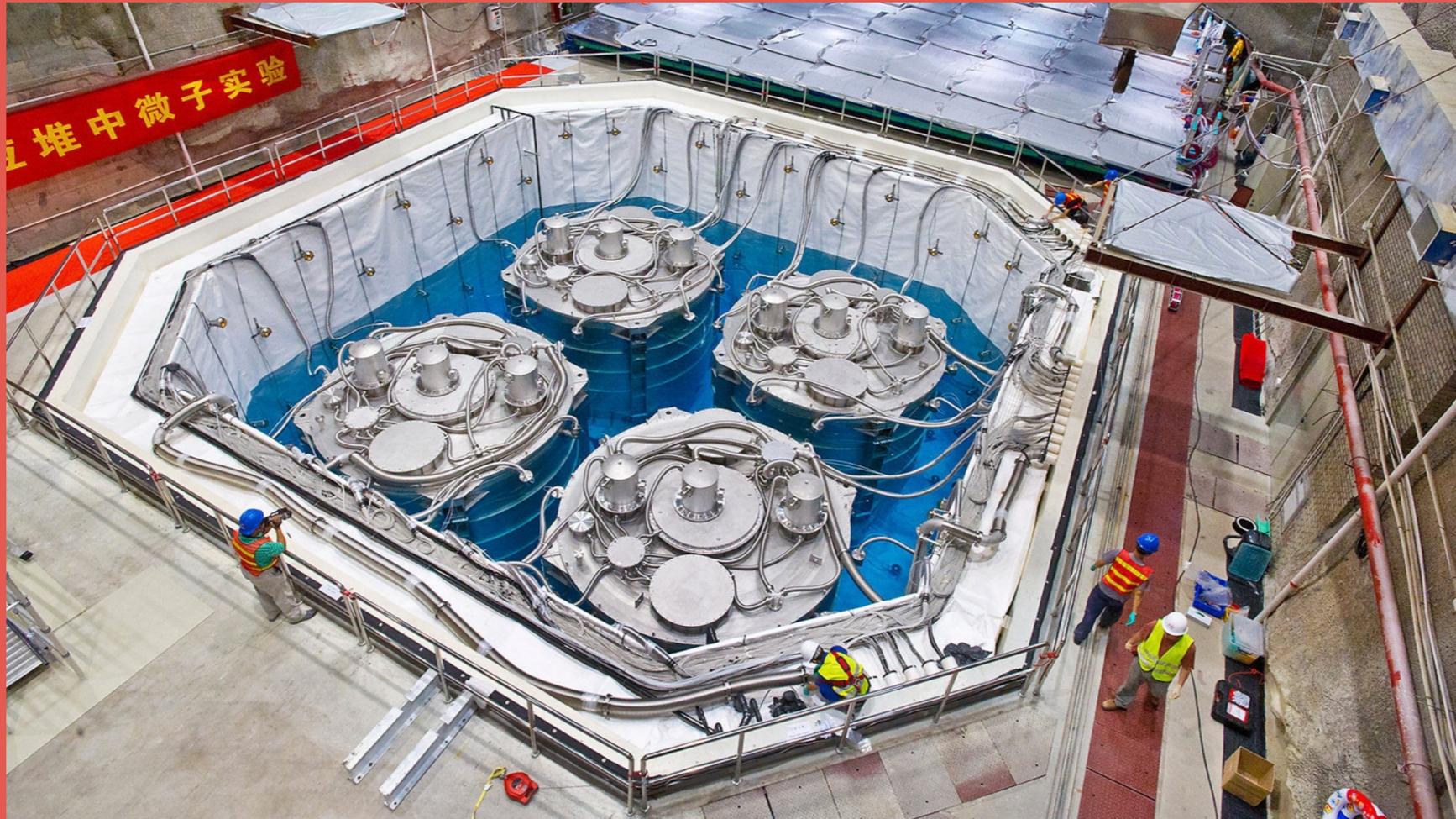


Photo courtesy of Brookhaven National Laboratory

## Daya Bay discovers a mismatch

02/12/16 | By Kathryn Jepsen

The latest measurements from the Daya Bay neutrino experiment in China don't align with predictions from nuclear theory.

actually, prediction is (much) based on data (ILL, fits, etc)

no collaboration (so far) has claimed any “discovery” → **why symmetry does?**

# what to remember?



DC Proposal: hep-ex/0606025

$\theta_{13}$ ” [6]. But since its publication the worldwide situation has changed and the projects still being considered are Angra [7] in Brazil, Daya Bay [8] in China, Double Chooz in France (see [9, 10] and this proposal), KASKA [11] in Japan and RENO [12] in South Korea. A recent comparison of the capabilities of these experiments can be found in [13, 14]. Double Chooz is particularly attractive because it could limit  $\sin^2(2\theta_{13})$  to 0.022-030 (for  $\Delta m_{31}^2 = 3.5 - 2.5 \times 10^{-3} eV^2$ ), within an unrivaled time scale and a modest cost. Installation of the experiment will start with the far detector located

**Double Chooz performance is better than ever...**

reached proposal goal (@3years) with only 9months of ND+FD data  
(indeed with a “unrivalled time scale”)

(DC delays arguably not worse than most HEP experiment → DYB & RENO superb timing!!)

# Summary

- Double Chooz collaboration reported first  $\theta_{13}$  measurement with two detectors (FD-I: 460.93 days + FD-II: 212.21 days + ND: 150.76 days)
  - **$\sin^2 2\theta_{13} = 0.111 \pm 0.018$**  (stat.+syst.)
  - reactor flux uncertainty strongly suppressed to  $< 0.1\%$  (almost iso-flux)
    - ND: un-oscillation spectrum direct observable of experiment
  - other systematic uncertainties  $\leq 0.5\%$  (soon: improvements expected)  
(after analysis improvements made in single detector phase)
- reactor- $\theta_{13}$  is a key for current and future neutrino projects aim to solve still unknown CP-violation and mass hierarchy (and other observables)  
 $\Rightarrow$  **validation by multiple-experiments is essential: agreement?**
- precision now limited by statistics (systematics also dominated by stats)  
 $\Rightarrow$  **further improvements expected from Double Chooz (very soon!)**

**1.6 $\sigma$  means 1.6 $\sigma$**

(no more & no less)

DC: much better error **soon!**  
(best value: up to Nature's will)

merci....  
thank you....

## • integrated data and MC calibration scheme...

- MC treated independently (as two detectors)
- MC (no free knobs → lab measurement + calibration)

## • Linearised-PE & “ $\alpha$ ” Calibration...

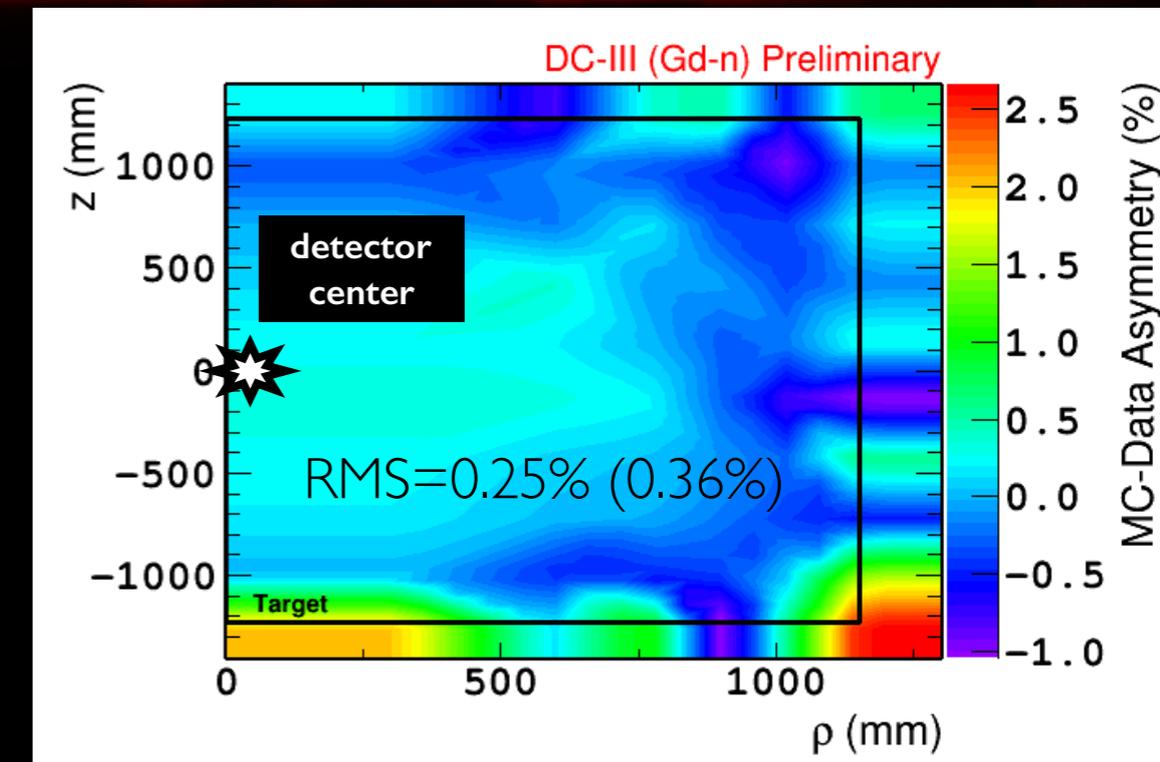
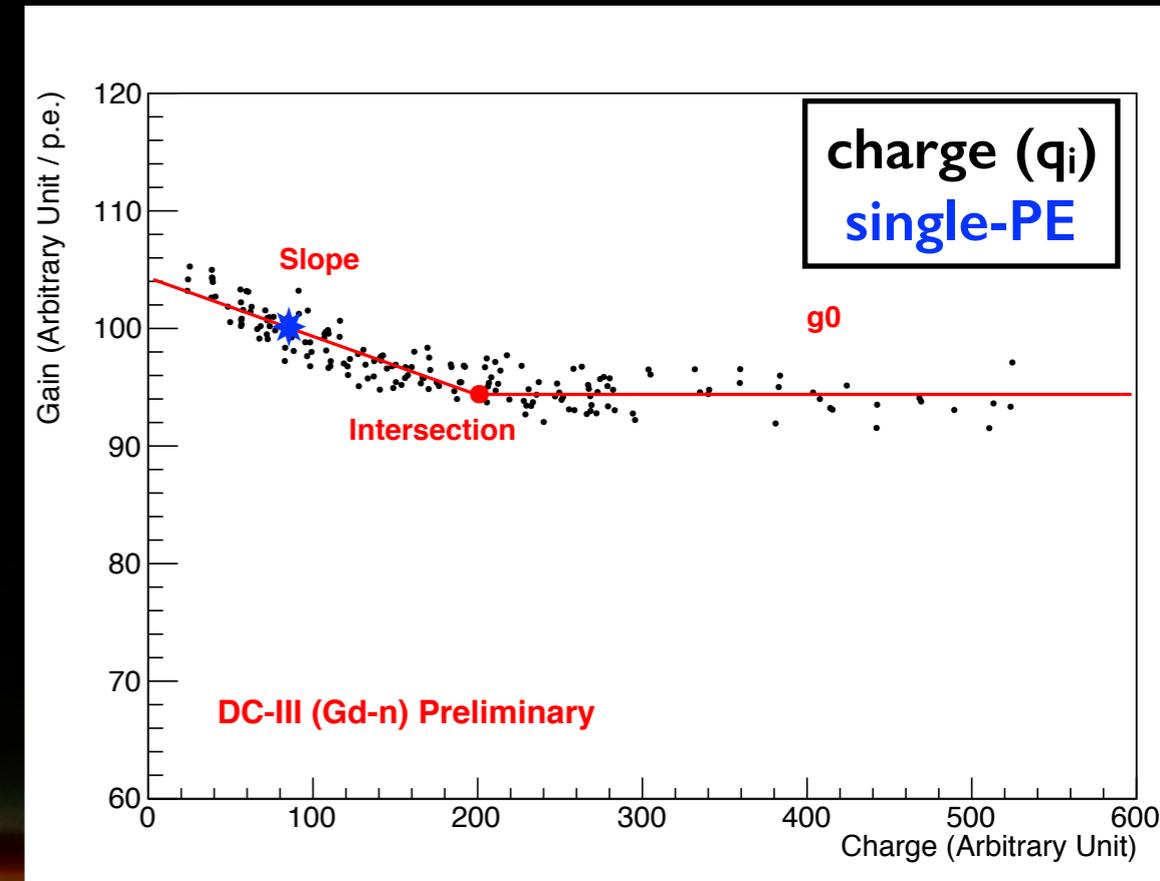
- def:  $PE = \alpha(PE, \#PMT \text{ hit}) \times \{\sum q_i \times g(q_i)\}$
- conversion  $Q[\Delta \sim 5\%] \rightarrow PE[\Delta \leq 0.5\%]$  @ H-n peak center
- impact: **stability (+++)**, **linearity (++)**, **uniformity (+)**
- source: gain non-linear [@electronics] + other (zeroes, etc)

## • Uniformity Calibration...

- def: create H-n response full volume MAP
- conversion  $PE(\rho, z)[\Delta \leq 8\%] \rightarrow PE(\text{center})[\Delta \leq 0.5\%]$
- impact: **uniformity (+++)**

## • MeV (or absolute) Energy Calibration...

- conversion:  $PE(0, \tau) \rightarrow \text{MeV}(0, \tau)$
- use  $^{252}\text{Cf}$  @  $(\rho=0, z=0, t=\tau) \rightarrow$  H-n peak: 2.223 MeV
- DATA to MC equalisation (prior  $<0.5\%$  agreement)



### • Drift Stability Calibration...

- def:  $PE(t) \rightarrow PE(\tau)$ , where  $\tau$ : time MeV definition
- response drift by +0.5%/years (unknown)
- impact: **stability (+)**

### • Charge Non-Linearity Calibration...

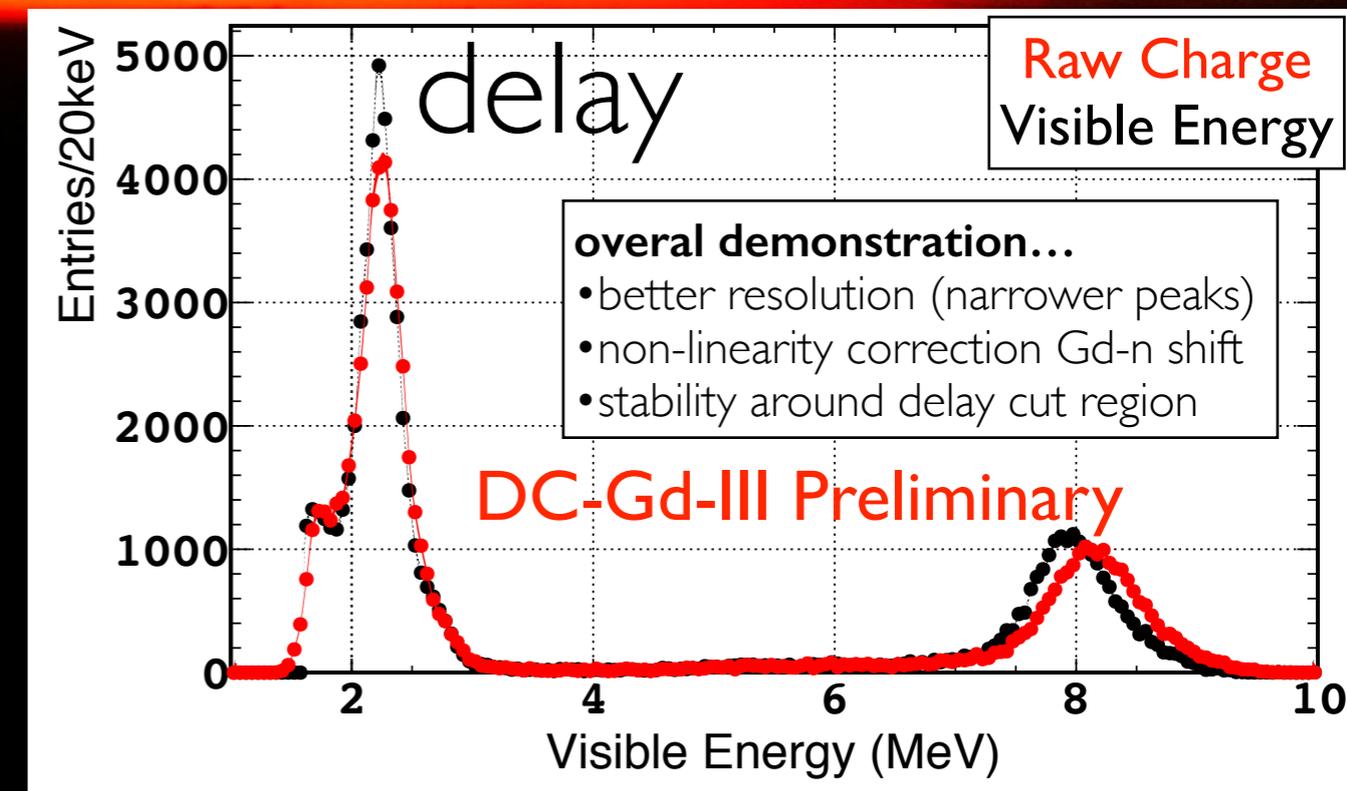
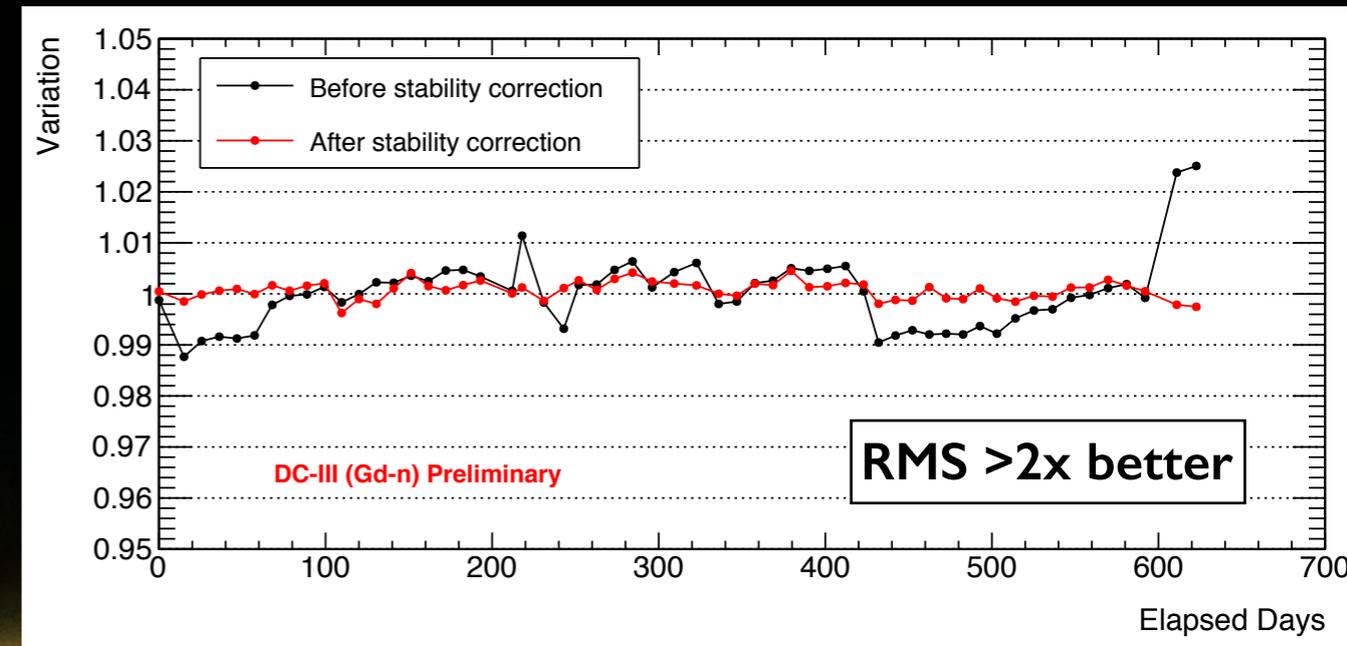
- readout driven-non-linearity  $\rightarrow \Delta(H-n, Gd-n) \sim 1\%$
- validation with C-n peak @ 5MeV &  $^{12}B$  spectrum
- impact: **linearity (++)**

### • Light Non-Linearity Calibration...

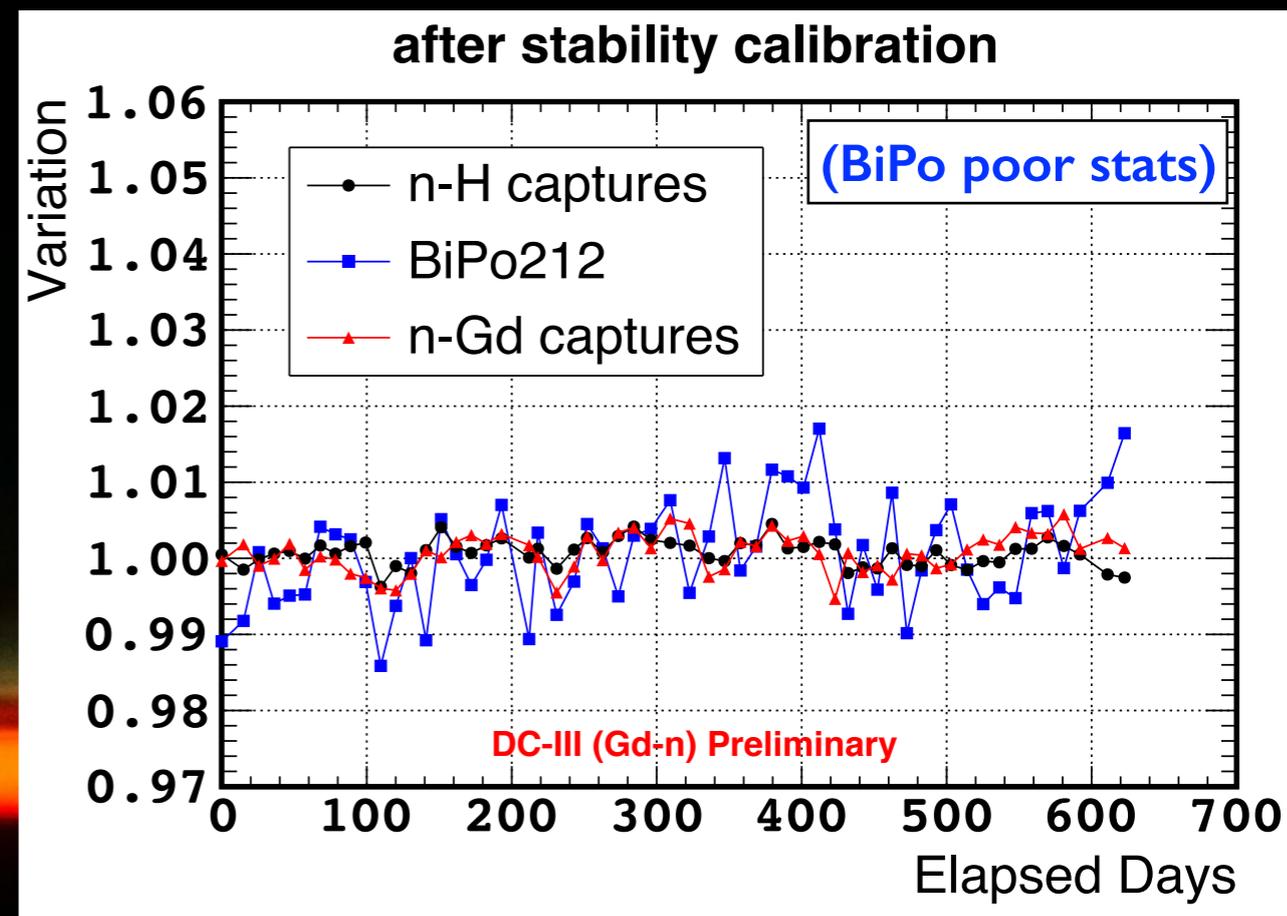
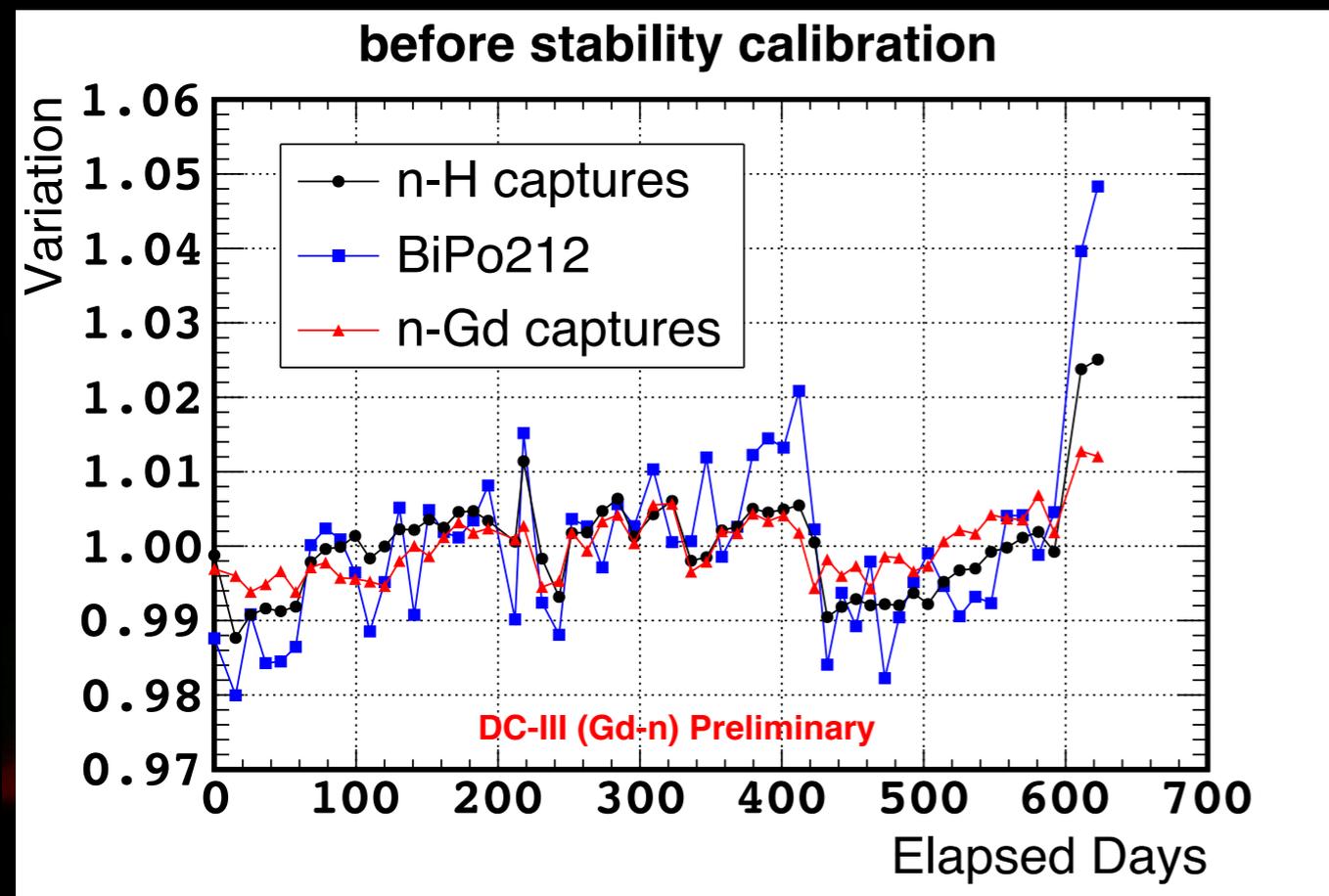
- single- $\gamma$  scintillation quenching measurement
  - many calibration sources @ center
- conversion:  $MeV(\text{single-}\gamma) \rightarrow MeV(e^+)$  [only MC]
- impact: **linearity (++)**

### • Overall performance...

- from  $Q(q, p, z, t)$  [RMS  $\sim 10\%$ ] to MeV [RMS  $\leq 1.0\%$ ]
- better detection systematics  $\rightarrow \theta^{13}$ , BGs,  $\Delta m^2$ .



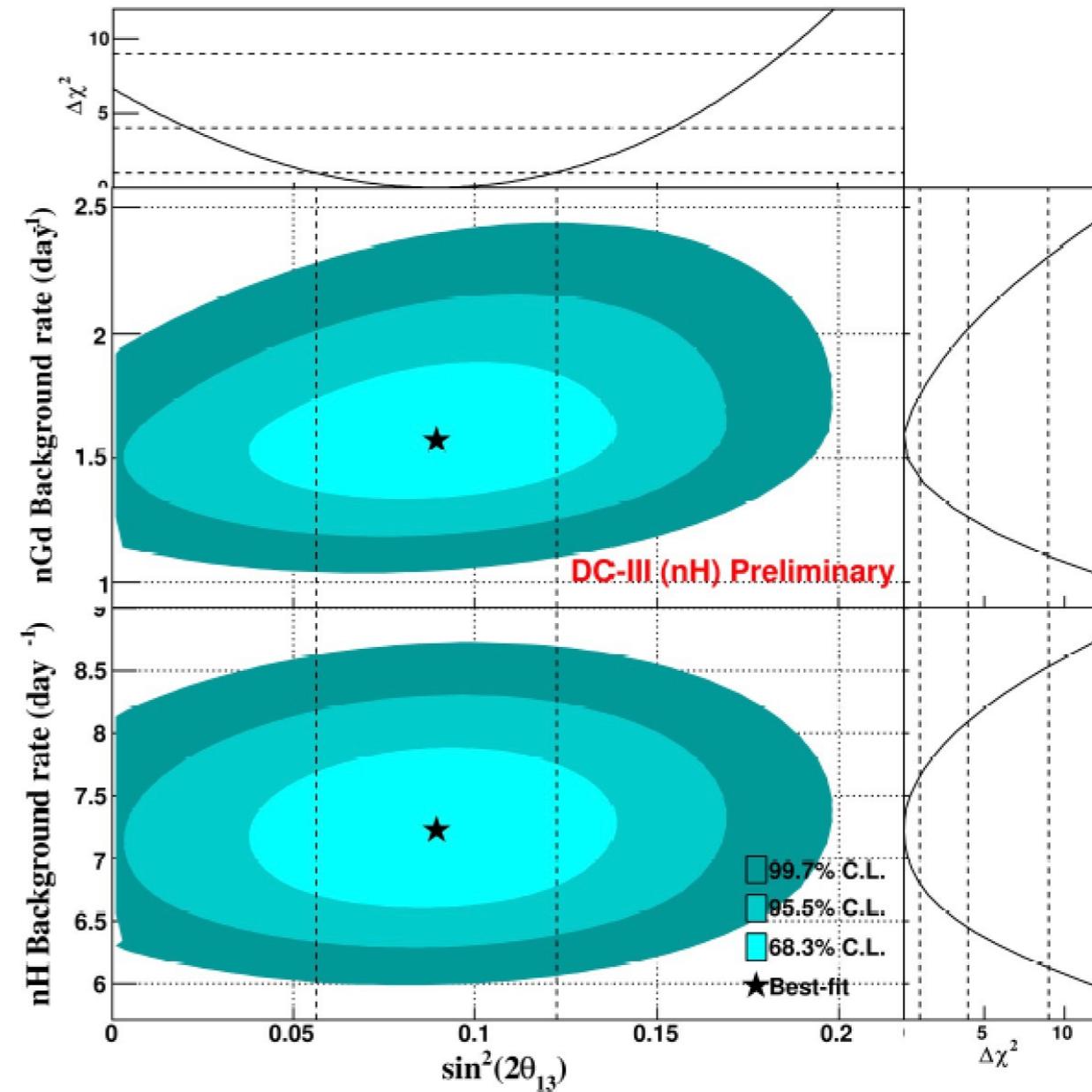
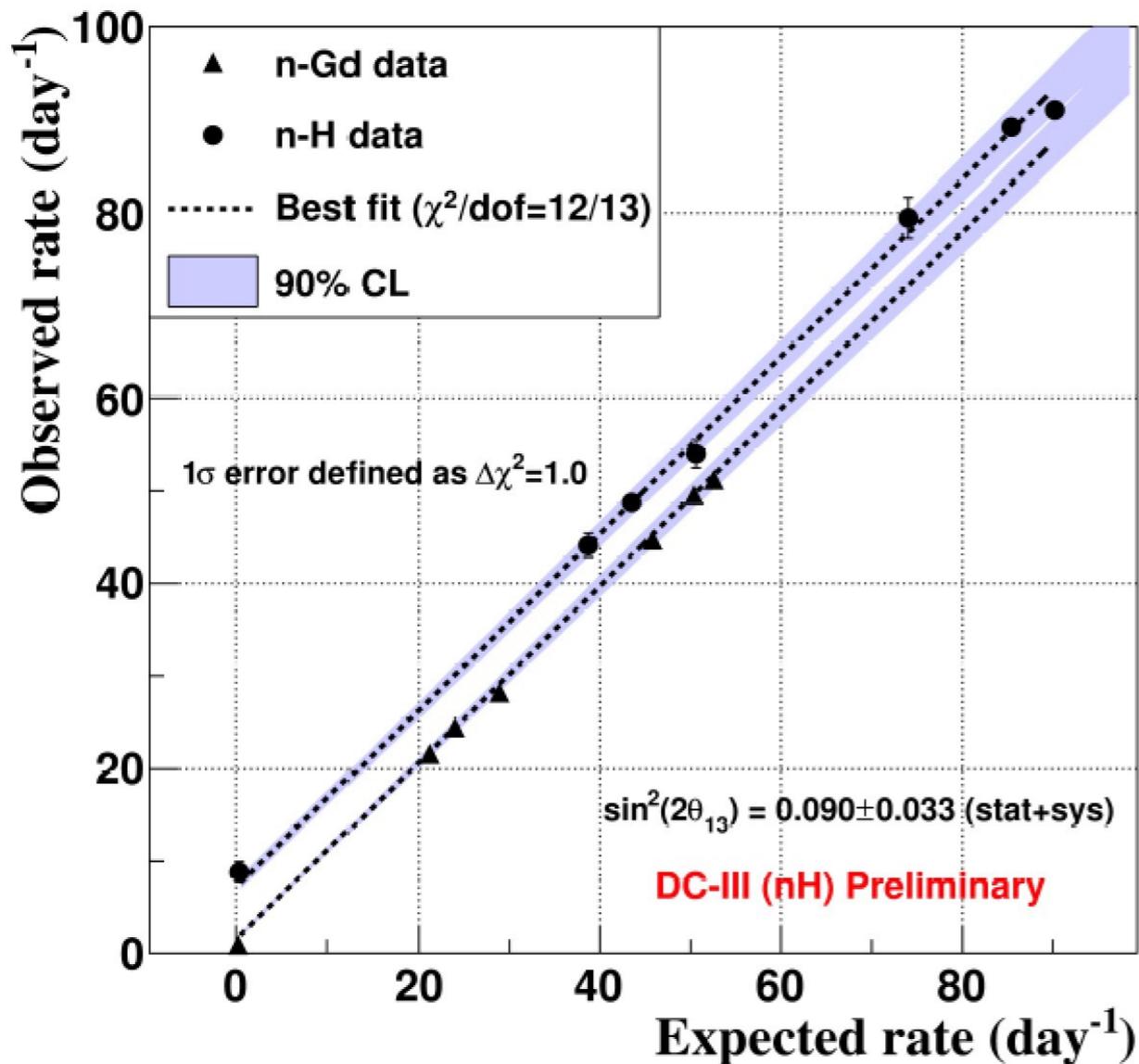
# response stability (with energy dependence)...



**raw response stability** → large non-statistical variations (electronics power cycle)  
 ( $\Delta$  up to 7%)

**corrected response** → pattern consistent with statistics dominated  
 ( $\Delta \leq 0.5\%$  + RMS > 2x better)

Combining this H-based result with latest Gd-based result (2014):



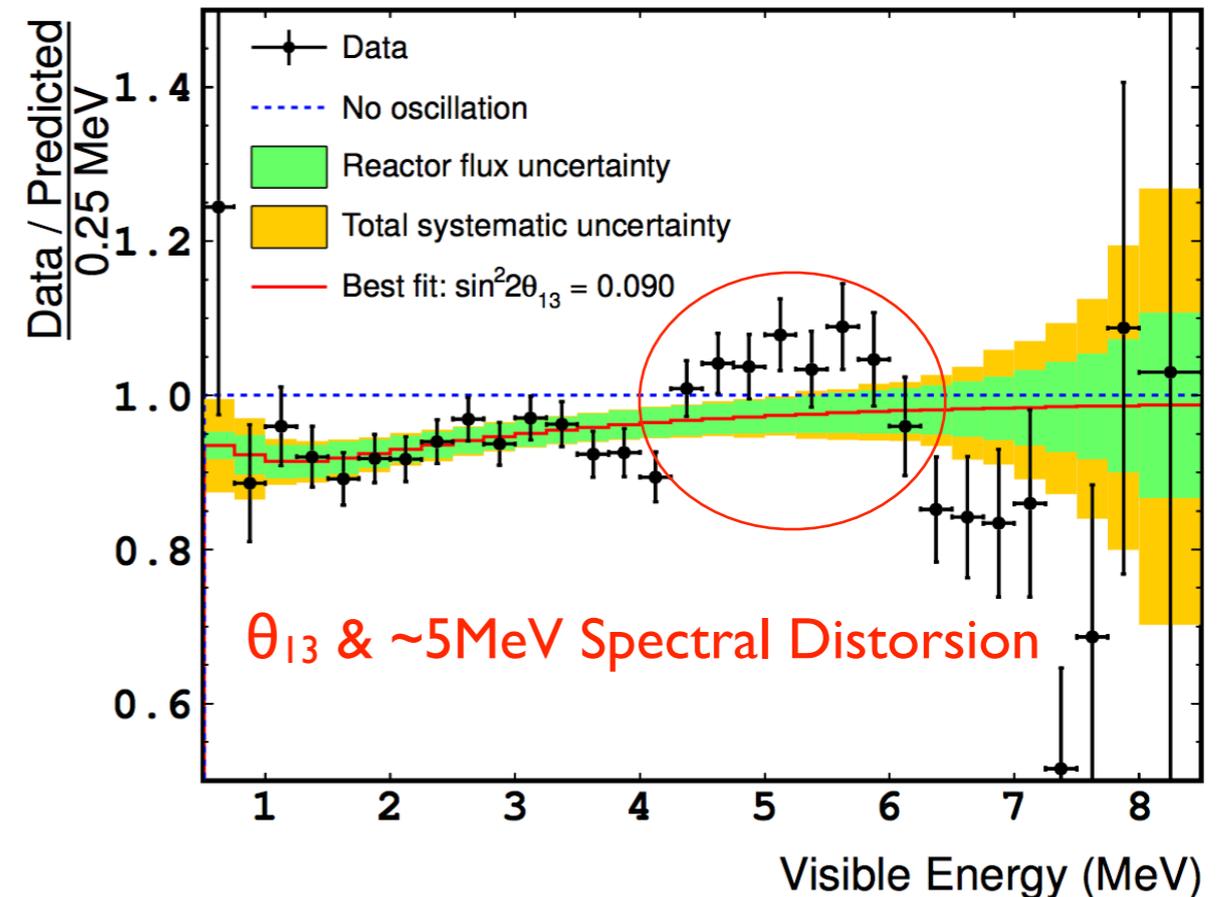
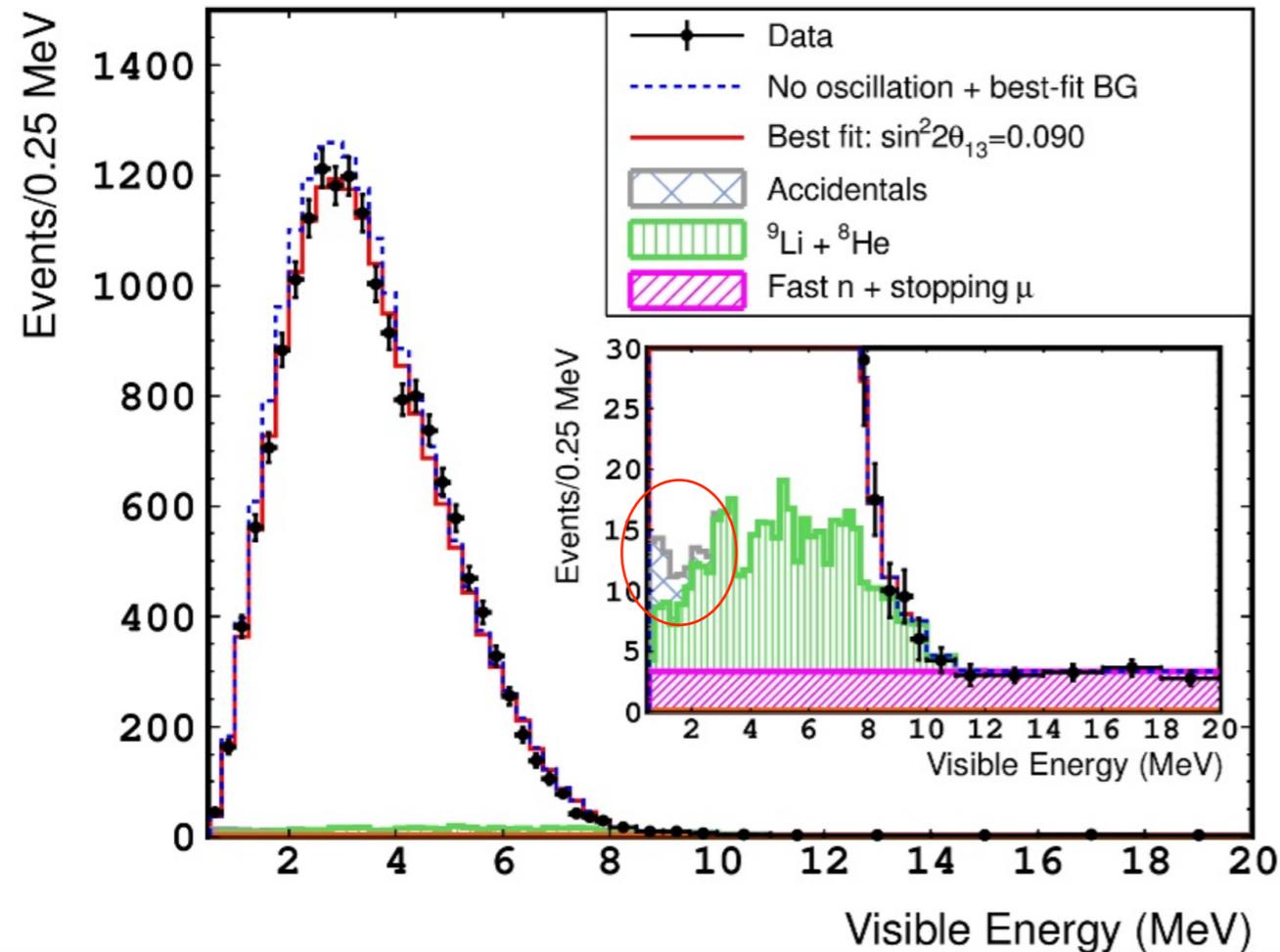
$$\sin^2 2\theta_{13} = 0.090 \pm 0.033$$

$$\text{H only: } \sin^2 2\theta_{13} = 0.098^{+0.038}_{-0.039}, \quad \text{Gd only: } \sin^2 2\theta_{13} = 0.090^{+0.034}_{-0.035}$$

Correlations between Gd and H have minimal impact. This result assumes no correlation.

# Oscillation fit: rate + spectrum shape

JHEP 1410 (2014) 086



- background and other uncertainties constrained by shape information

- $\sin^2 2\theta_{13} = 0.090^{+0.032}_{-0.029}$

- unexpected spectrum distortion observed at 4-6MeV (May 2014)

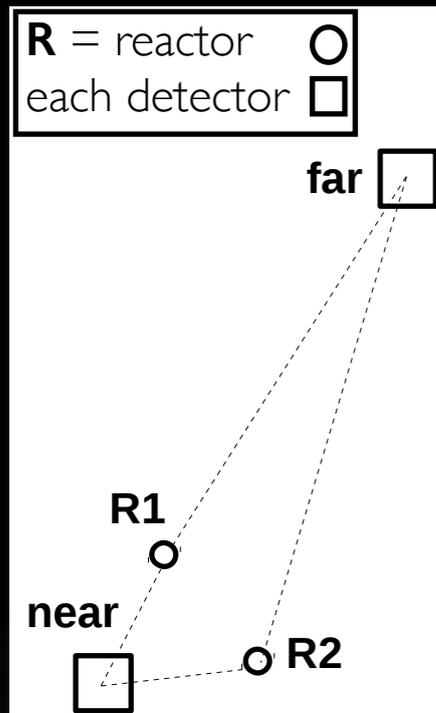
- ✓ Negligible impact to  $\theta_{13}$  measurement

- ✓ Magnitude of excess proportional to reactor power

- ✓ Same distortion confirmed by RENO (June-14), Daya Bay (July-14) and our n-H analysis

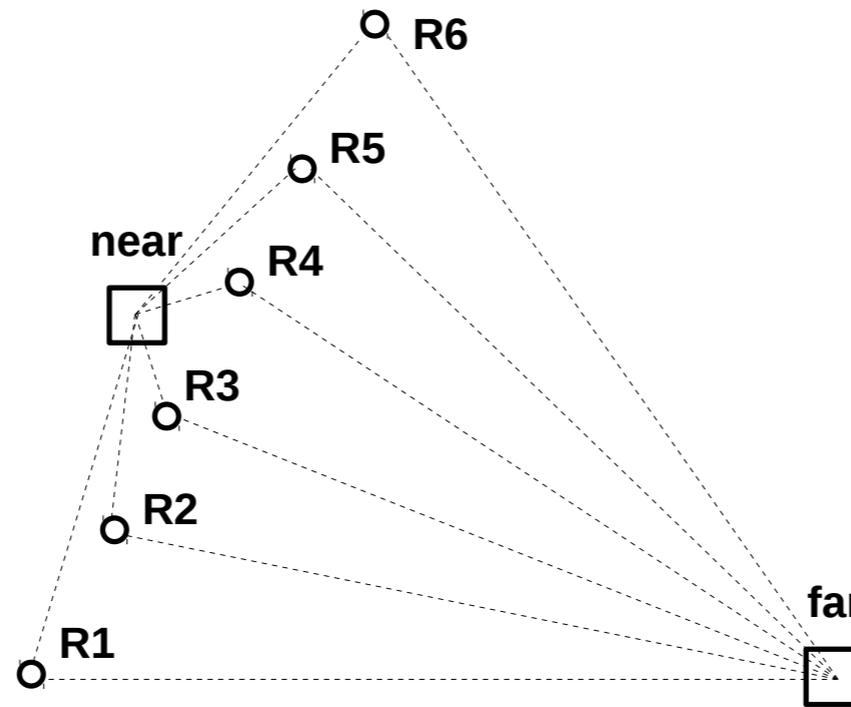
## reactor systematics → geometry...

## Double Chooz



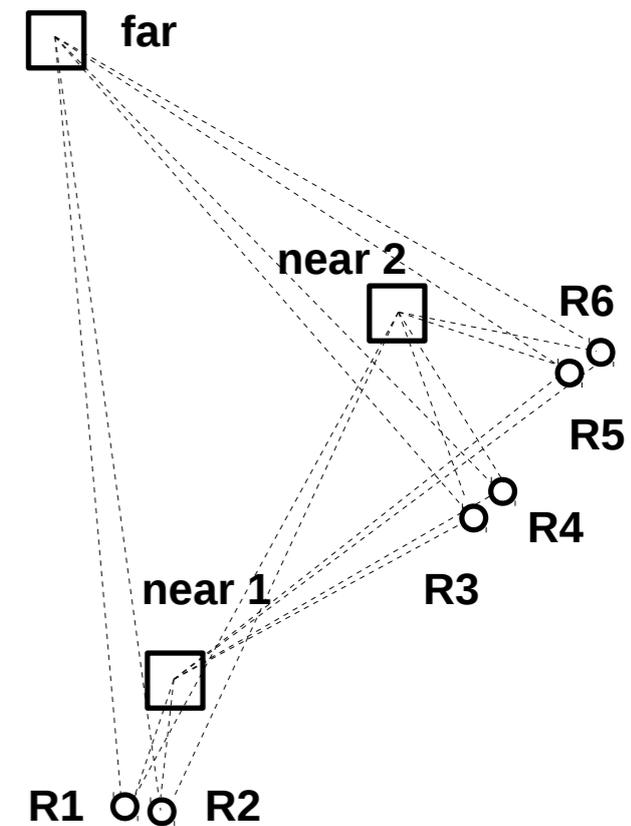
FD @ 300mwe  
(target: 8t)

## RENO



FD @ 450mwe  
(target: 16t) [2x DC]

## Daya Bay



FD @ 850mwe  
(target: 80t) [10x DC]

key design feature of experiment: **multi-detector(s) error cancellation**

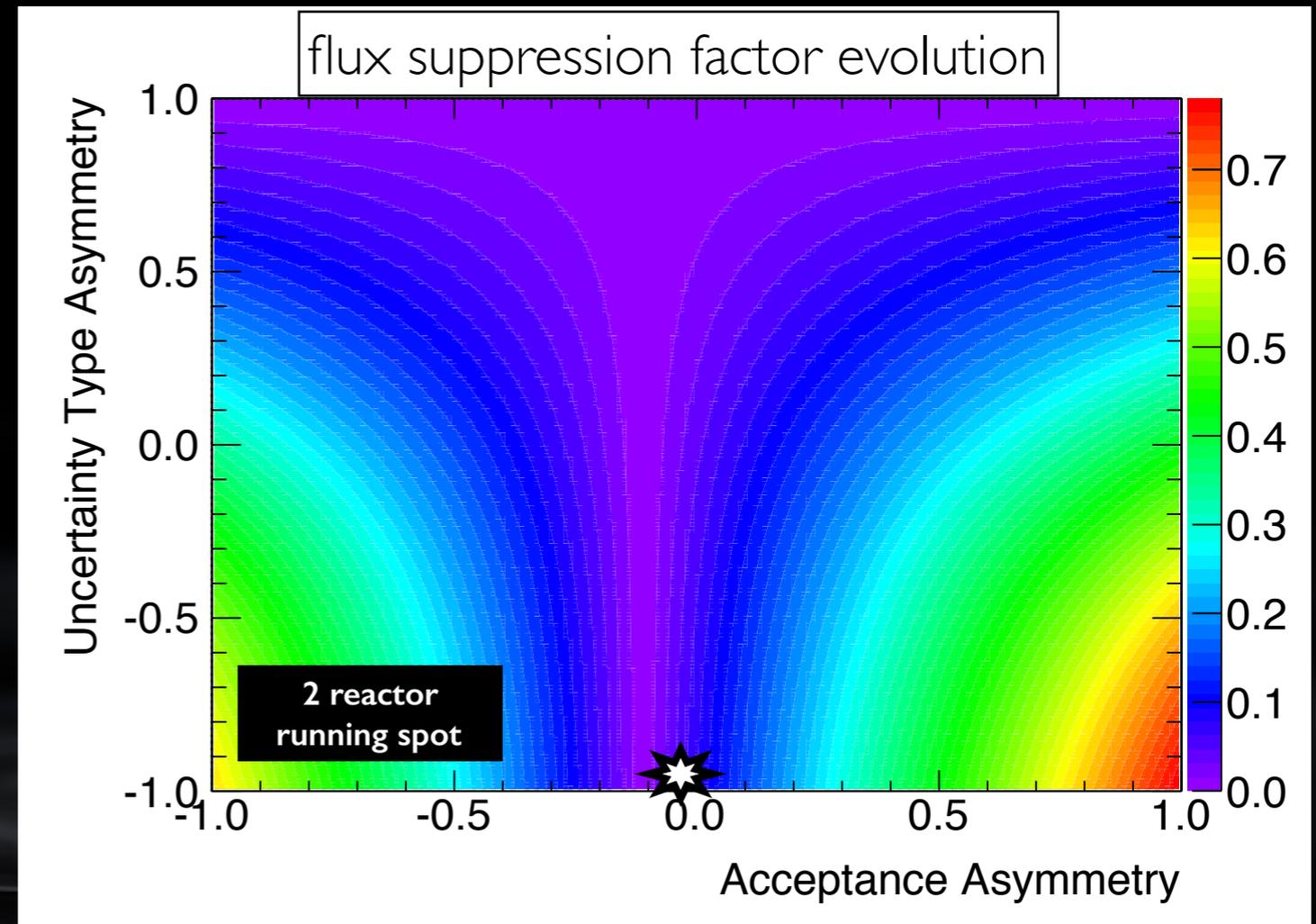
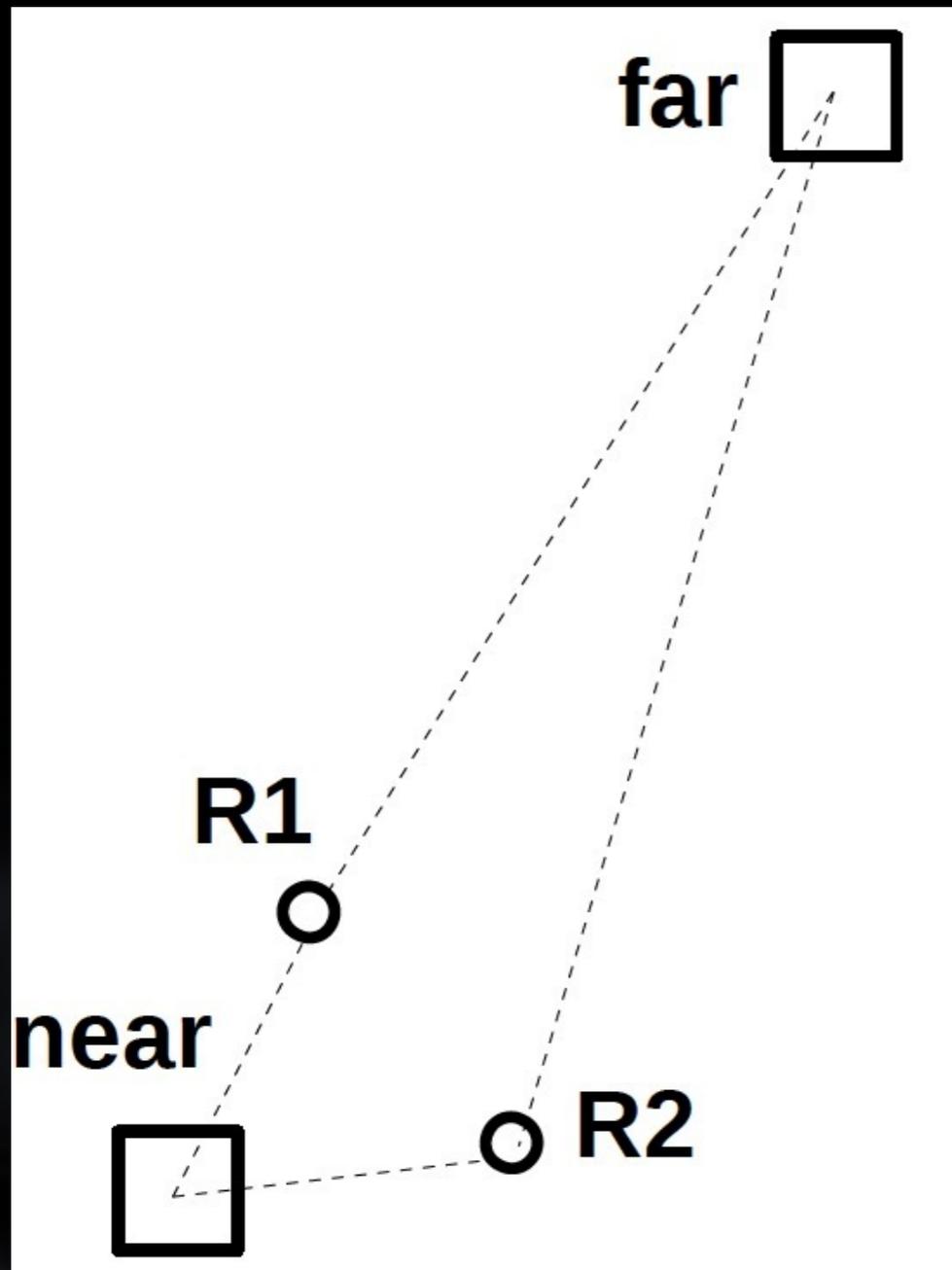
- (if identical) **detection-systematics** (target composition, efficiency, response, etc):  $\sim 2\% \rightarrow \sim 0.2\%$
- (if near/far) **flux systematics**:  $\sim 3.0\%$  (raw?)  $\rightarrow \sim 1.0\%$  (uncorrelated)  $\rightarrow \ll 0.1\%$  (geometry)

**[deeper underground + radio-purity  $\Rightarrow$  less background per detector]**

# DC major $\delta(\text{flux})$ cancellation with ND...

DC most iso-flux experimental setup

$\Rightarrow \sim 90\%$   $\delta(\text{flux})$  suppression



reactor error correlations  $\rightarrow \delta(\text{flux})$  suppression

$\delta(\text{flux})^{\text{FD}} = 1.7\% \rightarrow \delta(\text{flux})^{\text{FD+ND}} \leq 0.1\%$  (preliminary)

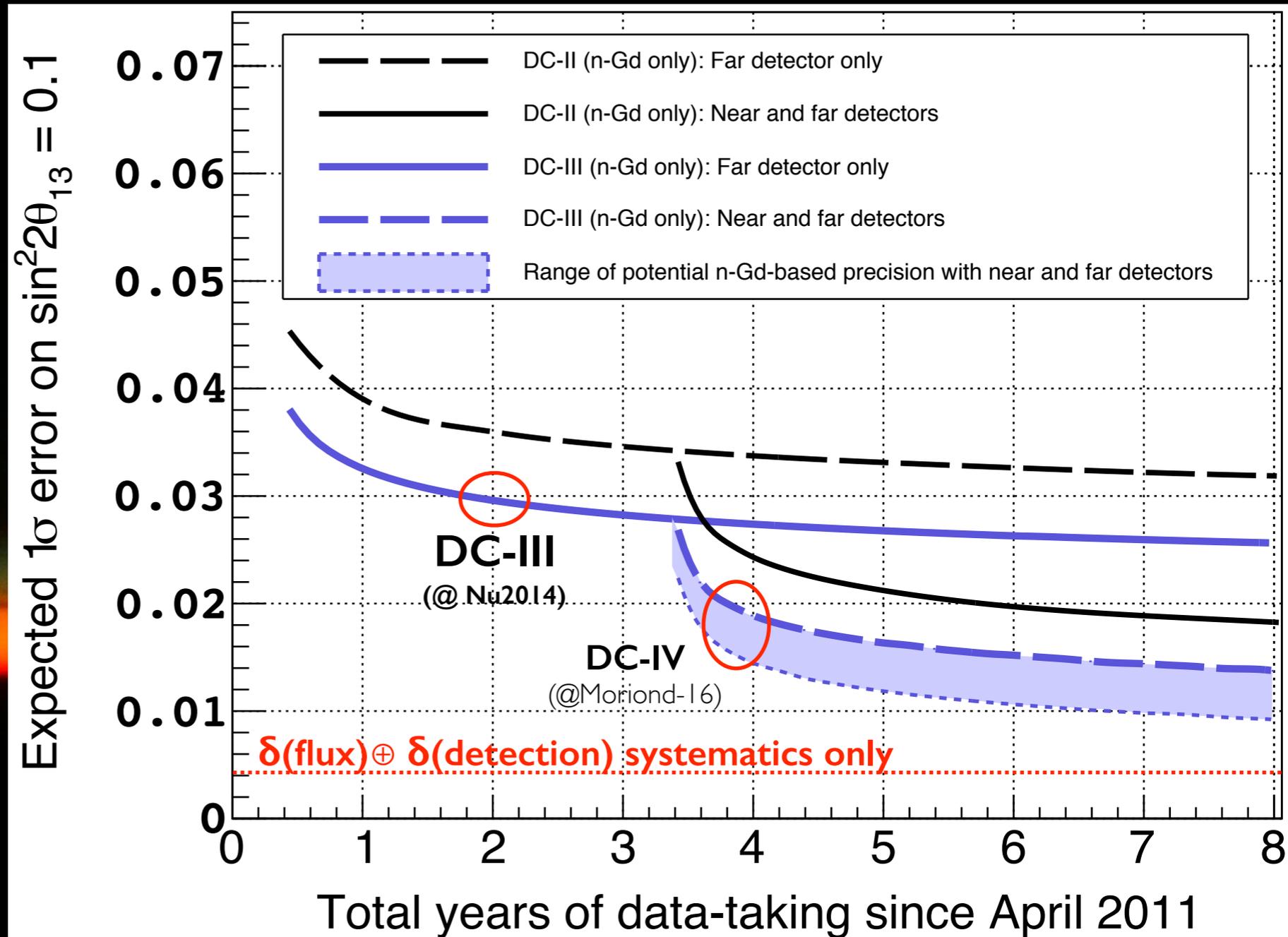
“Reactor Neutrino Flux Uncertainty Suppression on Multiple Detector Experiments”

Cucoanes, Novella, Cabrera et al. ([arXiv:1501.00356](https://arxiv.org/abs/1501.00356)) Anatael Cabrera (CNRS-IN2P3 & APC)

# $1\sigma$ error projection (via R+S analysis)...

## Gd-n analysis FD+ND prospect inputs

- $\delta(\text{flux}) \sim 0.1\%$  (**preliminary**)
    - iso-flux suppression dominated
  - $\delta(\text{detection}) \sim 0.2\%$  (**preliminary**)
    - à la Daya Bay / RENO
  - $\delta(\text{BG}) \sim \text{DC-III} + \text{R+S constraint}$ 
    - @DC-III  $\sim 0.3\%$  (2 years data)
- note:**
- $\delta(\text{stat})$  not just  $1/\sqrt{N^{\text{FD}}}$  (**dominant**)
    - several effects  $N^{\text{BG}}, N^{\text{ND}}, \text{etc}$



remarkable improvement of DC-III new analysis (wrt DC-II)

$1\sigma$  within  $[0.010, 0.014]$  with 3 years FD+ND: BG systematics dependent  $\rightarrow$  **statistics dominated**  
 (rate+spectrum projection uses latest BG model from DC-III)